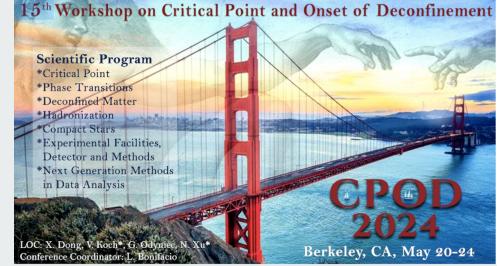




p_{T} - p_{T} Correlators at High Baryon Density Region

Rutik Manikandhan (University of Houston) for the STAR Collaboration











STAR-FXT Setup

Transverse Momentum Correlations

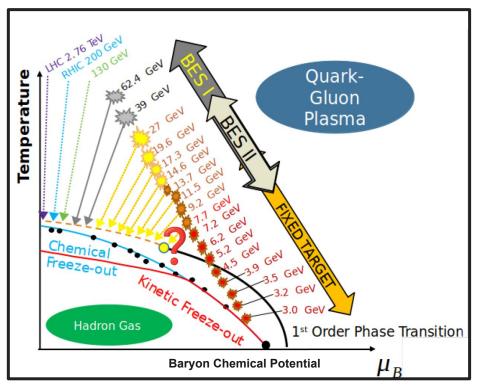
Results



Phases of QCD Matter



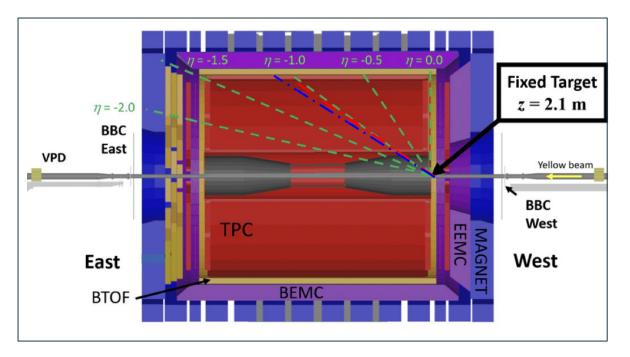
- BES-II collider program at the Relativistic Heavy-Ion Collider scans phase space of QCD matter by colliding gold ions at varying energies.
- Seeking to map onset of deconfinement, and the predicted QCD critical point.
- ★ The BES-II collider program provided the energies $\sqrt{s_{NN}} >= 7.7$ GeV and the BES-II FXT program provided the ones below, down to $\sqrt{s_{NN}} = 3$ GeV.



STAR-FXT Setup



- Gold Target fixed at west end of the detector
- ★ TPC Acceptance :
 ▶ η : [-2,0] (lab frame)
- ◆ PID Acceptance :
 ▶ η : [-1.5,0] (lab frame)
- Mid rapidity :

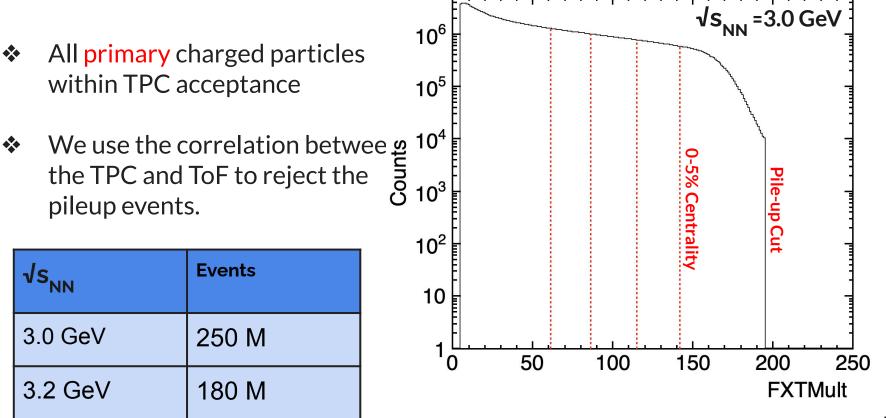


https://www.star.bnl.gov

η ≅ -1.05 (3.0 GeV)
 η ≅ -1.13 (3.2 GeV)

Centrality Definition





Transverse Momentum Correlations



- Transverse momentum correlations have been proposed as a measure of thermalization and as a probe for the critical point of quantum chromodynamics [1].
- Correlation measurements generally have finer 'resolution' than fluctuation measurements and can be looked at more differentially [2].
- The correlator is the mean of covariances of all pairs of particles i and j in the same event with respect to the mean.

 $C_m = <\Delta p_{t,i}, \Delta p_{t,j}> \ < (p_{t,i} - < p_t>)(p_{t,j} - < p_t>)>$



[1]: ALICE, Phys. Part. Nuclei 51,2020
[2]: Pruneau CA. Data Analysis Techniques for Physical Scientists. Cambridge University Press; 2017.

Transverse Momentum Correlations

STAR

- Dynamical fluctuations have no contributions from statistical fluctuations.
- Statistical fluctuations are Poissonian.
- Two body correlation function.

$$<\Delta p_{t1},\Delta p_{t2}>= \ \int dp_1 dp_2 rac{r(p_1,p_2)}{< N(N-1)>} \Delta p_{t1} \Delta p_{t2}$$

S. Gavin, Phys. Rev. Lett. 92, 162301

$$r(p_1,p_2)=N(p_1,p_2)-N(p_1)N(p_2)$$

Transverse Momentum Correlations



Locally thermalized systems. (at all energies?)

$$<\Delta p_{t,i}, \Delta p_{t,j}>=Frac{R}{1+R}$$

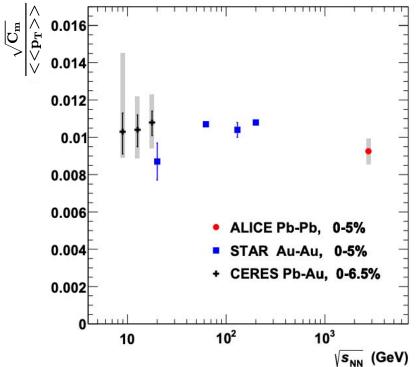
- F depends on the ratio of the correlation length (ζ_T) to the transverse size.
- R is the scaled variance and depends on N_{part}
- If matter is locally equilibrated in the most central collisions, F(ζ_T) is energy independent.

$$R = rac{<\!N^2>-<\!N>^2-<\!N>}{<\!N>^2}$$

prated in the

$$\zeta_{T}$$
) is
S. Gavin, Phys. Rev. Lett. 92, 162301 CONST!!!

- The correlation observable may have a dependence on energy, so we scale it with <<p_>>>.
- Efficiency independent observable.
- Make a direct comparison with the CERES and ALICE.

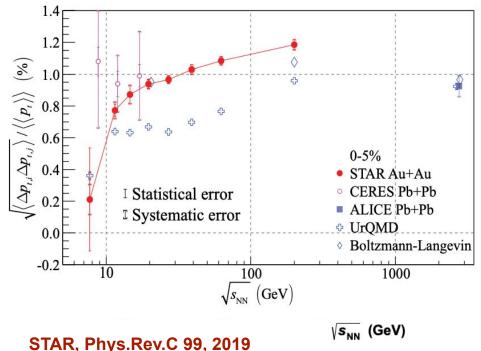


ALICE, Eur. Phys. J. C 74, 2014





- The correlation observable may have a dependence on energy, so we scale it with <<p_T>>.
- Efficiency independent observable.
- Make a direct comparison with the CERES and ALICE.
- A significant beam energy dependence was found for dynamical correlations.



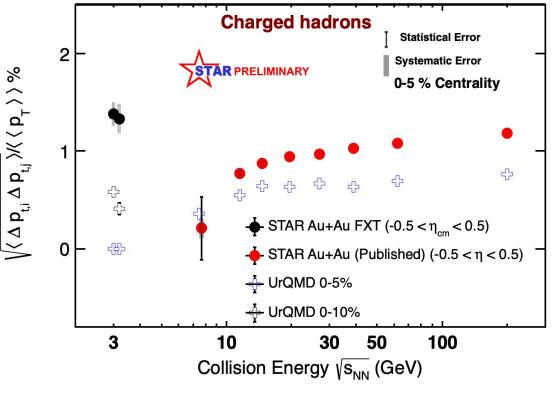


- We see a departure from monotonicity
- Change in correlation length ζ_{T} ?

 $< p_T > \implies I_{eff}$

 $T_{eff} = T_{kin} + m_0 < eta_T >^2$

 Temperature fluctuations should be reflected in p_T fluctuations.



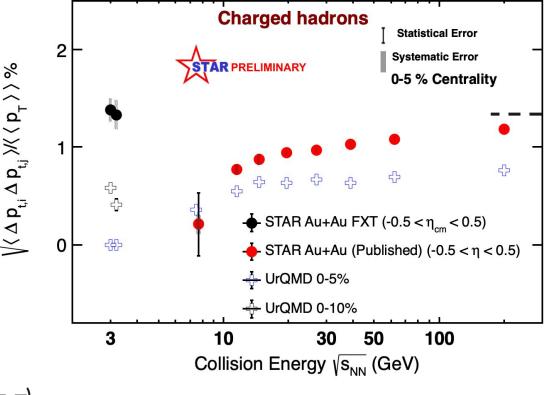
Sumit Basu et. al., Phys.Rev.C 94, 2016



- F(ζ_T) and R to be constant as a function of collision energy.
- $F(\zeta_T) = 0.046$
- R = 0.0037 (Central Au+Au at 200 GeV)

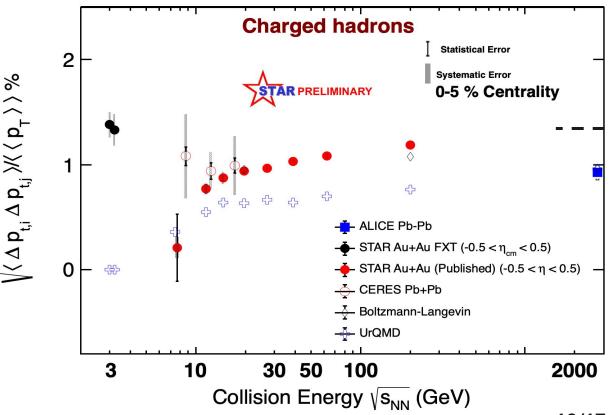
S. Gavin, Phys. Rev. Lett. 92, 162301

$$(rac{F(\zeta_T)R}{1+R})^{1/2}= ext{constant}$$
 (---)





- CERES in agreement with STAR.
- Boltzmann-Langevin implies thermalization.
- ALICE lower than STAR, due to different N_{part}

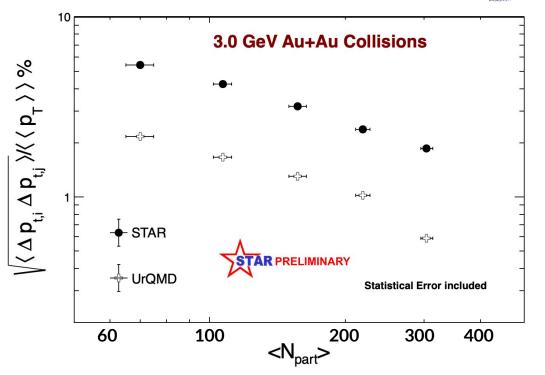


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Correlator Vs Centrality

STAR

- Monotonic increase in decreasing centrality.
- UrQMD underpredicts the data.
- ♦ UrQMD Acceptance:
 ▶ η:[-0.5,0.5] (Collider mode)
 - > p_T:[0.2,2.0] GeV/c



> All charged particles

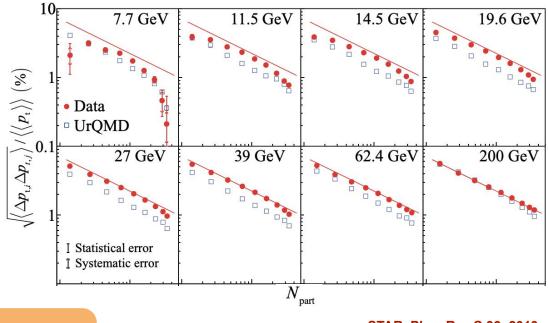
Correlator Vs Centrality



- Power law seems to describe the data at 200 GeV, implying an independent sources scenario.
- We see significant departure from this power law dependence at the lower energies.
- UrQMD tends to underpredict the data at all energies.

Power Law:

 $rac{\sqrt{C_m}}{<< n_{\pi}>>} \propto < N_{part}>^b$

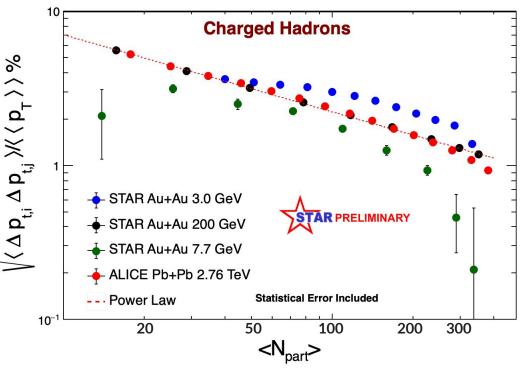


STAR, Phys.Rev.C 99, 2019

Correlator Vs Centrality



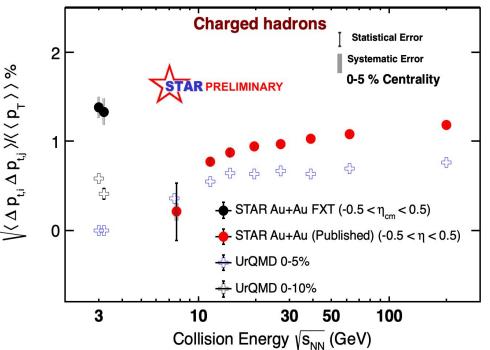
- Power law implies uncorrelated sources (b=-0.5).
- STAR data from 200 GeV Au+Au collision shows minimal deviation.
- Deviation increases as we go down the collision energy
- Deviation holds at STAR 3.0 GeV Au+Au collisions as well.



Conclusions

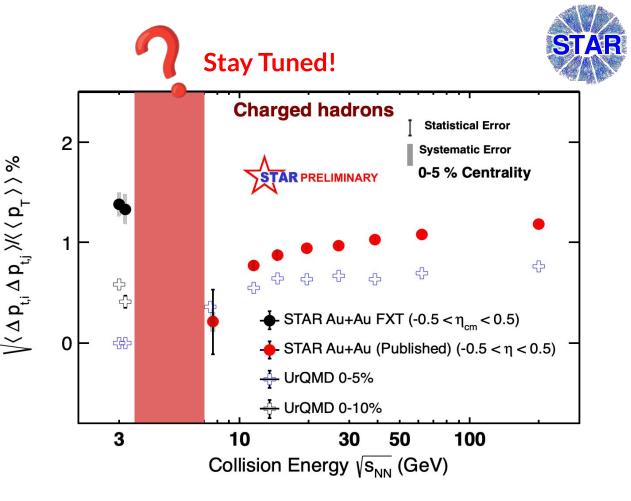


- First measurement of Δp_T-Δp_T
 correlators at high baryon density region
 - > $\Delta p_T \Delta p_T$ show a non-monotonic behaviour.
 - Possibility of correlation length changing in between ?
- We need to delve deeper into the disparity observed between UrQMD and experimental data at Fixed-Target (FXT) energies.



<u>Outlook</u>

- BES-II FXT energies are crucial to understand.
- Account for detector acceptance effects.
- Look into higher order moments.
- Thermal model predictions.



References



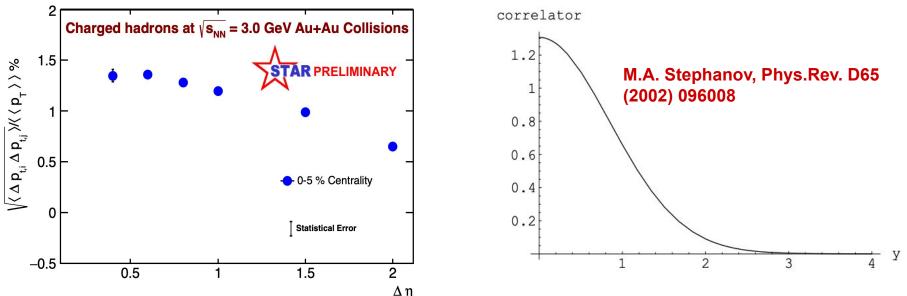
- 1. Temperature Fluctuations in Multiparticle Production Phys. Rev. Lett. 75, 1044
- 2. Incident energy dependence of pt correlations at relativistic energies Phys.Rev.C72:044902,2005
- 3. Event-by-event fluctuations in mean p_T and mean e_T in s(NN)**(1/2) = 130-GeV Au+Au collisions Phys.Rev.C 66 (2002) 024901
- 4. Collision-energy dependence of p_T correlations in Au + Au collisions at energies available at the BNL Relativistic Heavy Ion Collider Phys.Rev.C 99 (2019) 4, 044918
- 5. Event-by-event mean p_T fluctuations in pp and Pb-Pb collisions at the LHC Eur. Phys. J. C 74 (2014) 3077
- 6. Specific Heat of Matter Formed in Relativistic Nuclear Collisions Phys.Rev.C 94 (2016) 4, 044901
- 7. Baryon Stopping and Associated Production of Mesons in Au+Au Collisions at s(NN)**(1/2)=3.0 GeV at STAR Acta Phys. Pol. B Proc. Suppl. 16, 1-A49 (2023)
- 8. Traces of Thermalization from p_T Fluctuations in Nuclear Collisions S. Gavin, Phys. Rev. Lett. 92, 162301 (2004)



- BACKUP

Acceptance dependence



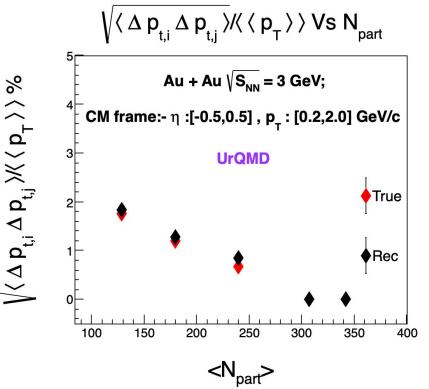


- The effect of primordial protons bring the correlator down for the whole acceptance.
- Closer to mid-rapidity where majority of the particle production takes place the value saturates.

Closure Test

- ★ The relative uncertainties $\sqrt{C_m} / \langle p_T \rangle >$ on are generally smaller than those on C_m because most of the sources of uncertainties lead to correlated variations of $\langle p_T \rangle >$ and C_m that tend to cancel in the ratio.
- Closure test was performed with UrQMD data, by incorporating 3.0 GeV efficiency curves.
- We see closure within the statistical error bars.
- No efficiency correction was employed on STAR Data.





Cumulants from moments



$$\left\langle \Delta p_{\mathrm{T},i} \Delta p_{\mathrm{T},j} \right\rangle = \left\langle \frac{\sum_{\substack{i,j,\ i\neq j}}^{N_{\mathrm{ch}}} (p_{\mathrm{T},i} - \left\langle \left\langle p_{\mathrm{T}} \right\rangle \right\rangle) (p_{\mathrm{T},j} - \left\langle \left\langle p_{\mathrm{T}} \right\rangle \right\rangle)}{N_{\mathrm{ch}} (N_{\mathrm{ch}} - 1)} \right\rangle_{\mathrm{ev}} = \left\langle \frac{Q_1^2 - Q_2}{N_{\mathrm{ch}} (N_{\mathrm{ch}} - 1)} \right\rangle_{\mathrm{ev}} - \left\langle \frac{Q_1}{N_{\mathrm{ch}}} \right\rangle_{\mathrm{ev}}^2, \quad (2)$$

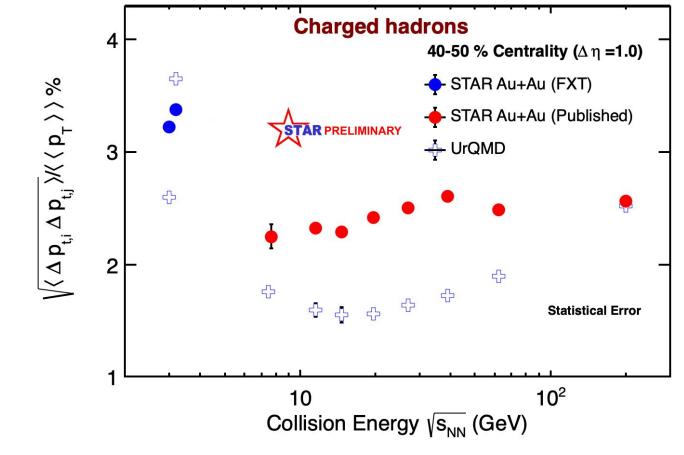
$$\langle \Delta p_{\mathrm{T},i} \Delta p_{\mathrm{T},j} \Delta p_{\mathrm{T},j} \rangle = \left\langle \frac{\sum_{\substack{i,j,k, \ i \neq j \neq k}}^{N_{\mathrm{ch}}} (p_{\mathrm{T},i} - \langle \langle p_{\mathrm{T}} \rangle \rangle) (p_{\mathrm{T},j} - \langle \langle p_{\mathrm{T}} \rangle \rangle) (p_{\mathrm{T},k} - \langle \langle p_{\mathrm{T}} \rangle \rangle)}{N_{\mathrm{ch}} (N_{\mathrm{ch}} - 1) (N_{\mathrm{ch}} - 2)} \right\rangle_{\mathrm{ev}}$$
$$= \left\langle \frac{Q_{1}^{3} - 3Q_{2}Q_{1} + 2Q_{3}}{N_{\mathrm{ch}} (N_{\mathrm{ch}} - 1) (N_{\mathrm{ch}} - 2)} \right\rangle_{\mathrm{ev}} - 3 \left\langle \frac{Q_{1}^{2} - Q_{2}}{N_{\mathrm{ch}} (N_{\mathrm{ch}} - 1)} \right\rangle_{\mathrm{ev}} \left\langle \frac{Q_{1}}{N_{\mathrm{ch}}} \right\rangle_{\mathrm{ev}} + 2 \left\langle \frac{Q_{1}}{N_{\mathrm{ch}}} \right\rangle_{\mathrm{ev}}^{3}, \quad (3)$$

 $Q_n = \sum_{i=1}^{N_{\mathrm{ch}}} p_{\mathrm{T},i}^n$.

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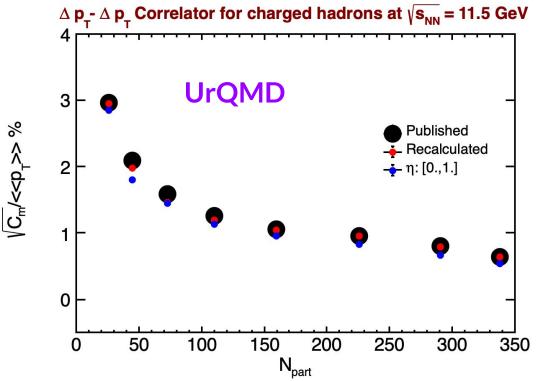
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UrQMD with asymmetric Acceptance

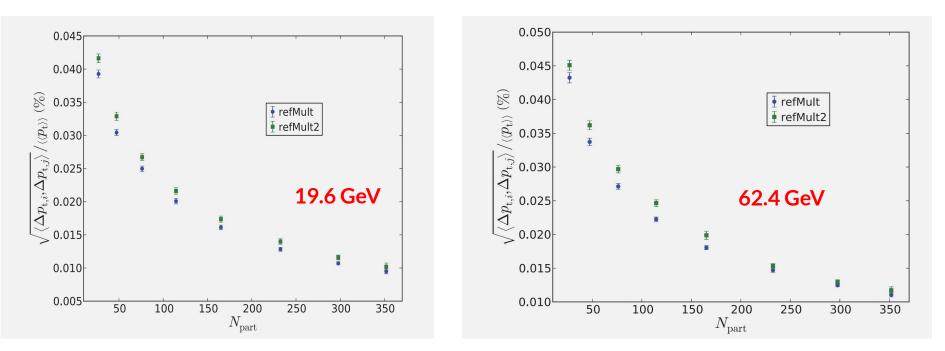
- To verify the UrQMD calculations, the analysis was carried out at a published energy.
- The analysis was also done with an asymmetric acceptance of **n** : [0,1]





Auto Correlation Studies

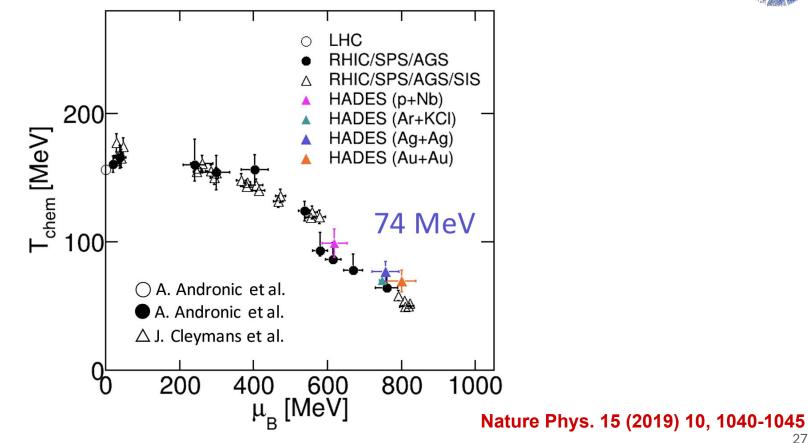




https://groups.nscl.msu.edu/nscl_library/Thesis/Novak,%20John.pdf

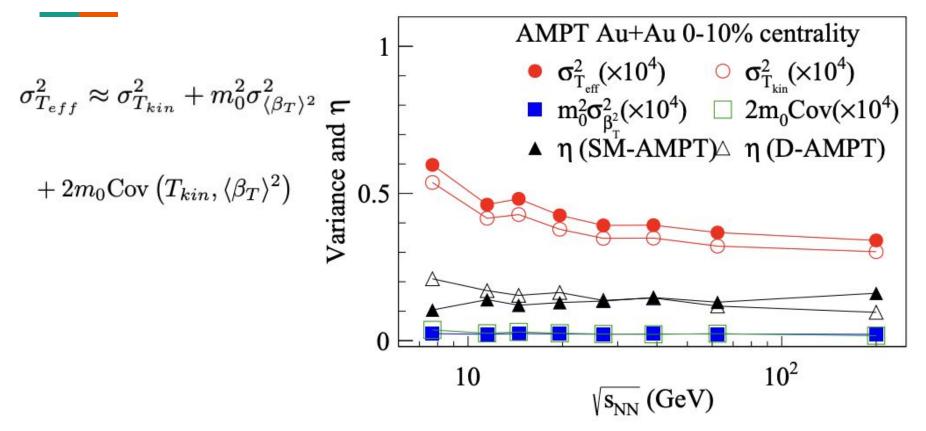
Vs μ_{B} chem





Contributions to temperature fluctuations





Phys.Rev.C 106 (2022) 1, 014910

Proton Multiplicity fluctuations



