

A Pseudo-Dirac Bino and Neutrino Masses at the LHC

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Pilar Coloma, SI, *PRL*. 117 no.11 111803, *arXiv:1606.06372*

Julia Gehrlein, Patrick Fox, SI, *in preparation*

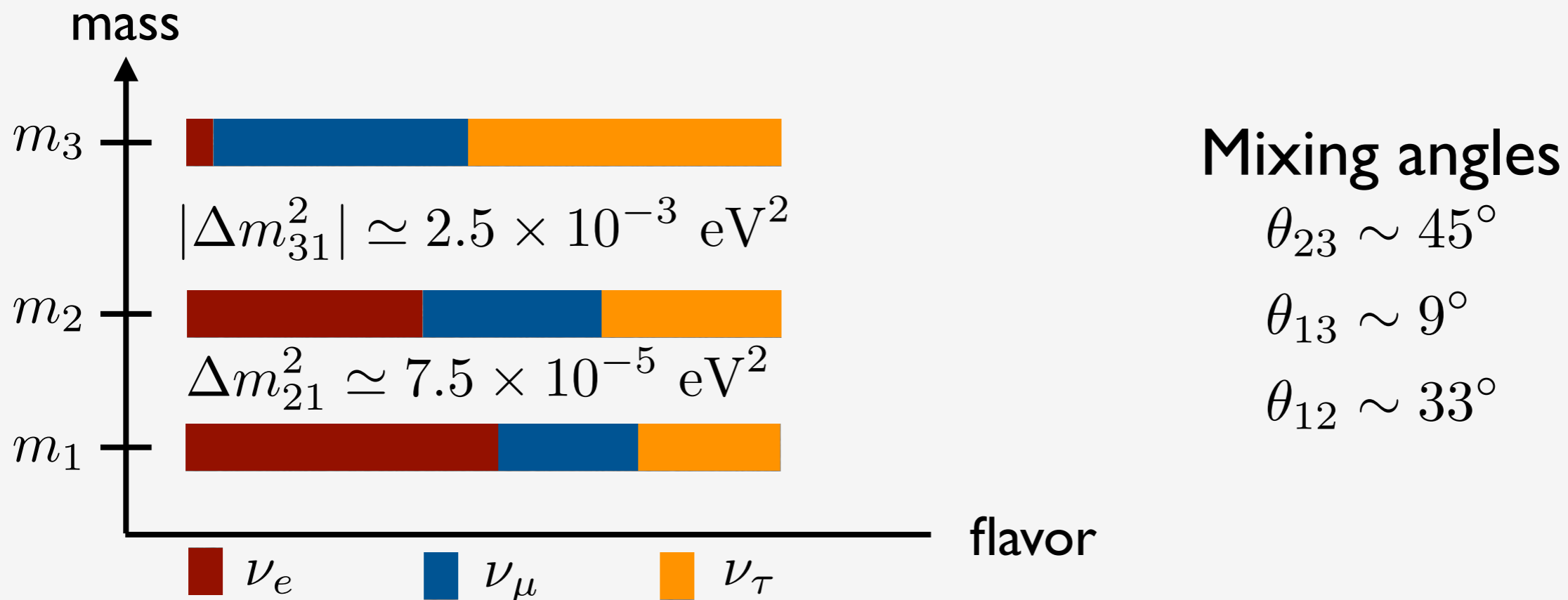
Neutrinos have mass

PDG review, www.nu-fit.org, ...

Mass eigenstates are different than flavor eigenstates



Neutrinos are massive particles



Origin of the masses is not known

Massless in the SM — no right-handed neutrinos

SM singlet fermions

1- Can have Dirac masses

$$y_\nu \bar{\ell} \tilde{\phi} \nu_R \rightarrow m_\nu \bar{\nu}_L \nu_R \quad \longleftrightarrow \quad y_\nu \sim 10^{-12}$$

Fermion masses are technically natural...

‘We’ don’t like very small numbers

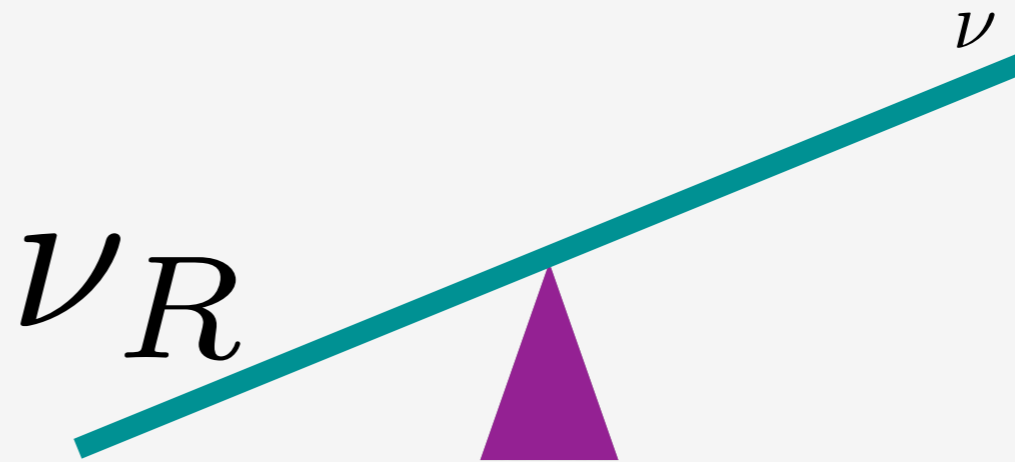
Origin of the masses is not known

2- Or Majorana masses — Seesaw mechanism

$$y_\nu \bar{\ell} \tilde{\phi} \nu_R + \frac{1}{2} M_R \nu_R^c \nu_R \quad \longrightarrow \quad m_\nu \sim \frac{y_\nu^2 v^2}{M_R}$$

$$y_\nu \sim \mathcal{O}(1)$$

$$M_R \sim \mathcal{O}(10^{14} \text{ GeV})$$



Majorana particles are their own antiparticles

Lepton number is violated

Particle numbers are accidental symmetries of the SM...

Origin of the masses is not known

3- Or pseudo-Dirac masses — Inverse Seesaw Mechanism

Let's say lepton number is (approximately) conserved

SM singlets:

$$\begin{array}{cc} N, & N' \\ \downarrow & \downarrow \\ L = 1 & L = -1 \end{array}$$

$$Y_N \bar{\ell} \tilde{\phi} N + \epsilon Y_{N'} \bar{\ell} \tilde{\phi} N' + M_D \bar{N} N'^c + \mu \bar{N} N^c + \mu' \bar{N}' N'^c$$

lepton number violation

$$\psi = \begin{pmatrix} N \\ N'^{\dagger} \end{pmatrix} : \text{pseudo-Dirac fermion}$$

Origin of the masses is not known

3- Or pseudo-Dirac masses — Inverse Seesaw Mechanism

Light neutrino masses:

$$m_\nu \sim \frac{\epsilon Y_N^T Y_{N'} v^2}{M_D} + \mathcal{O}\left(\frac{Y_N^T \mu Y_N v^2}{M_D^2}\right)$$

For $Y_N \sim Y_{N'} \sim 1,$
 $M_D \sim \text{TeV}$

need:

$$\mu \sim \text{keV}$$

$$\epsilon \sim 10^{-12}$$

Why so small?

What we need

(Usually) SM singlet fermions as
right-handed neutrinos

High mass scale

and/or

Small (lepton-number-violating) parameters

makes LHC searches very hard (impossible?)

Hall, Randall, *Nuc.Phys.B-352.2 1991*

Kribs, Poppitz, Weiner, *arXiv: 0712.2039*

Frugiuele, Gregoire, *arXiv:1107.4634*

SI, McKeen, Nelson, *arXiv: 1407.8193*

SI, John March-Russell, *arXiv:1604.00009*

My favorite model for everything!

$U(1)_R$ -symmetric SUSY

Pilar Coloma, SI, *PRL. 117 no.11 111803*

Has Dirac gauginos

Dirac gauginos are awesome - less tuning for heavier stops

Solves SUSY CP and flavor problems, ...

$U(1)_{R-L}$ - symmetric SUSY

SUSY particles are charged under $U(1)_{R-L}$



Superfields	$SU_c(3)$	$SU_L(2)$	$U_Y(1)$	$U(1)_R$	$U(1)_{R-L}$
L_i	1	2	-1/2	1	0
E_i^c	1	1	1	1	2
H_u	1	2	1/2	0	0
$W_{\tilde{B}}^\alpha$	1	1	0	1	1
$\Phi_S = \phi_S + \theta S$	1	1	0	0	0
$W'_\alpha = \theta D$	1	1	0	1	1



Spurion D-term

SUSY is broken in a hidden sector

Dirac bino mass

No Majorana gaugino masses due to the R -charges

Dirac masses come from the spurion D-term

$$\int d^2\theta c \frac{W'_\alpha}{\Lambda_M} W_{\tilde{B}}^\alpha \Phi_S \rightarrow \boxed{\frac{cD}{\Lambda_M}} \tilde{B} S$$

Λ_M : messenger scale

Dirac bino mass: M_D

$$\begin{array}{l} \tilde{B} \equiv (1, 1, 0)_{+1} \\ S \equiv (1, 1, 0)_{-1} \end{array} \quad \longrightarrow \quad \Psi = \begin{pmatrix} \tilde{B} \\ S^\dagger \end{pmatrix} \quad : \text{Dirac bino}$$

$U(1)_{R-L}$ must be broken

...because (anomaly mediation)

(Small) Majorana mass
for the bino

$$m_{\tilde{B}} = \frac{\beta(g_Y)}{g_Y} m_{3/2}$$

$m_{3/2}$: gravitino mass

Can also have a singlino Majorana mass

$$m_{\tilde{B}} \sim m_S \ll M_D$$

$U(1)_{R-L}$ is only
approximately broken

$$\Psi = \begin{pmatrix} \tilde{B} \\ S^\dagger \end{pmatrix} \quad : \text{pseudo-Dirac bino}$$

All the mass terms

$$\begin{aligned}
 -\mathcal{L} \supset & \frac{f_i M_D}{\Lambda_M} \ell_i h_u \tilde{B} + \frac{d_i m_{3/2}}{\Lambda_M} \ell_i h_u S \\
 & + M_D \tilde{B} S + m_{\tilde{B}} \tilde{B} \tilde{B} + m_S S S
 \end{aligned}
 \begin{array}{l}
 \nearrow \\
 \nearrow
 \end{array}
 U(1)_{R-L} \text{ violating}$$

This is an Inverse SeeSaw scenario!

$U(1)_{R-L}$ violation $\propto m_{3/2}$ \longrightarrow important for small neutrino masses

$\Psi = \begin{pmatrix} \tilde{B} \\ S^\dagger \end{pmatrix}$: We call this “bivo” (pronounced exactly like ‘bino’)
(like ‘too’ and ‘two’)

Neutrino masses

$$m_1 = 0 \quad \longrightarrow \quad 1 \text{ neutrino is massless}$$

$$m_2 = \frac{m_{3/2} v^2}{\Lambda_M^2} (1 - \rho)$$
$$m_3 = \frac{m_{3/2} v^2}{\Lambda_M^2} (1 + \rho)$$

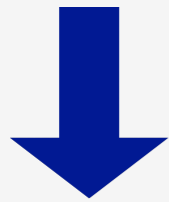
Neutrino masses are prop
to the gravitino mass

$\rho \simeq 0.7$ from mass splittings

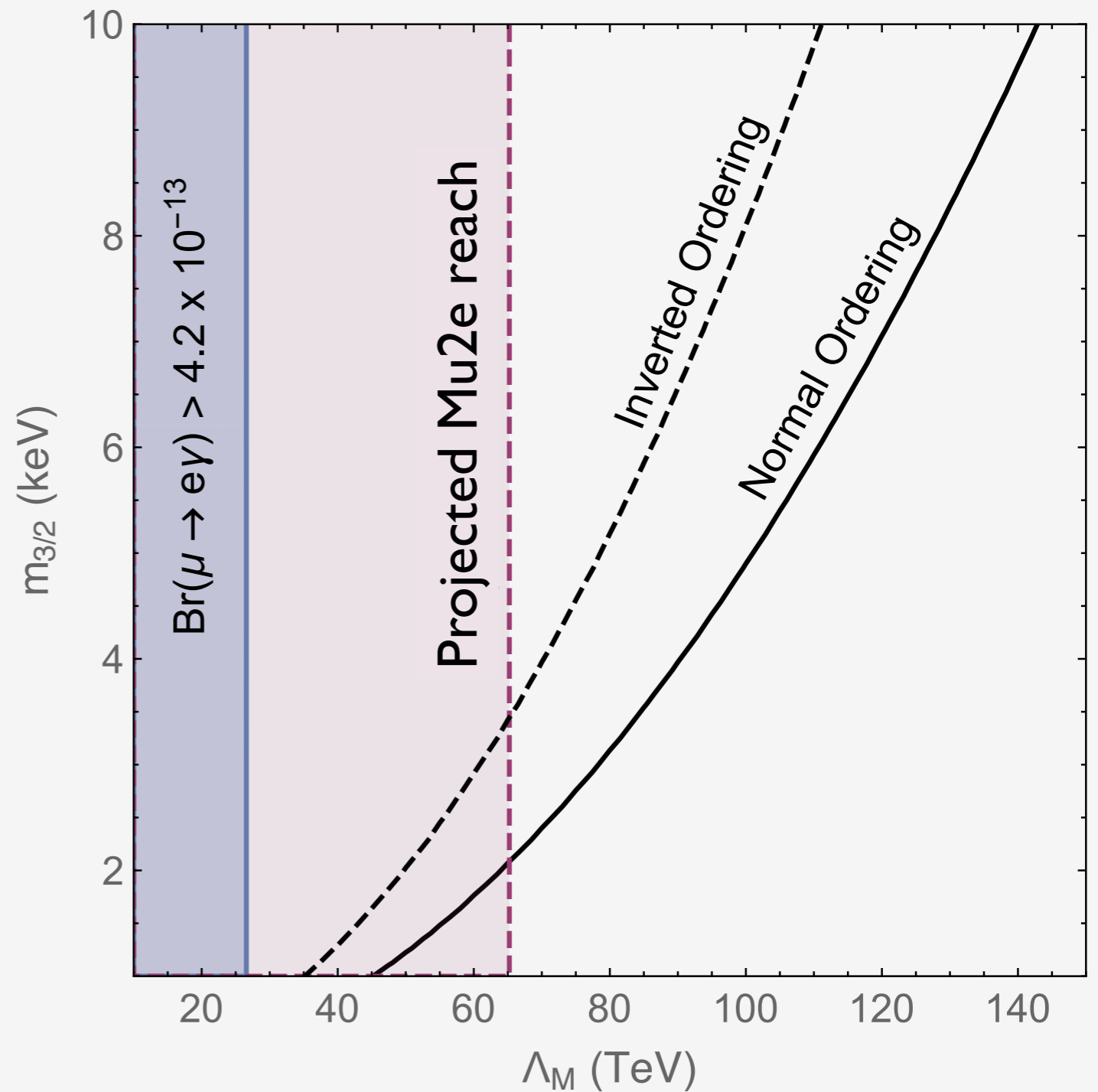
No dependence on M_D

Parameter space

Neutrino masses
+
Lepton flavor
violation constraints



$$m_{3/2} \sim \mathcal{O}(\text{keV})$$
$$\Lambda_M \sim \mathcal{O}(100 \text{ TeV})$$



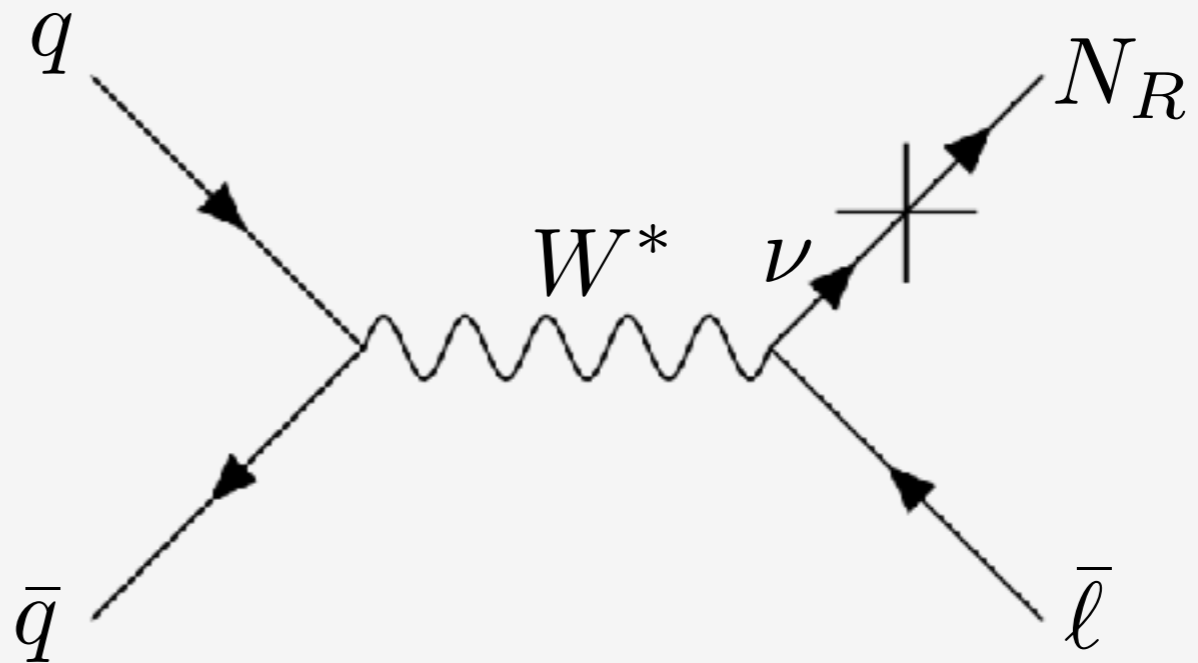
No constraints from neutrinoless double-beta decay

LHC Phenomenology of Bivo

Julia Gehrlein, Patrick Fox, SI, *in preparation*

LHC pheno: production

RH neutrinos are usually produced via mixing with the SM neutrinos



Pay a mixing price on top of EW interactions:

$$\theta^2 \sim 10^{-5}$$

There are better ways to produce a bino!

LHC pheno: production

Assumptions for the phenomenology:

Lightest neutralino is a pure bino

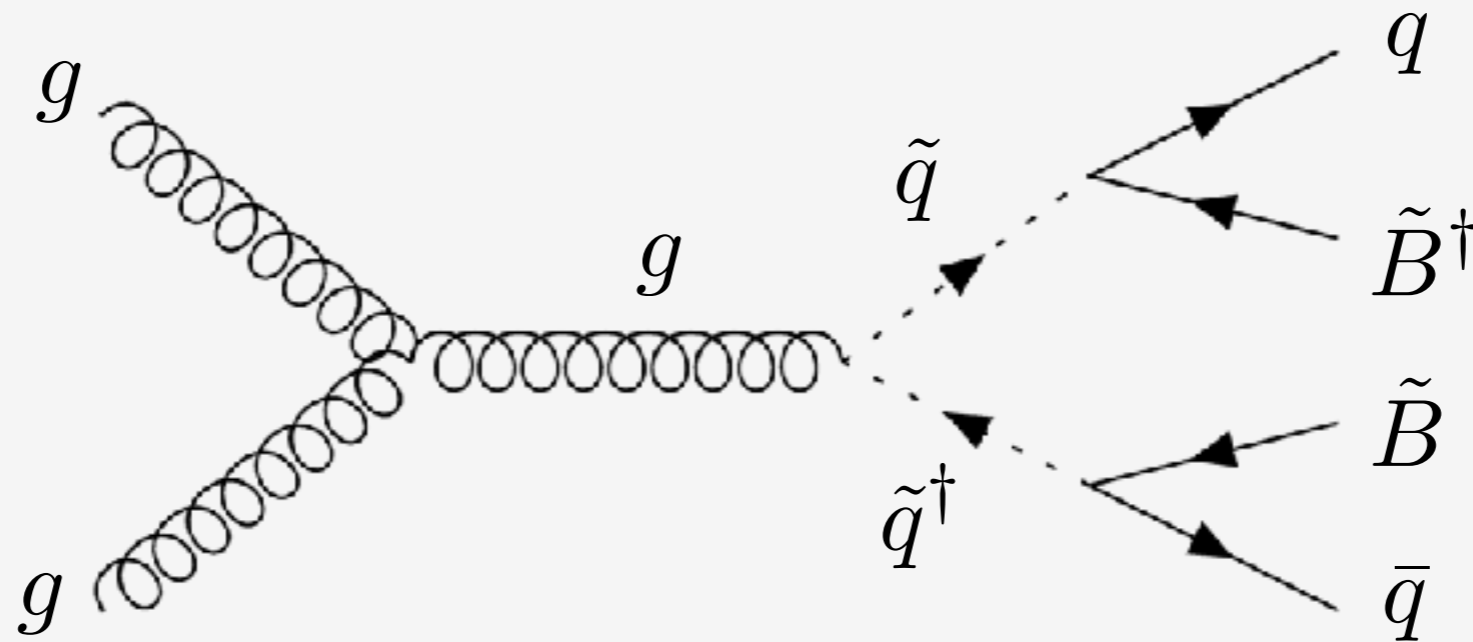
Gravitino is the LSP - a few keV

Bino is the NLSP

Degenerate squarks

LHC pheno: production

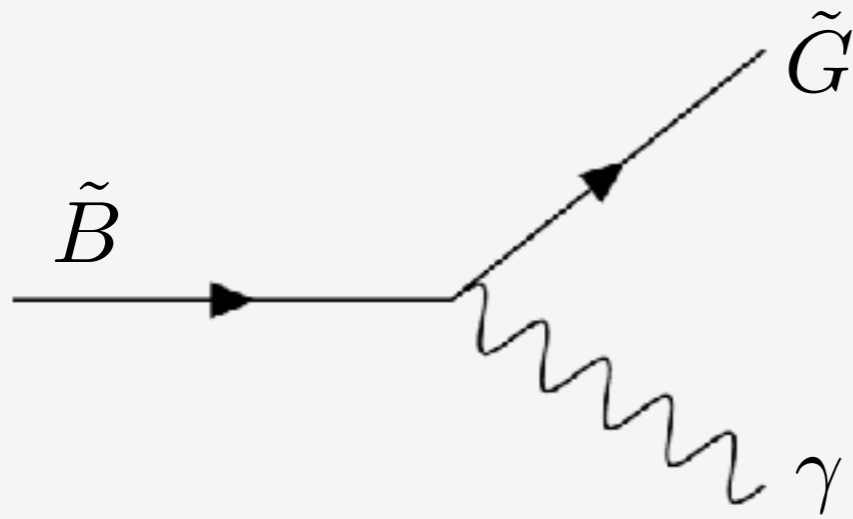
Can produce bino via squark decays:



Gluon fusion is the main production channel at 14 TeV LHC

LHC pheno: decays

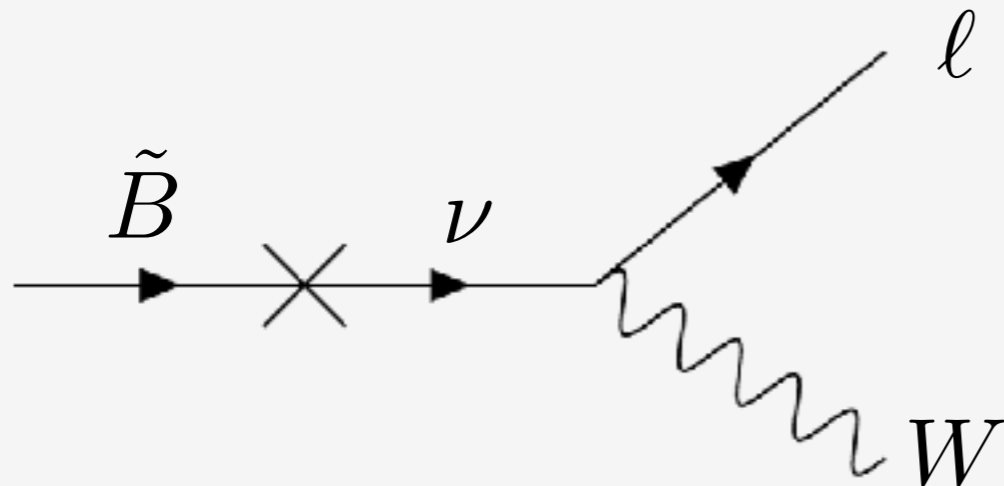
1) Decays to a gravitino and a photon



very small

$$\Gamma(\tilde{B} \rightarrow \tilde{G}\gamma) \sim \frac{M_D^5}{M_{Pl}^2 m_{3/2}^2} \sim 10^{-8} \text{ eV}$$

2-4) Decays to $W\ell$, $Z\nu$, $h\nu$



etc

total width:

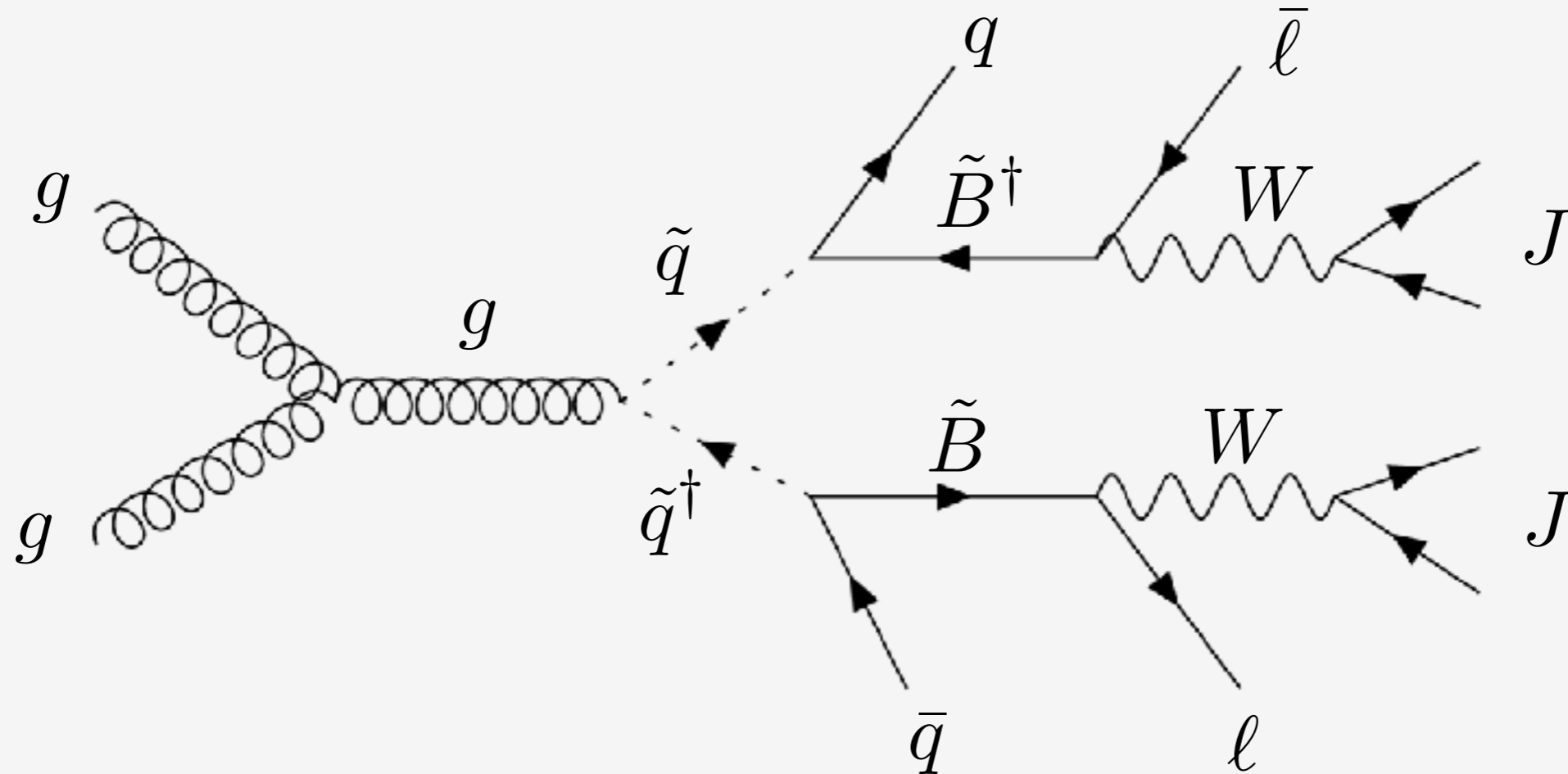
$$\Gamma_{tot} \sim \frac{M_D^3}{\Lambda_M^2} \sim O(\text{MeV})$$

LHC pheno: signals

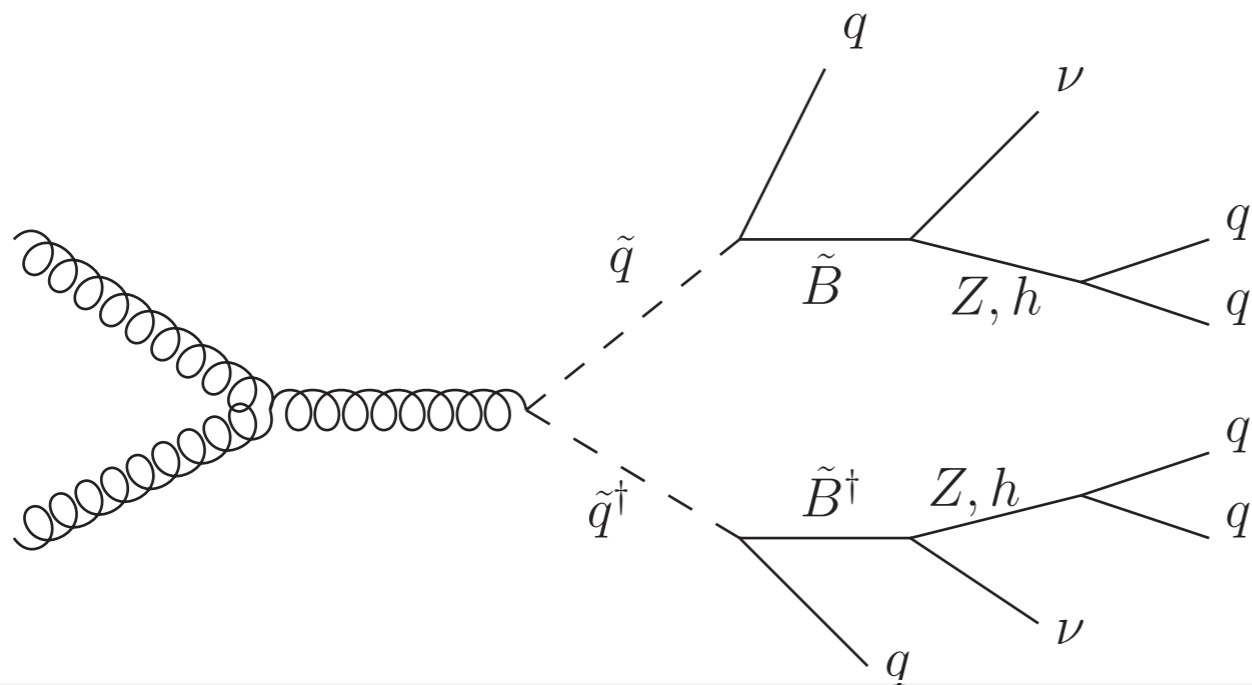
Bino decays promptly, no displaced vertices

A combination of jets, leptons and missing energy is expected

e.g.



LHC pheno: signals

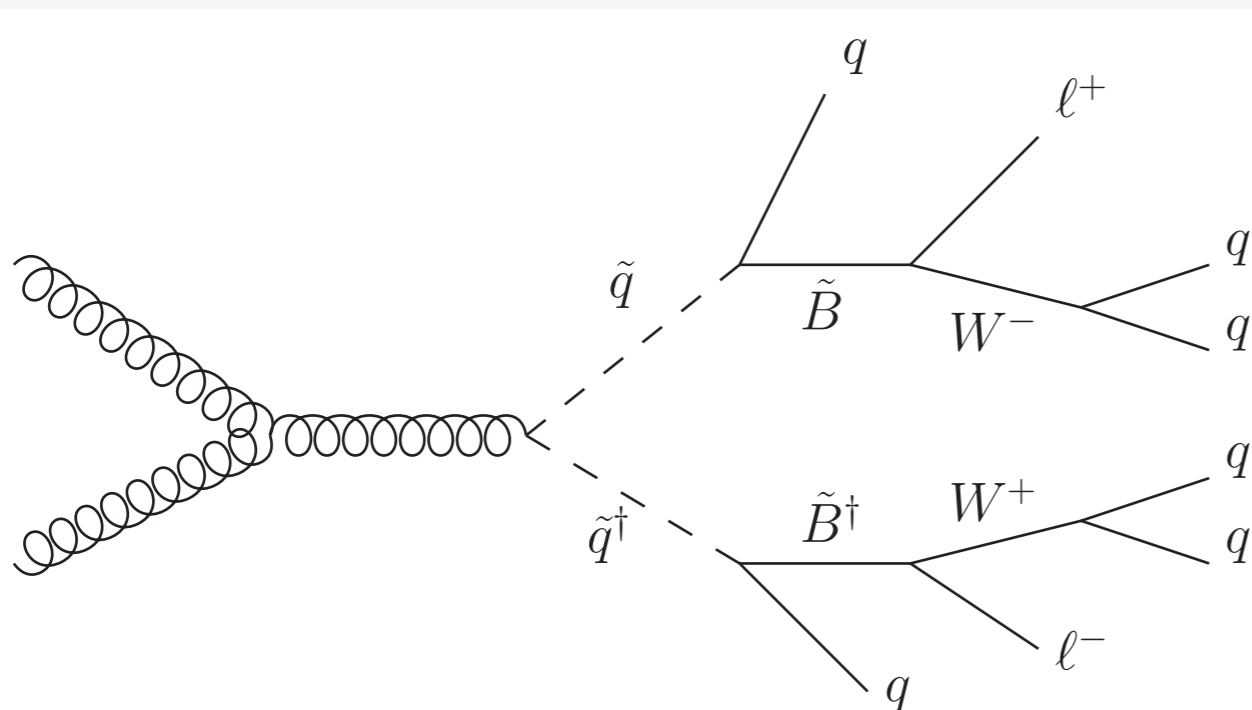


~ 20% branching ratio



ATLAS-CONF-2017-022

CMS-SUS-16-033



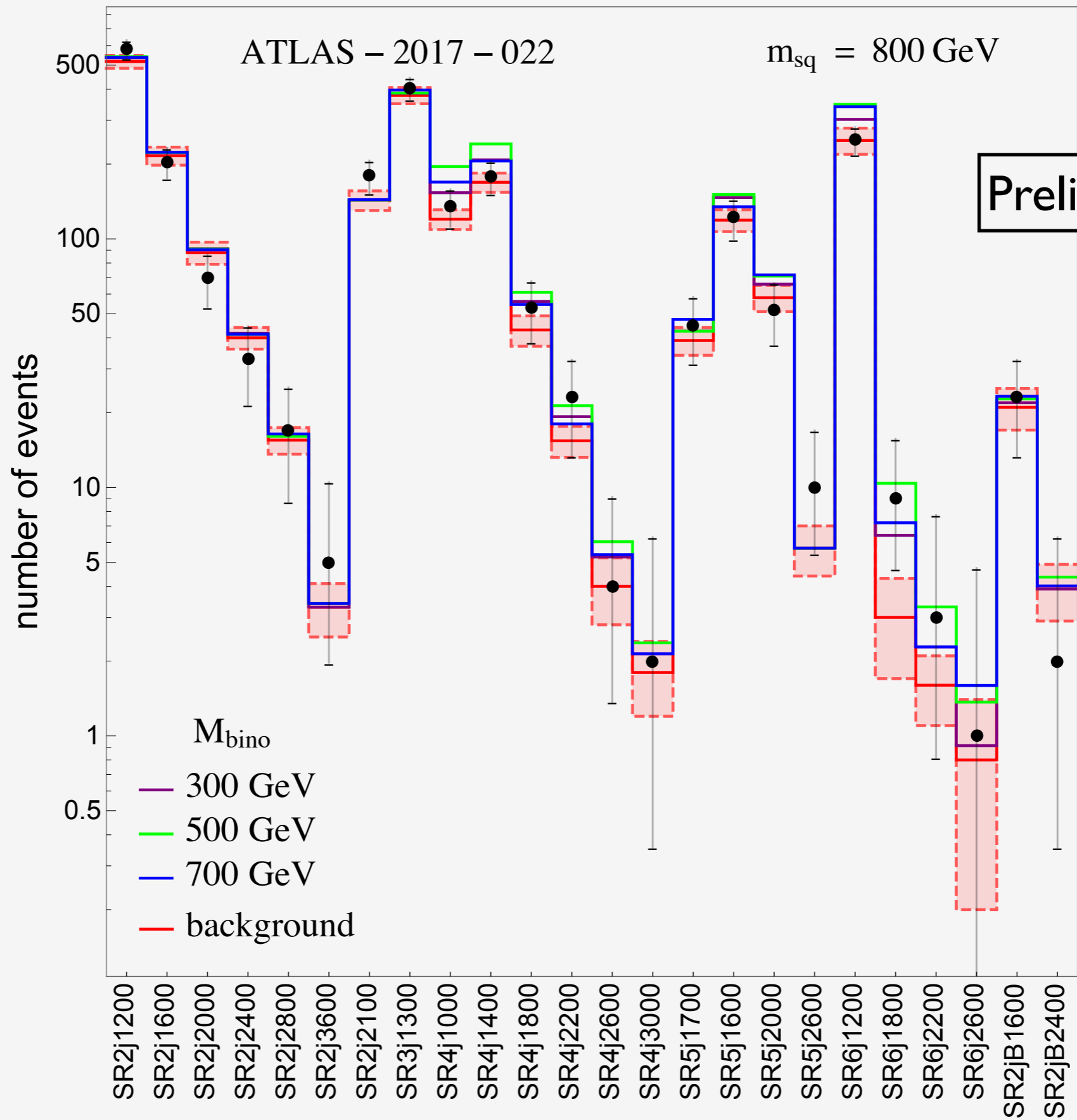
~ 3% branching ratio



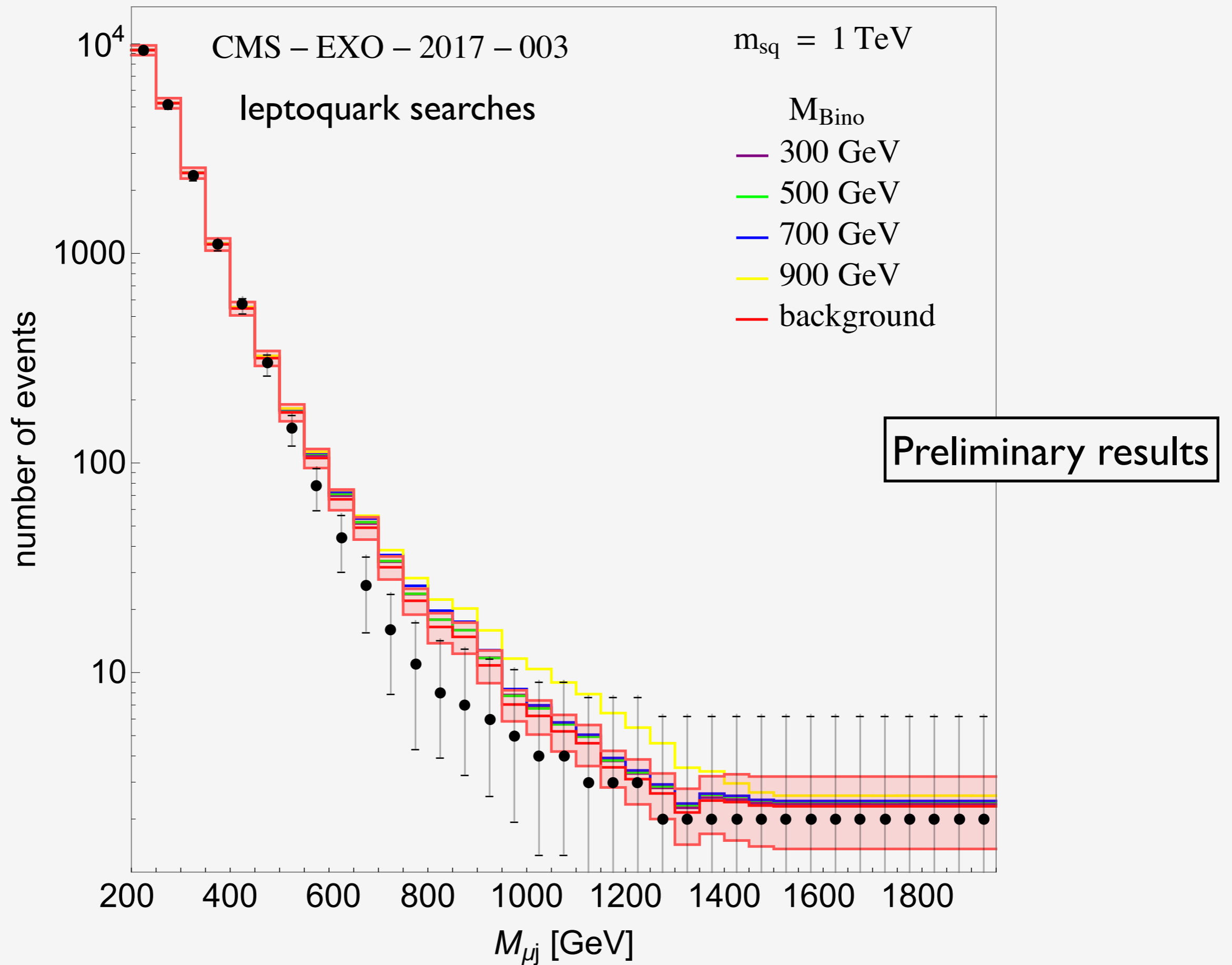
CMS-PAS-EXO-17-003

CERN-EP-2016-074 (ATLAS)

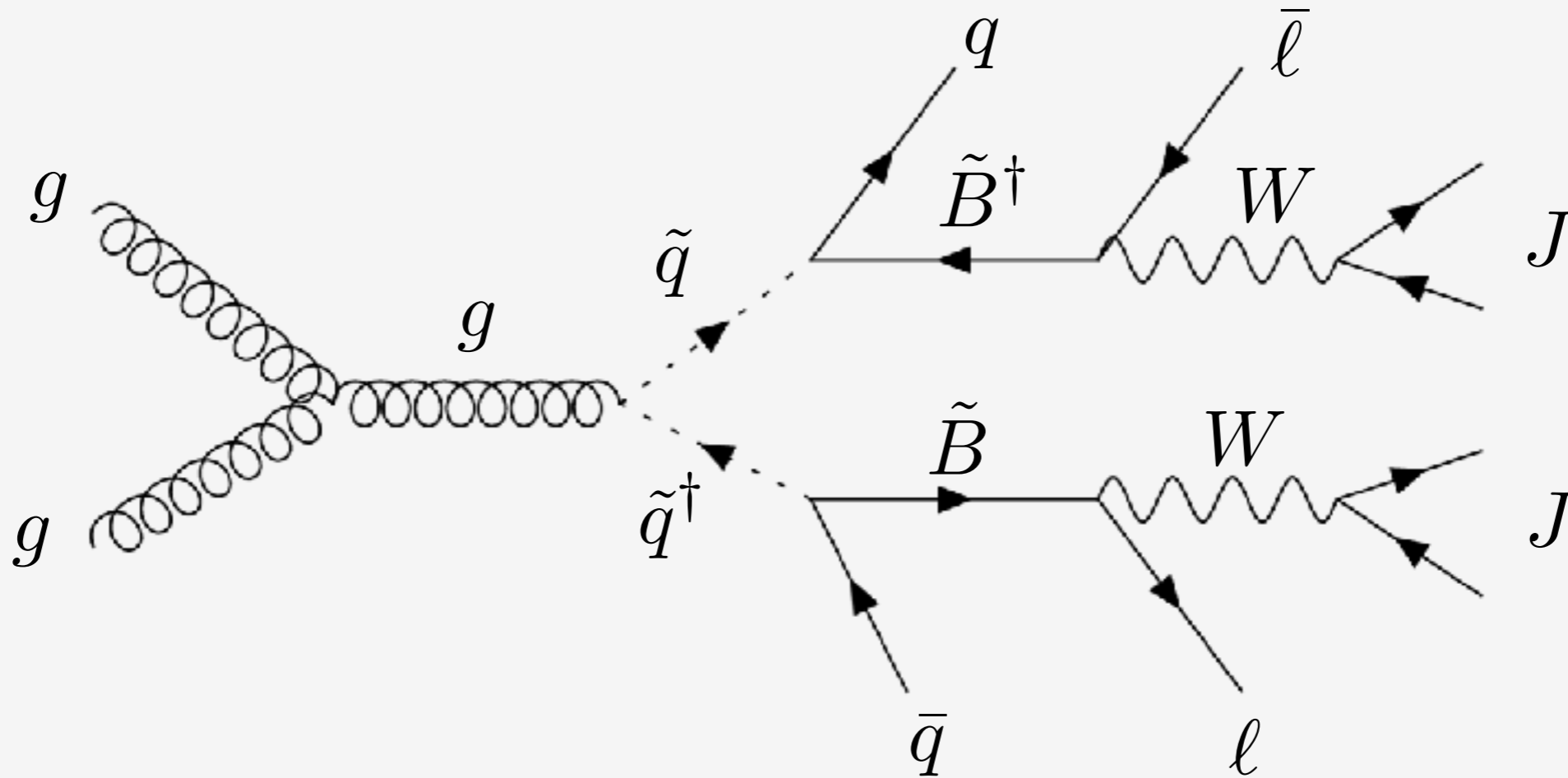
Huge thanks to Angelo Monteux for the analysis codes!



Preliminary results



Smoking gun signal



electron : muon : tau
ratios are fully determined by the neutrino sector:
1:2:1

Preliminary results

binos heavier than squarks

