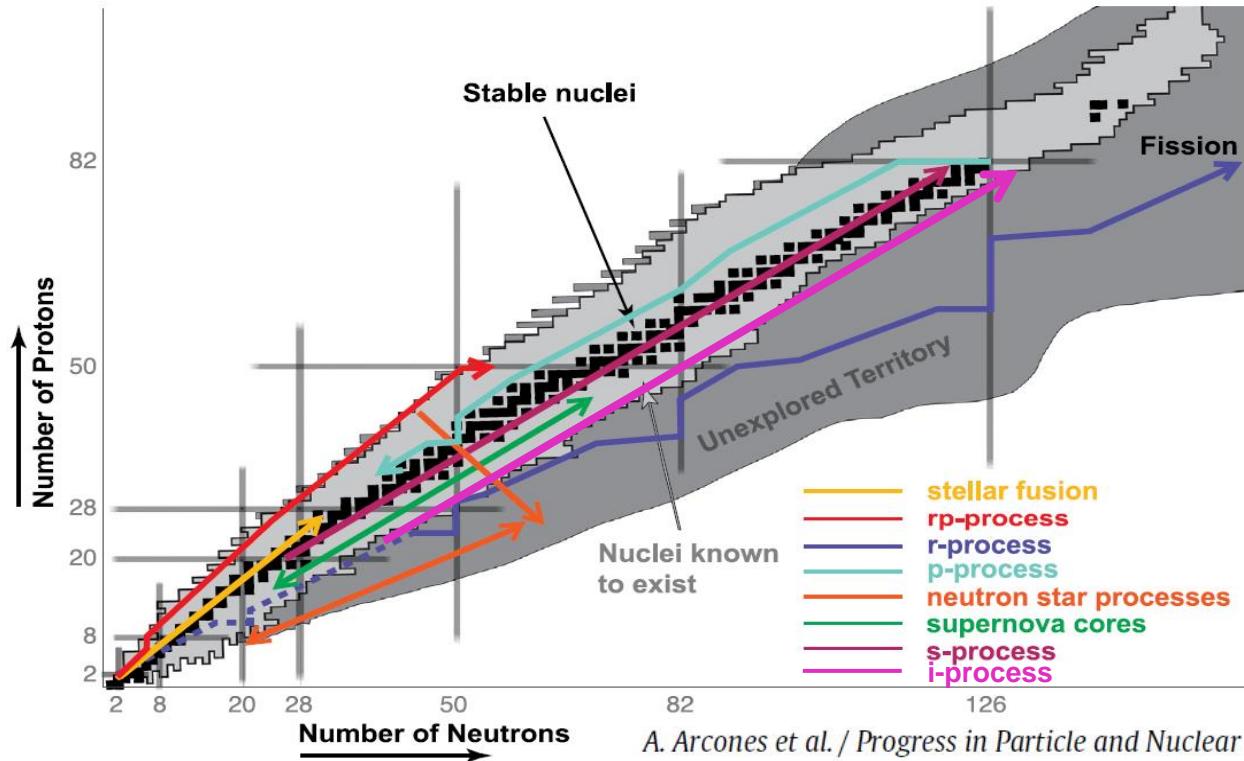


# Overview of theoretical efforts on microscopic nuclear level densities

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WANDA 2023

# Nuclear level densities (NLD)

Definition: number of levels per energy bin



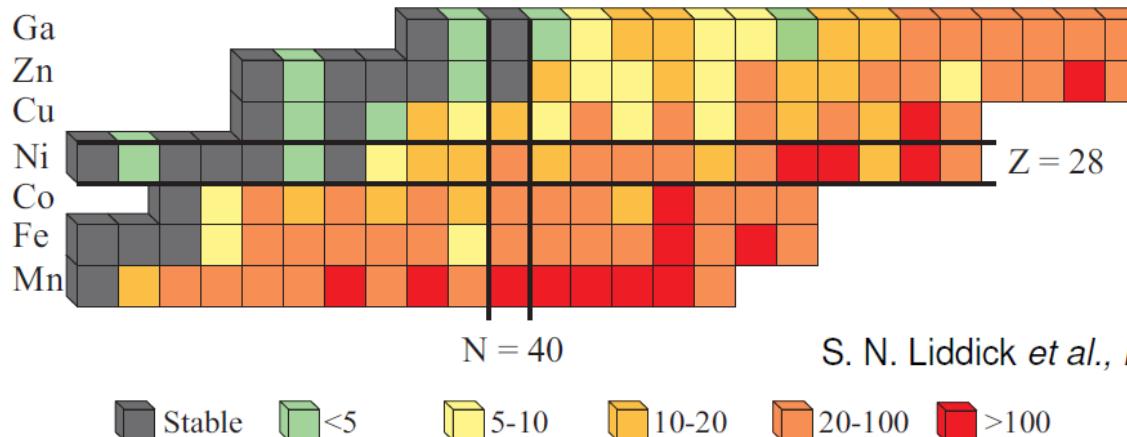
# Hauser-Feshbach nuclear input

- Main ingredients of statistical model calculations within the Hauser-Feshbach approach
  - Nuclear Level Densities
  - $\gamma$ -strength functions ( $\gamma$ SF)
  - Optical model potentials

# Impact of NLDs, $\gamma$ SF in neutron capture rates

## Variations of neutron capture rates at 1.5 GK

Nuclear Level Density	$\gamma$ ray Strength Function
Constant Temperature matched to the Fermi Gas model (CT+BSFG)[19]	Kopecky-Uhl generalized Lorentzian (KU) [17]
Back-shifted Fermi Gas model (BSFG)[19],[20]	Hartree-Fock BCS + QRPA (HF-BCS+QRPA) [21]
Generalized Super fluid model (GSM)[22], [23]	Hartree-Fock-Bogolyubov + QRPA (HFB+QRPA) [24]
Hartree Fock using Skyrme force (HFS) [25]	Modified Lorentzian (Gor-ML)[26]
Hartree-Fock-Bogoliubov (Skyrme force) + combinatorial method (HFBS-C) [27]	



S. N. Liddick *et al.*, PRL 116 242502 (2016)

# Experimental NLDs

- Low energy discrete experimental levels
- Level density from neutron resonance spacings at the neutron separation energy (available only for specific spins)
- Oslo method and  $\beta$ -Oslo technique (require normalization to the low energy discrete levels and to the level density at neutron separation energies or theoretical estimates)
- Particle evaporation technique

# NLDs from theory

Phenomenological models

- Fermi gas model

$$\rho(E_x, J, \pi) = \frac{1}{2} \frac{2J+1}{2\sigma^2} \exp\left[-\frac{(J+1/2)^2}{\sigma^2}\right] \frac{1}{\sqrt{2\pi}\sigma} \frac{\sqrt{\pi}}{12} \frac{\exp[2\sqrt{aU}]}{a^{1/4} U^{5/4}}$$

- Constant temperature model

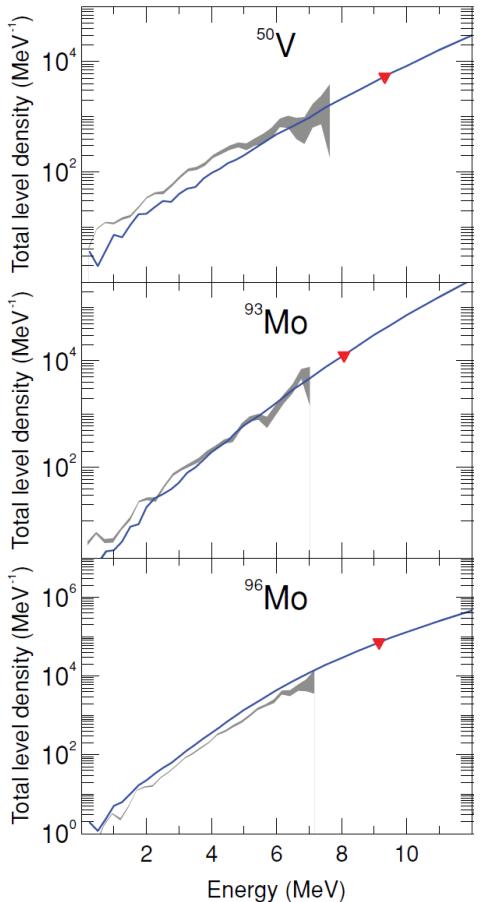
$$\rho(E_x, J, \pi) = \underbrace{\frac{1}{2} \frac{2J+1}{2\sigma^2} \exp\left[-\frac{(J+1/2)^2}{\sigma^2}\right]}_{\text{spin distribution}} \frac{1}{T} \exp\left[\frac{E_x - E_0}{T}\right]$$

- ✓ Extensively studied, available in reaction codes (e.g. TALYS)
- ✗ Parameters  $(a, U, E_0, T)$  must be determined from the available experimental data or from empirical expressions, knowledge of the spin distribution and spin cut-off parameter  $\sigma$  is required

# NLDs from theory

## Microscopic models

- Based on Hartree-Fock calculations
- ✓ spin and parity dependent NLDs
- ✓ Level densities available for thousands of nuclides, high excitation energies and spins
- ✓ available in reaction codes (TALYS)
- ✗ many-body correlations missing



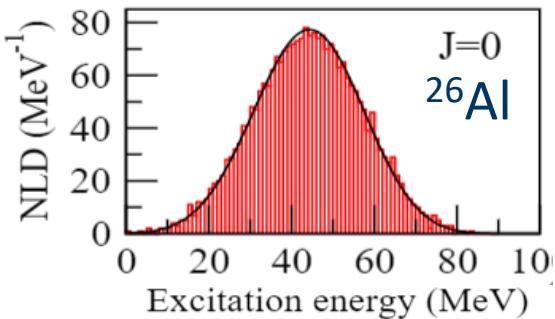
S. Goriely et al., ADNDT 77 311 (2001) S. Hilaire et al., PRC 86 064317 (2012)

S. Goriely et al., PRC 78 064307 (2008)

# NLDs from theory

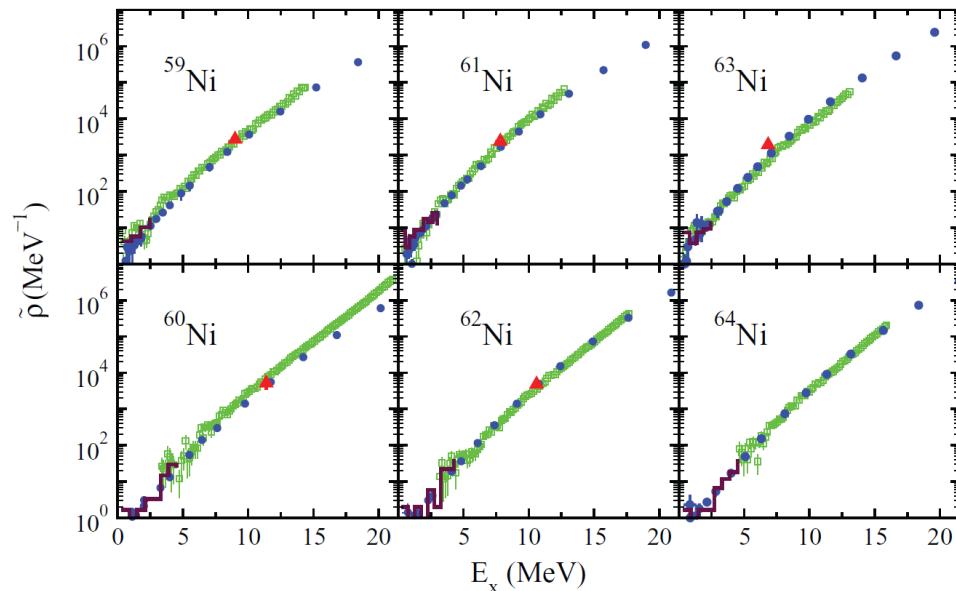
## SHELL MODEL APPROACHES (spin and parity dependent NLDs)

- Configuration interaction shell model calculations using conventional diagonalization;  $\times$  *sd*-nuclei



Y. Alhassid et al., PRL 99 162504 (2007)  
M. Bonett-Matiz et al. PRC 88 011302 (R) (2013)

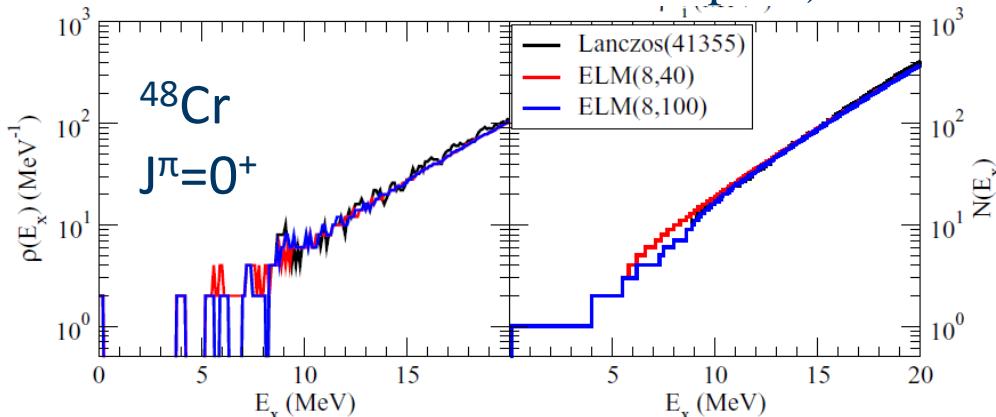
- Shell Model Monte Carlo;  
mid-mass and heavy nuclei



# NLDs from theory

## SHELL MODEL APPROACHES (spin and parity dependent NLDs)

- Lanczos method, computes moments of the Hamiltonian; mid-mass nuclei (requires about 100 iterations in the full model space)



- Moments method, computes the first two moments of the Hamiltonian; does not require diagonalization in the full model space

W. E. Ormand et al., PRC 102 014315 (2020)

R. Sen'kov et al., PRC 93 064304 (2016)

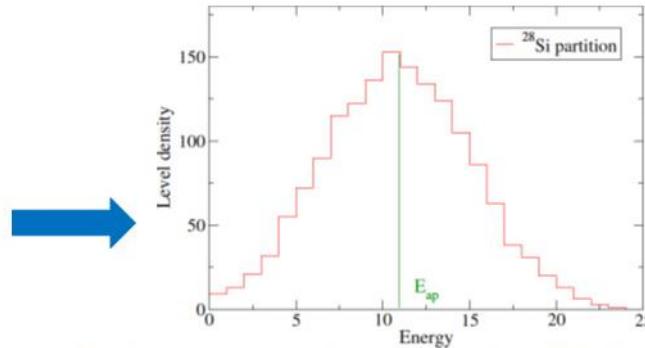
# Calculation of level density – Moments method

Partitions, p	$d_{5/2}$	$s_{1/2}$	$d_{3/2}$
1	6	0	0
2	5	1	0
3	5	0	1
4	4	2	0
...	...	...	...
15	0	2	4

$E_{g.s.}$  : Shell model

$\eta$  : cut-off (2.8)

$$\rho(E; a) = \sum_p D_{ap} G_{ap}(E)$$



Each partition p, gives a gaussian distribution

$$E_{ap} = \frac{1}{D_{ap}} \text{Tr}^{ap} H$$

$$\sigma_{ap}^2 = \frac{1}{D_{ap}} \text{Tr}^{ap} H^2 - E_{ap}^2$$

$$G_{ap} = G(E - E_{ap} + E_{g.s.}; \sigma_{ap})$$

$$G(x; \sigma) = C \begin{cases} e^{-x^2/2\sigma^2}, & |x| \leq \eta\sigma, \\ 0, & |x| > \eta\sigma \end{cases}$$



R. Sen'kov et al., CPC 184, 215 (2013)

# Model spaces

Tested with

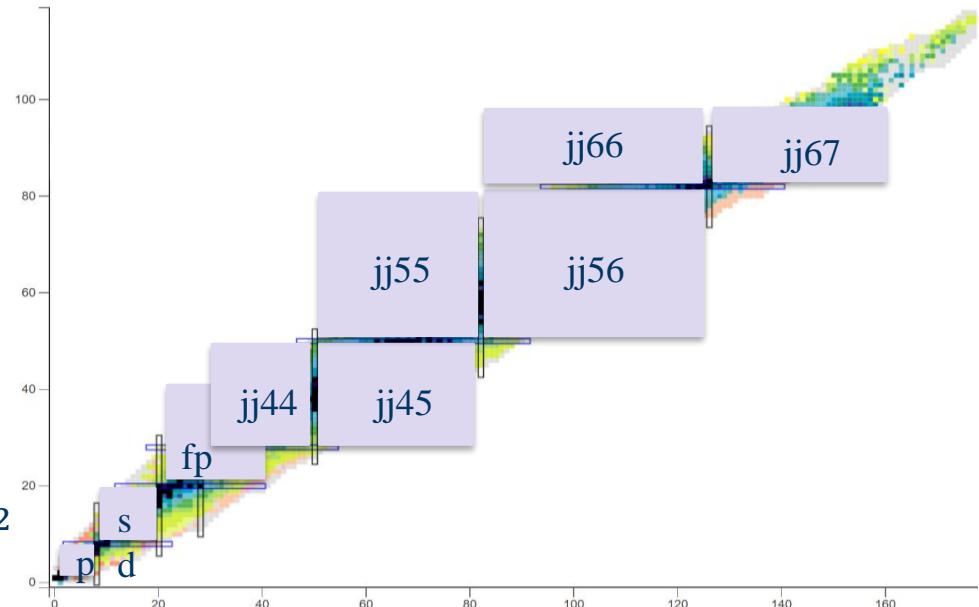
- $sd - 0d_{5/2}, 0d_{3/2}, 1s_{1/2}$
- $pf - 0f_{7/2}, 0f_{5/2}, 1p_{3/2}, 1p_{1/2}$
- $jj44 - 0f_{5/2}, 1p_{3/2}, 1p_{1/2}, 0g_{9/2}$
- $pf + 0g_{9/2}$

Extensions

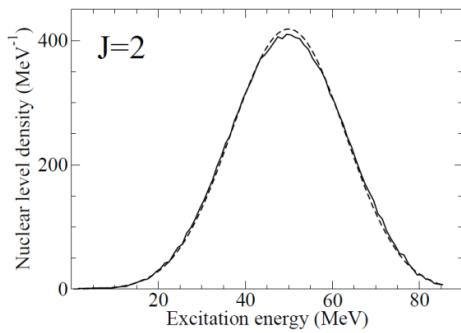
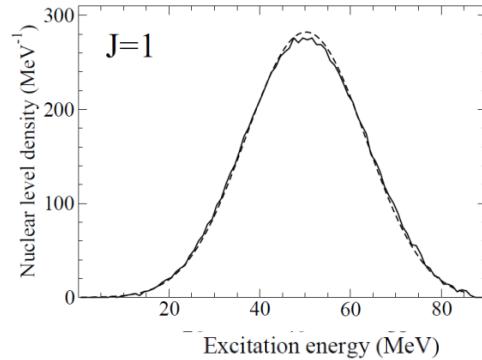
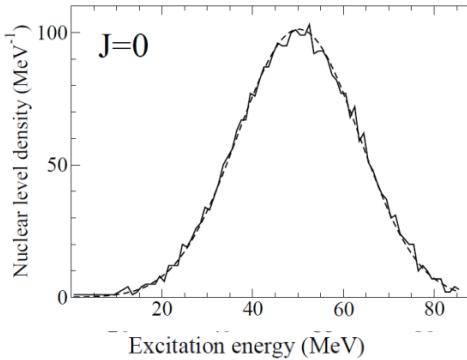
- $jj55 - 0g_{7/2}, 1d_{5/2}, 1d_{3/2}, 2s_{1/2}, 0h_{11/2}$

...

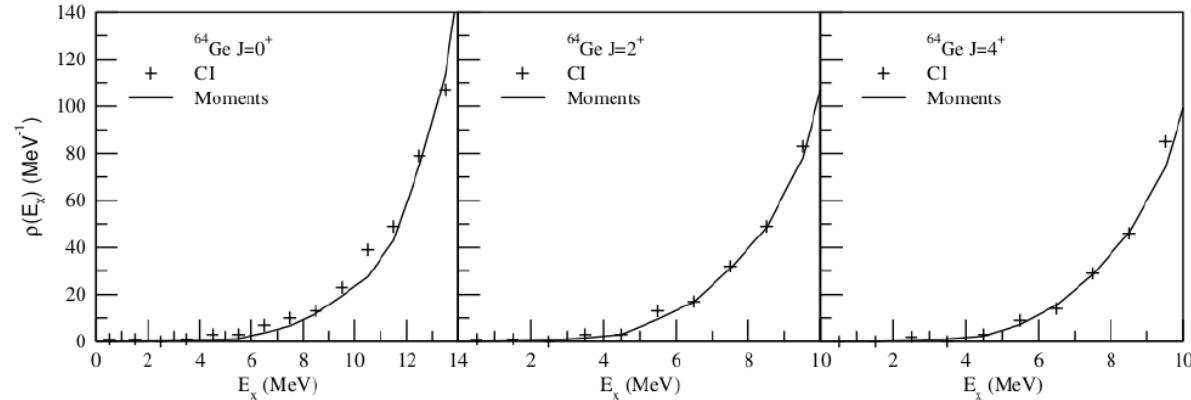
- Any model space for which an effective shell model Hamiltonian is available



# MM vs Exact SM calculations NLDs



sd –  $^{28}\text{Si}$

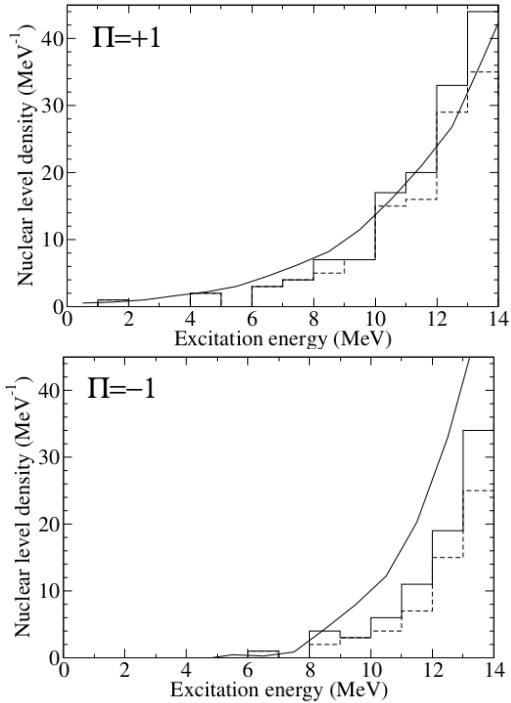


jj44 –  $^{64}\text{Ge}$

R. Sen'kov et al., PRC 93 064304 (2016),

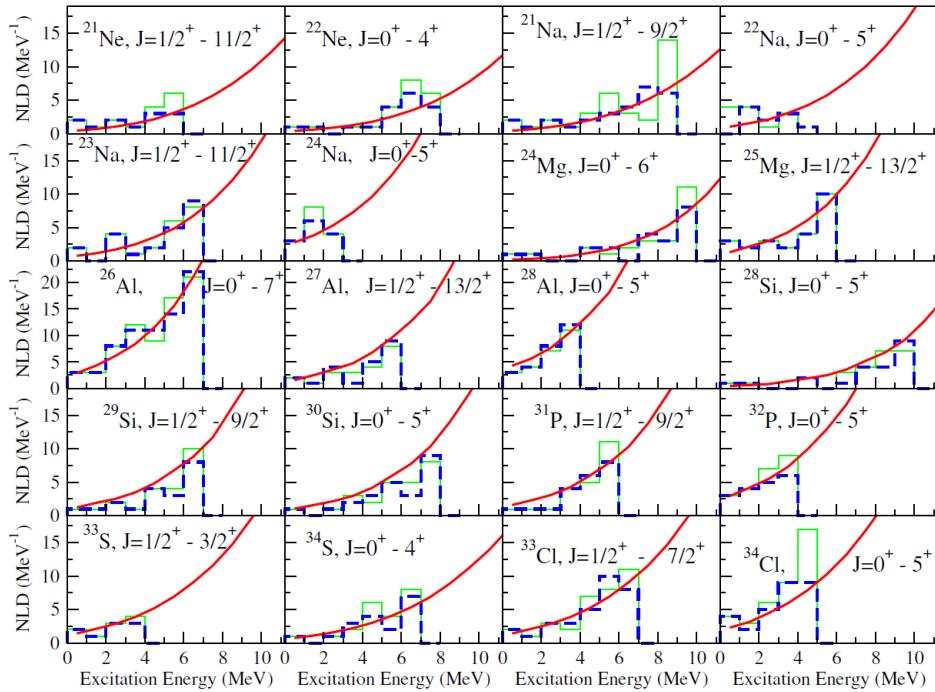
M. Scott et al., PoS (2008)

# MM vs Experimental NLDs



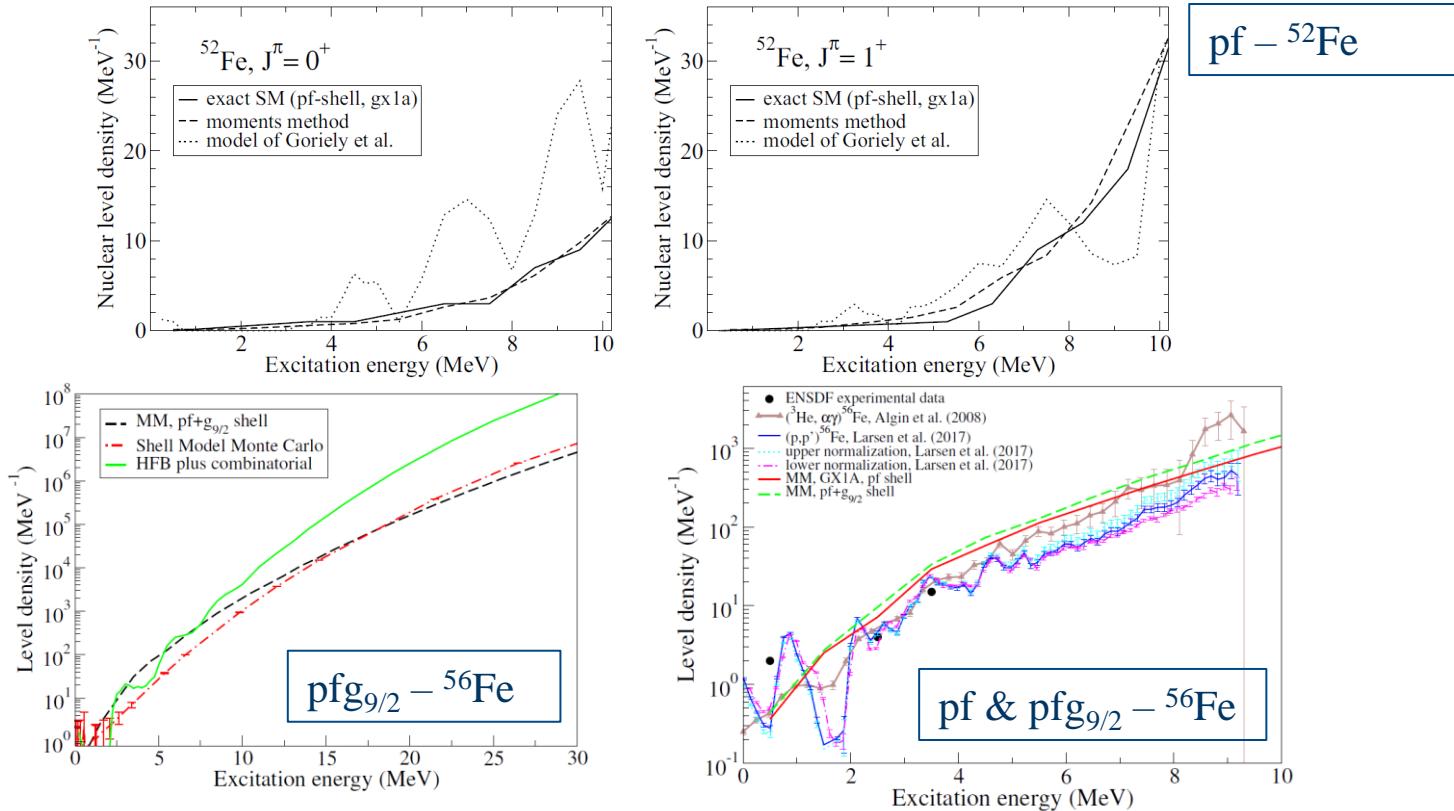
$s + p + sd + pf - {}^{28}\text{Si}$

R. Sen'kov et al., PRC 93 064304 (2016)



S. Karampagia et al., ADNT 1, 120 (2017)

# MM vs other models & Oslo method



# Challenges

- Shell model level densities have a finite excitation range ( $\sim 12$  MeV); need for an algorithm to continue to higher excitation energies
- Ground state is required; directly from a shell model calculation, other extrapolation techniques
- Complex process to insert moments method level density in tables of nuclear reaction codes, such as TALYS.
- Availability of reliable shell model interactions (away from stability what?)

# Thank you

## Collaborators

- MSU: Vladimir Zelevinsky, Alex Brown
- CMU: George Perdikakis, Mihai Horoi