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CIPANP-2018 13th conference on the intersections of particle and nuclear physics

Palm Springs May 30, 2018

Dark Matter (DM) is there!

What do we know about it? Not much

1. It gravitates





Coma cluster (of galaxies)



Andromeda Galaxy

- 2. It is dark (i.e. it does not interact with photons)
- **3.** It is stable on cosmological scales

Fun fact: There is lots of DM in the Universe, but

for DM particles weighing several hundred times the mass of the proton, there should be about one DM particle per coffee-cup-sized volume of space.

Stars, Planets

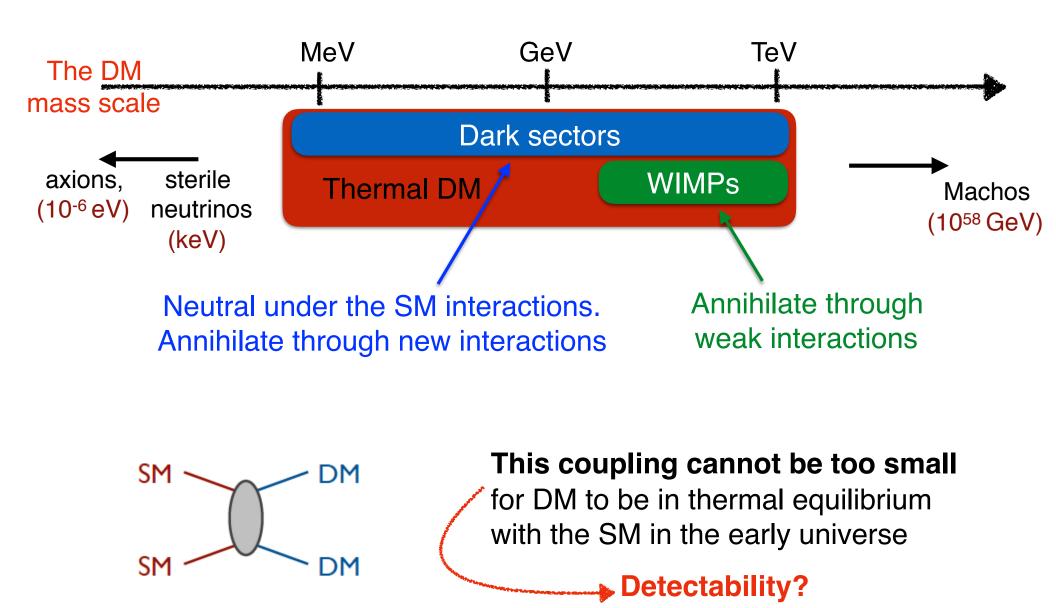
🗕 Dark Matte

23%

Dark Energy

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Thermal dark matter



Dark sectors ... beyond Dark Matter

Further motivations?

- Several anomalies in data can be addressed by dark sectors (eg. $(g-2)_{\mu}$, B-physics anomalies, Dark Matter anomalies (galactic center excess), ...);

- Neutrino mass model building

What theories?

DM theories, Supersymmetric theories (NMSSM), neutral naturalness theories, theories for baryogengesis, ...

Dark sectors ... beyond Dark Matter

Further motivations?

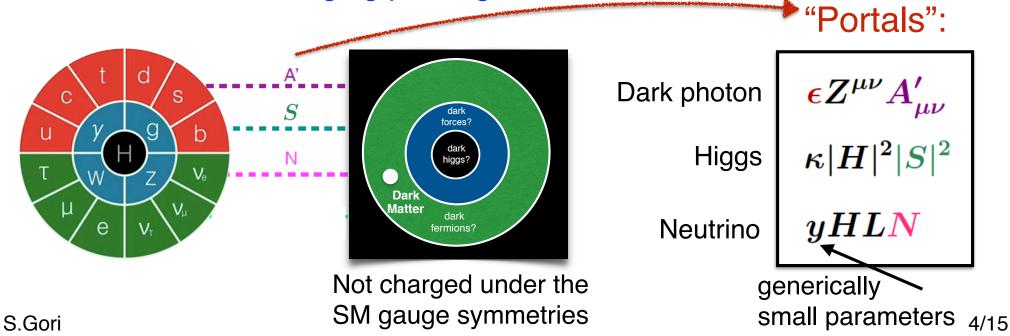
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How to test this emerging paradigm?



The dark photon

Nature seems well described by a SU(3) x SU(2)_L x U(1)_{em} gauge theory. We need to check this assumption!

Additional gauge symmetries in nature? U(1)'?

Holdom, '86

$$\mathcal{L}\subset \epsilon Z^{\mu
u}A'_{\mu
u}+m^2_{A'}A'_{\mu}A'^{\mu}$$
 + couplings with the dark sector

Mixing with the SM hyper-charge gauge boson

arising from

* dark Higgs mechanism or



* Stueckelberg mechanism

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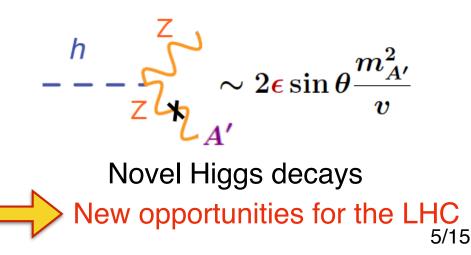
* Stueckelberg mechanism

➡ Massive photon

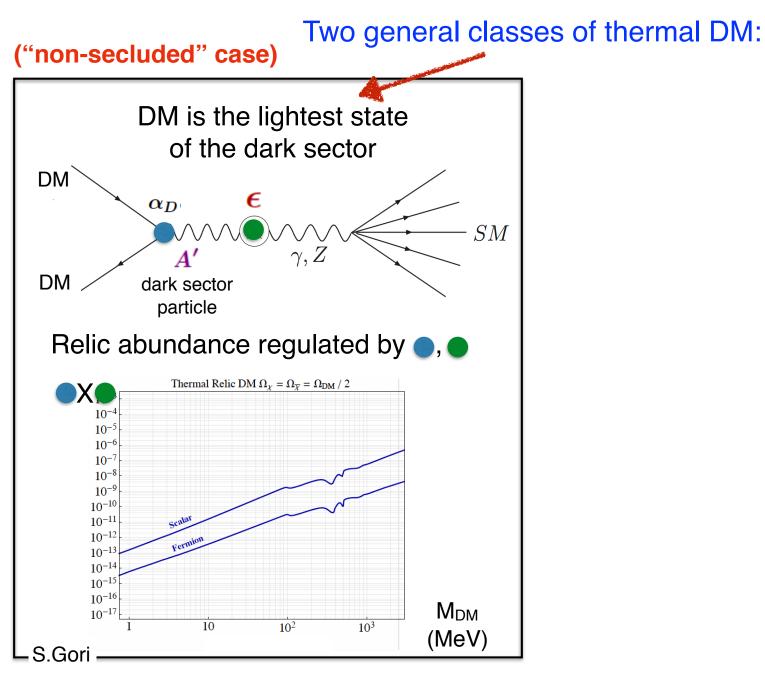
The SM Z boson is affected

Mass of the Z boson: $m_Z^2 \sim m_{Z0}^2 (1 + \epsilon^2 \sin^2 heta)$

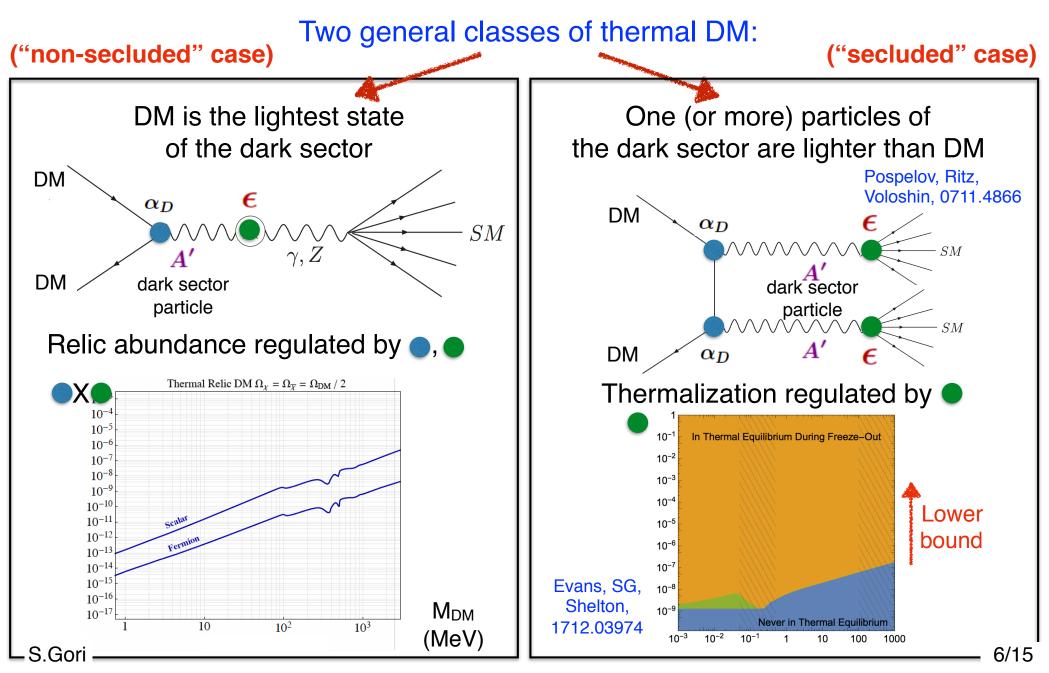
Couplings of the Z boson with fermions: $(Zf\bar{f}) (1 + \epsilon^2 \sin^2 \theta F(T_3, Q))$ The SM Higgs boson is affected



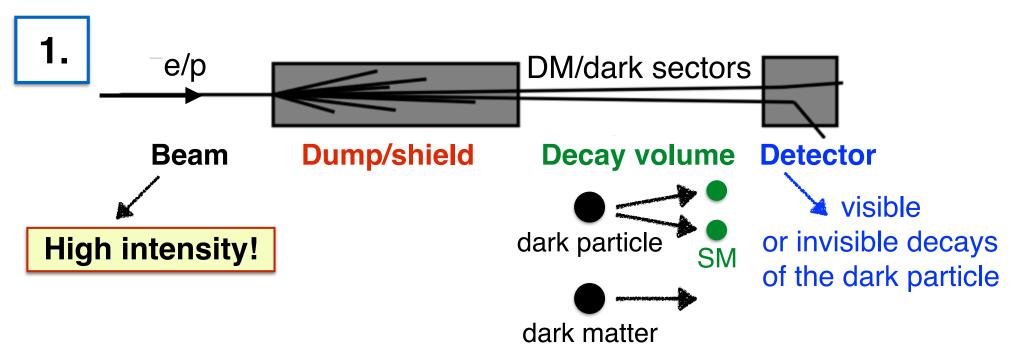
"Thermal goals"

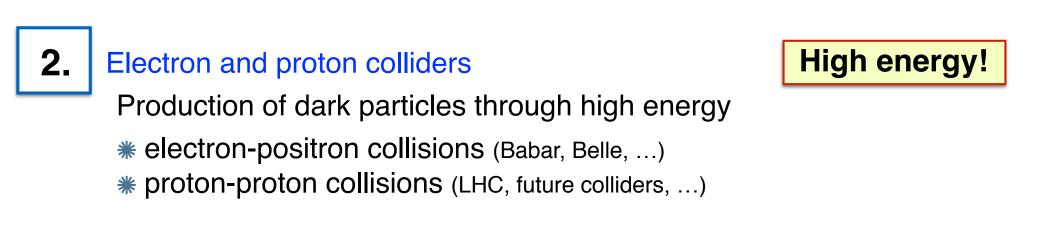


"Thermal goals"



Fixed target & colliders experiments





The production of dark photons $\epsilon Z^{\mu u}A'_{\mu u}$ Electron fixed target experiments Proton fixed target experiments e bremsbremsstrahlung strahlung meson **Electron-positron colliders** decay 1 **Proton-proton** Drell-Yan colliders:

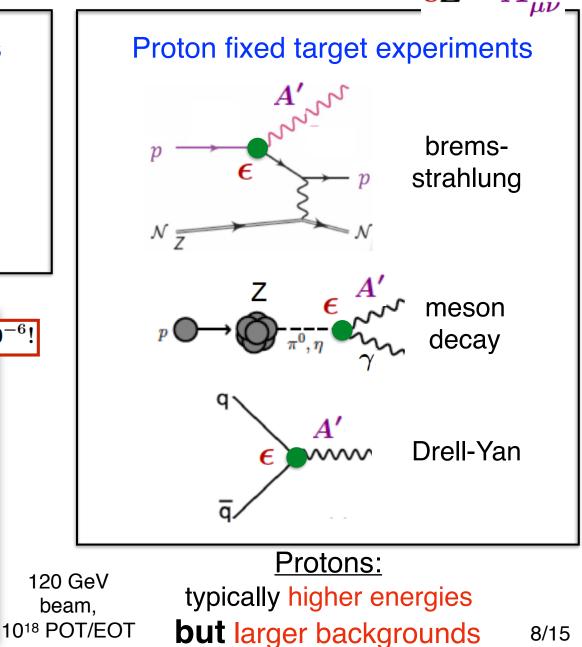
High intensity High energy

The production of dark photons $\epsilon Z^{\mu u}A$

Electron fixed target experiments e bremsstrahlung Berlin, SG, Schuster, Toro, 1804.00661 $10^7 \xrightarrow{\pi^0 \rightarrow \gamma A'}$ $\epsilon = 10^{-6}!$ 106 10⁵ $\omega \rightarrow \pi^0 A'$ $N_{A'}$ 10⁴ 10^{3} 10^{2} 10¹ 120 GeV 10⁻¹ 10^{-2} beam,

 $m_{A'}$ [GeV]

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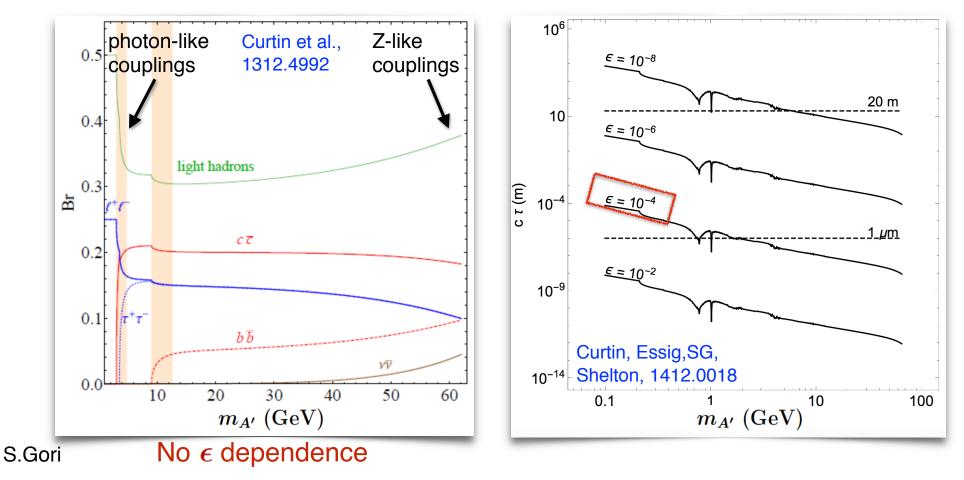


Minimal dark photon signatures (secluded*)

*DM is heavier

Lifetime and decay mode dictates search strategy

Only relevant free parameters of the minimal model: ϵ , $m_{A'}$ The dark photon can only decay to SM particles (visible decays)



Present & future reach

(secluded*) Berlin, SG, Schuster, Toro, 1804.00661 $A' \rightarrow \ell^+ \ell^-$ LEP bound is here 10^{-2} NA48/2 Babar **(g-2)**μ 10^{-3} HPS 10^{-4} LHCb EE774/ 10⁻⁵ E141 SeaQuest LHC ϵ FASER 10^{-6} Charm NA62 10^{-7} LSND/E137 SHiP 10^{-8} Supernova 10^{-9} 10^{-1} 10^{-2} $m_{A'}$ [GeV]

The entire parameter space is motivated by thermal DM! (our goal)

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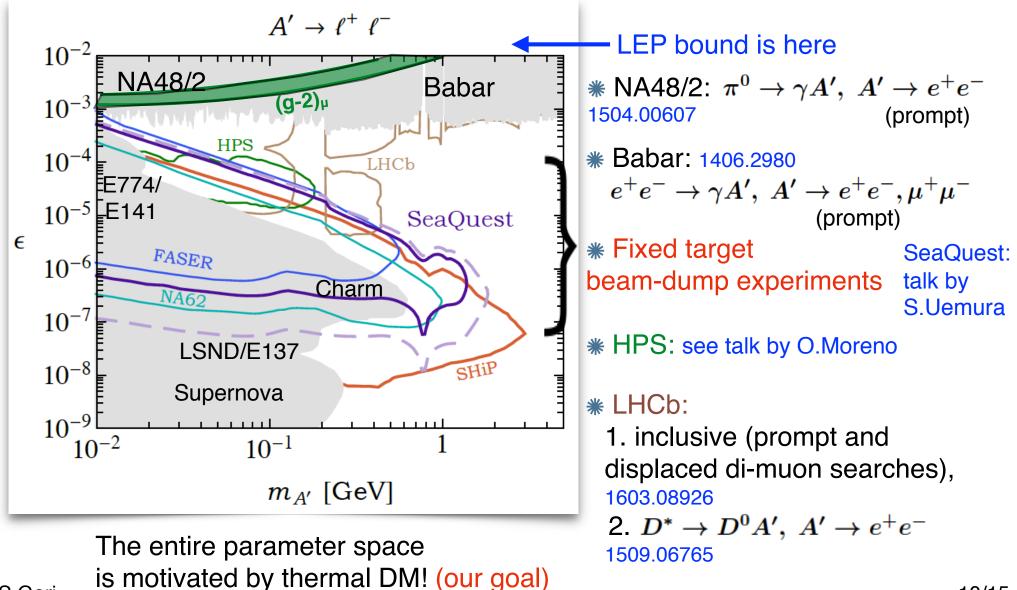
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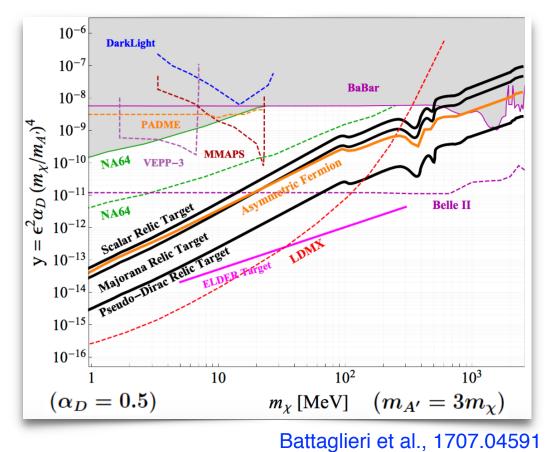


Minimal dark photon signatures (non-secluded*)

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Lifetime and decay mode dictates search strategy

Only relevant free parameters of the minimal model: ϵ , α_D , $m_{A'}$ The dark photon decays mainly to dark matter (invisible decays), $\alpha_D \gg \epsilon$



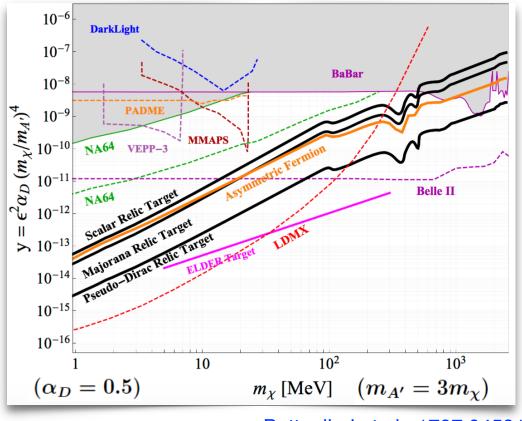
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Battaglieri et al., 1707.04591

* Babar/Belle (Babar, 1702.03327) $e^+e^- \rightarrow \gamma A', \ A' \rightarrow \mathrm{inv}$

* NA64 (1610.02988)

 $e^-p \rightarrow e^-pA', \ A' \rightarrow inv$ <u>Missing energy</u> measurement. 100 GeV electron beam

* LDMX (1610.02988)

 $e^-p \rightarrow e^-pA', \ A' \rightarrow \text{inv}$ <u>Missing momentum</u> measurement. few GeV electron beam, ~10¹⁶ EOT

* PADME, MMAPS, VEPP-3, DarkLight $e^+e^- \rightarrow \gamma A', A' \rightarrow \text{inv}$ <u>Missing mass</u> measurement. * See talk by R.Corliss 11/15

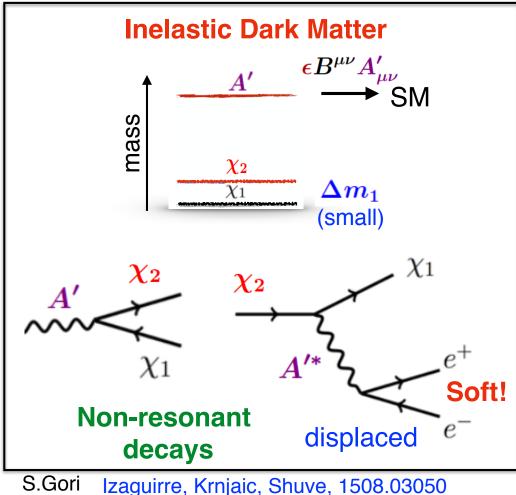
Beyond minimal dark photon models

(non-secluded*)

*DM is lighter

The non-secluded case can predict a <u>richer phenomenology</u> if the dark sector is not minimal

Examples:

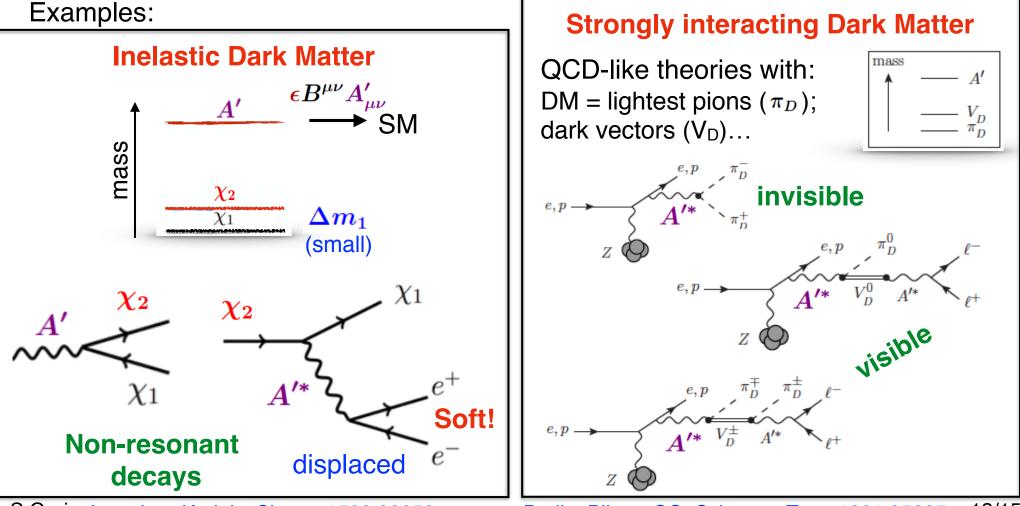


Beyond minimal dark photon models

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S.Gori Izaguirre, Krnjaic, Shuve, 1508.03050

Berlin, Blinov, SG, Schuster, Toro, 1801.05805 12/15

Beyond minimal dark photon models (secluded*) *DM is beau

*DM is heavier

The dark gauge boson can couple non-universally to the SM fermions $Z'e^+e^-
eq Z'\mu^+\mu^-
eq Z' au^+ au^-$

Example: models based on the L_{μ} - L_{τ} gauge symmetry \implies more hidden!

Motivations: B-physics anomalies, neutrino mass model building, ...

Free parameters of the model: g', m_{Z'}

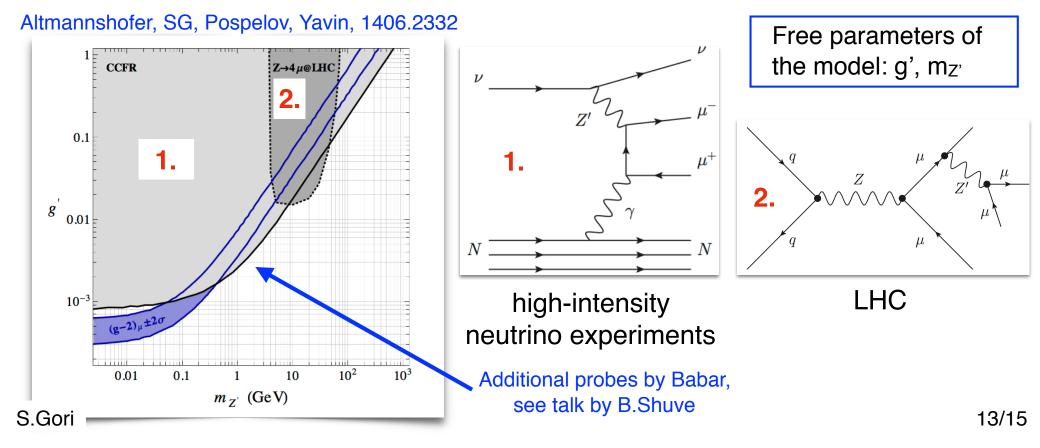
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Many other probes of dark sectors

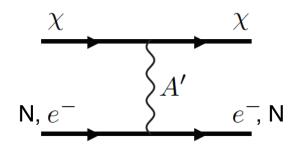
In this talk,

we have focused on the direct production and detection of dark photons

Additional probes include off-shell dark photon effects:

1. DM direct detection experiments:

Main challenge is that light (< 1GeV) DM deposits only a small recoil energy in detector

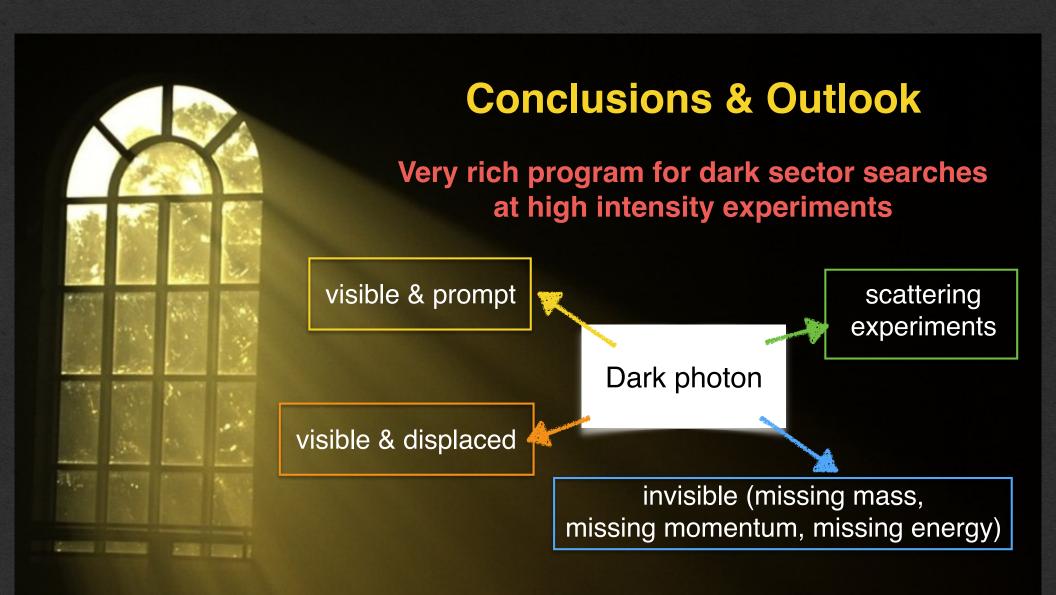


New dedicated experiments aim to see electron/nucleon recoils at lower energy than typical backgrounds (Sensei, Damic-1K, Ptolemy, SuperCDMS, ...)

2. Beam-dump experiments: direct production of a dark photon $A' \rightarrow \chi \bar{\chi}$ and DM - electron & nucleon scattering in a downstream detector (MiniBooNE, BDX (talk by M.Bondi), Coherent, Icarus, ...)

Beyond dark photon models:

Many additional tests for model with dark scalars, right-handed neutrinos, axions, ...



Complementarity with the high energy experimental program & the DM direct detection program