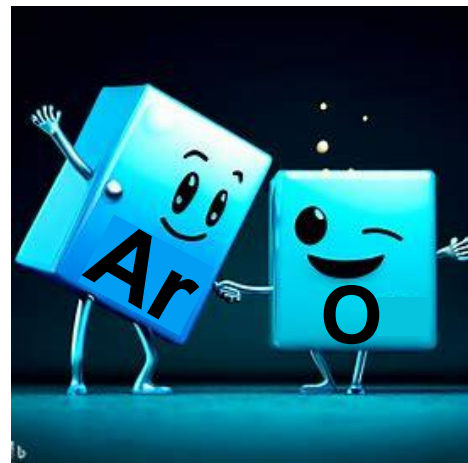


ND-GAr & Theia-ND

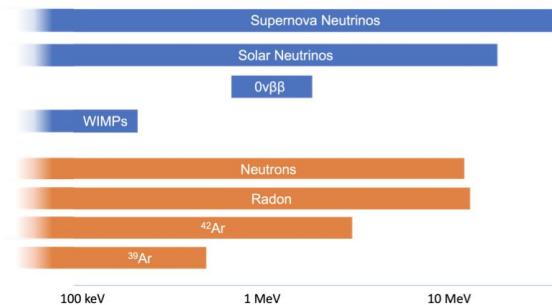
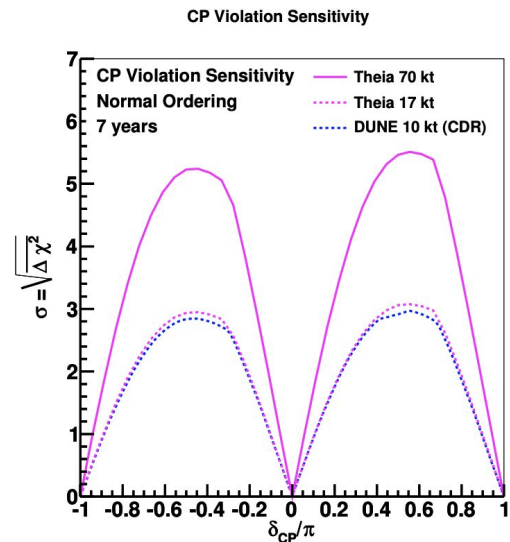
(A Match Made in Cavern?)

Mike Wilking
DUNE Phase 2 Workshop
Imperial College London
June 21st, 2023



FD Module of Opportunity

- DUNE Phase 2 goal is to produce 2 new FDs, upgrade the beam, and upgrade the ND
 - Ongoing US P5 process may end up weighing in on these elements with regard to US funding
- Additional funding sources / new collaborators can substantially improve the odds of completing DUNE Phase 2
 - A WbLS detector (Theia) would grow the collaboration and accessible funding sources
- The various proposed enhancements for DUNE Phase 2 all involve improving our access to low energy physics
 - However, many challenges exist to reach low background levels in a future LAr detector
 - Theia is specifically designed for low-E physics and will broaden our physics program (DSNB, SN burst, solar CNO, $0\nu\beta\beta$, ...) and provide a complementary target nucleus + event reconstruction for the LBL program
- New FD detector technologies require new ND capabilities



Brief Introduction to Theia

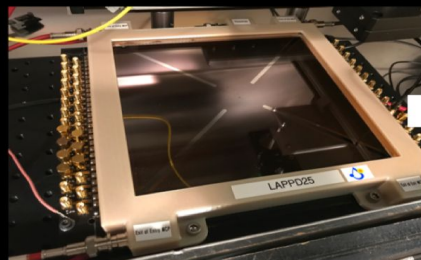
[The European Physical Journal
C 80 no. 5, \(May, 2020\)](#)

(slide from M. Wurm)

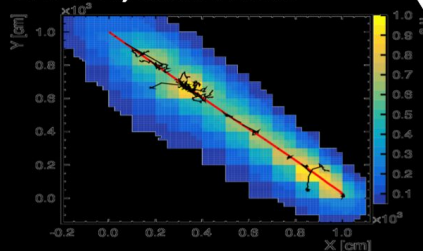
- Observe both Cherenkov and scintillation light for each event
- At High-E, operates much like Super-K
 - with additional information about below-Cherenkov-threshold hadrons
- Low-E threshold is lowered
 - More photons near and below Cherenkov threshold
- Much more detail in a recent Phase 2 biweekly talk from M. Wurm



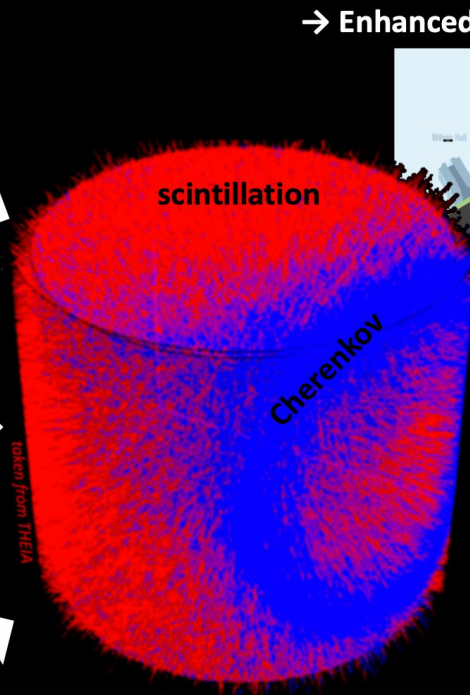
Novel target medium:
Water-based Liquid Scintillator



Novel light sensors:
LAPPDs, dichroicons

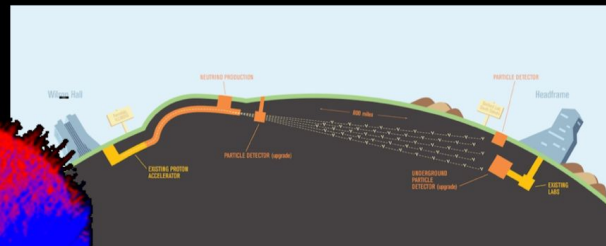


Novel reconstruction techniques



Large volume detector
able to exploit both
Cherenkov+Scintillation
signals

→ **Enhanced sensitivity to broad physics program**



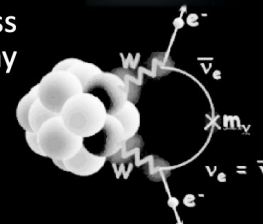
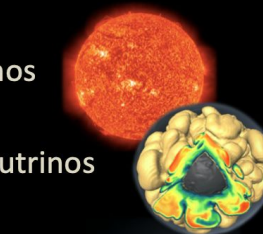
→ **Long-Baseline Oscillations**

→ Solar neutrinos

→ Supernova neutrinos

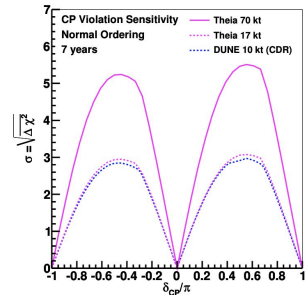
→ Diffuse SN neutrinos

→ Neutrinoless
Double-Beta Decay

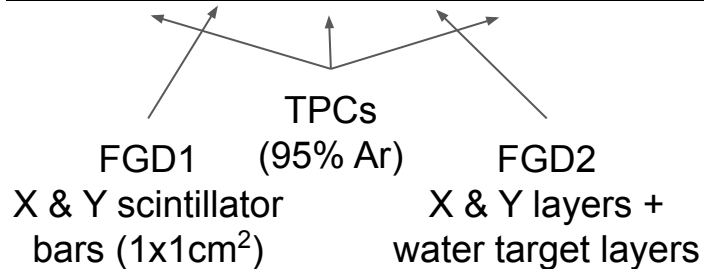
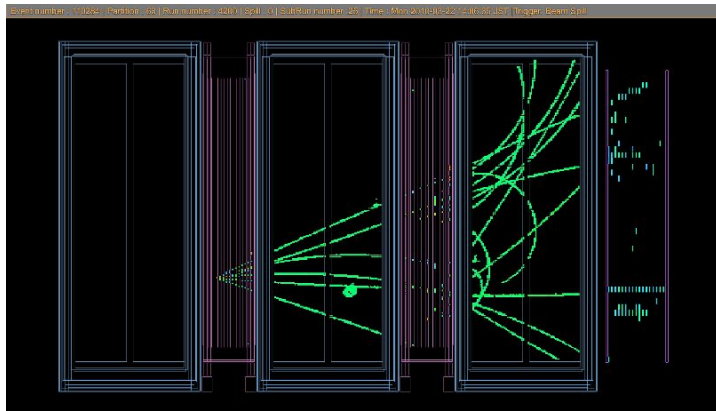


WbLS Near Detector Considerations

- A key component of LAr detectors is hadron calorimetry
 - Neutrino energy is the sum of the reconstructed lepton energy and the (corrected) deposited hadronic energy
- For water Cherenkov detectors, E_ν reconstruction is performed with above-Cherenkov particles
 - The Theia LBL sensitivity studies were performed without utilizing scintillation light
- The primary requirement for a Theia near detector is to measure above-Cherenkov-threshold particles
 - This is the approach used for the primary T2K / Hyper-K near detector
 - Additional external measurements of Cherenkov/scintillation ratio may be helpful
 - Large R&D program with several WbLS detectors is currently underway

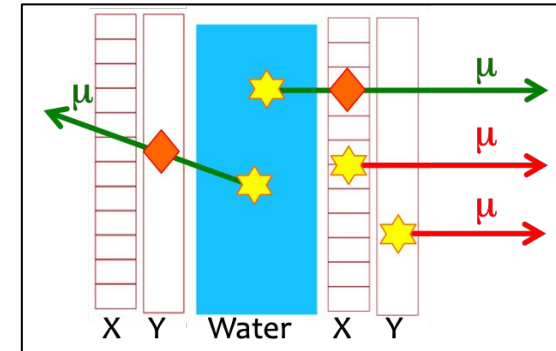
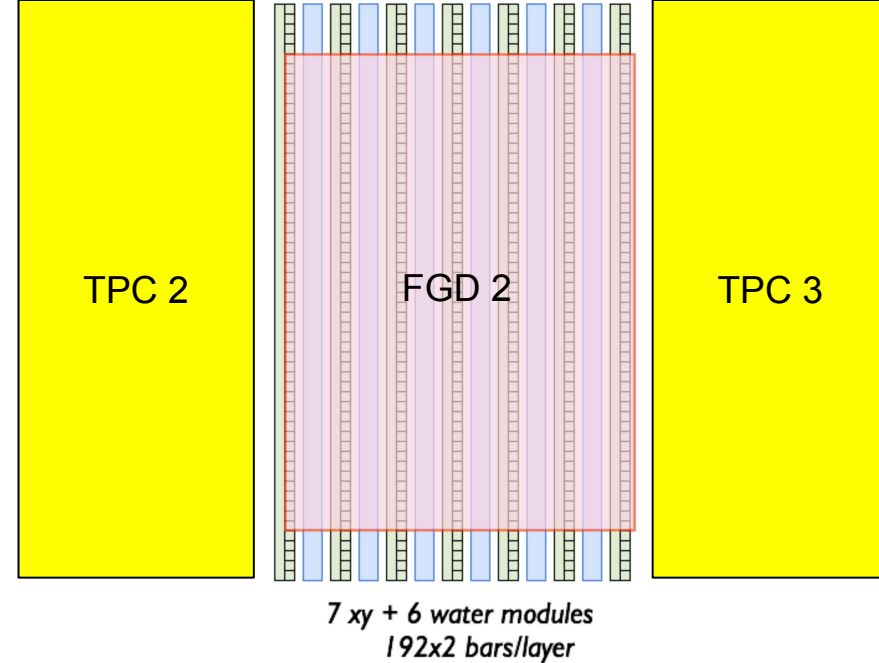


T2K Near Detector (ND280)



T2K Fine-Grained Detector (FGD2)

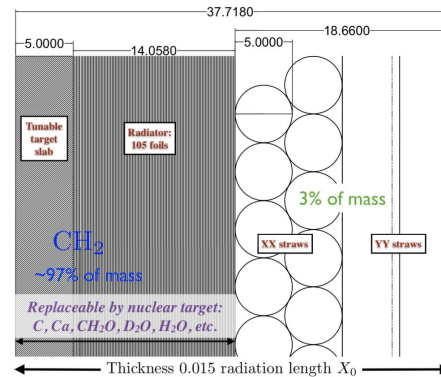
- T2K already employs water targets embedded within X & Y layers of scintillator bars
 - This reduced T2K's neutrino interaction uncertainties on water by $\sim 30\%$
- One of the most important detector uncertainties is disentangling events occurring within water to events occurring in adjacent scintillator layers
- The key difference using WbLS is the water layers themselves can be instrumented
 - Surrounding scintillator layers are no longer a strict requirement
 - Must ensure a sufficient light yield to record MIPs



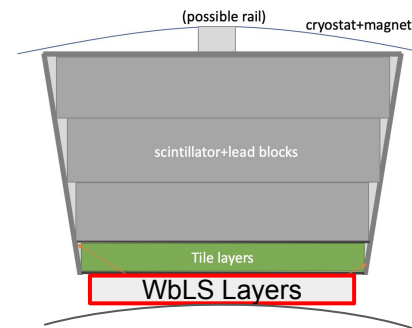
Near Detector Concepts

- Several concepts are possible with varying levels of complexity
- SAND already exists, so adding targets for studying WbLS nuclei is possible
 - On-axis only, but this can constrain extended xsec models
- ND-GAr is a primary goal for a DUNE Phase-2 ND
 - Adding WbLS targets in the upstream ECAL is possible
- A dedicated Theia ND
 - e.g. a WbLS Liquid-O
 - e.g. a WbLS NOvA ND

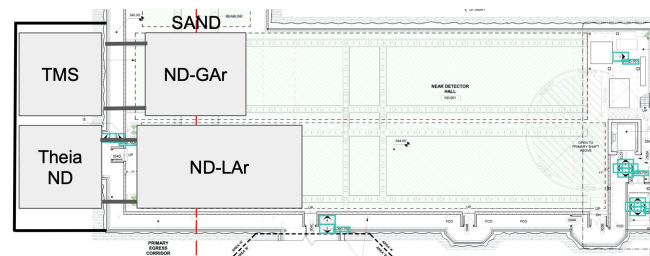
Additional nuclear targets in SAND



WbLS targets in the ND-GAr ECAL



Dedicated Theia ND

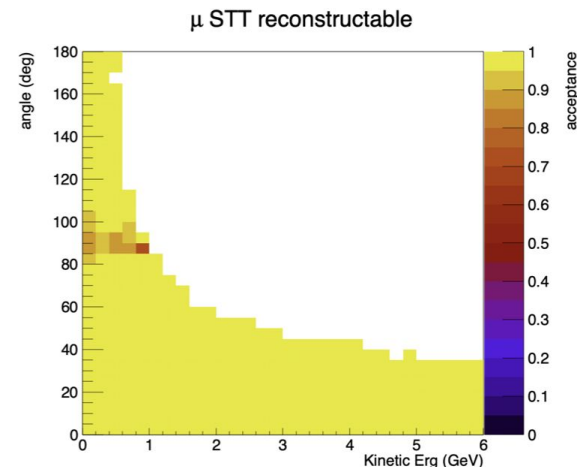
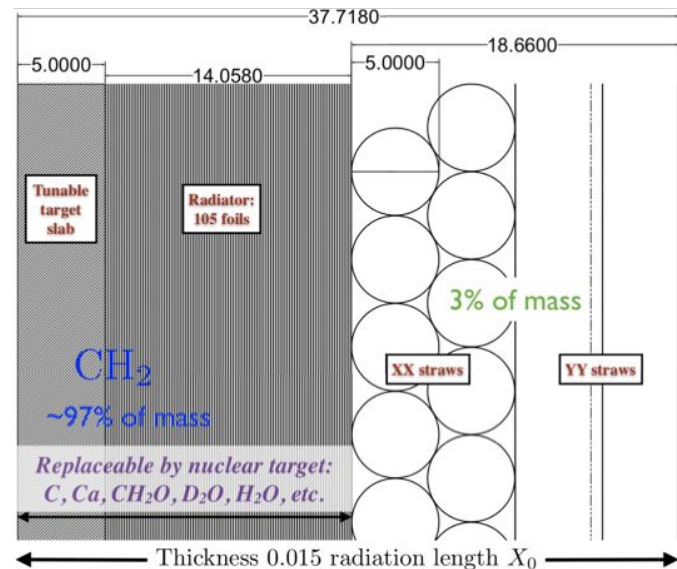


Least complex

Most complex

C/O/H Targets in SAND

- See previous talk at this workshop from R. Petti
- Identical layers of different target nuclei produce event samples that can be simultaneously fit to constrain differential nuclear effects
 - This approach was tried in T2K with some success (~30% reduction in neutrino cross section uncertainties)
 - SAND should do better, due to more precise tracking, better resolution, better acceptance, and much higher statistics

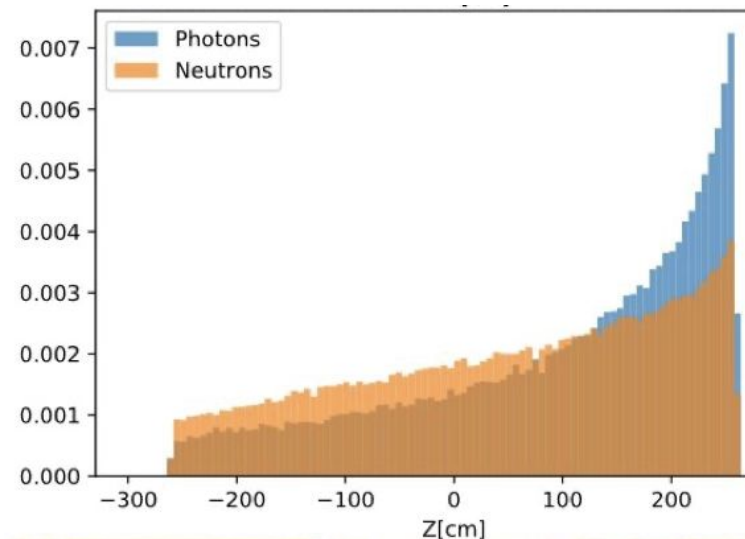
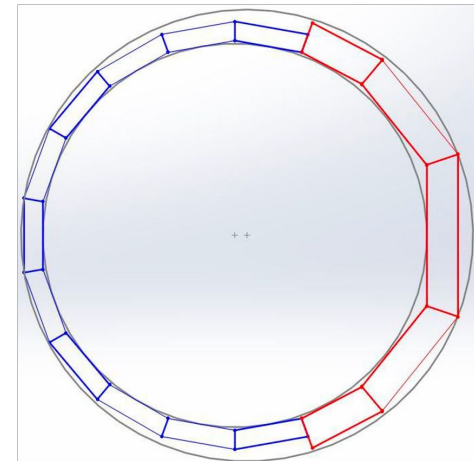


Target	CP optimized FHC (1.2MW, 2y)				CP optimized RHC (1.2MW, 2y)			
	ν_μ CC	$\bar{\nu}_\mu$ CC	ν_e CC	$\bar{\nu}_e$ CC	ν_μ CC	$\bar{\nu}_\mu$ CC	ν_e CC	$\bar{\nu}_e$ CC
<i>CH₂</i>	13,010,337	624,330	192,118	31,902	2,035,973	4,870,562	91,004	69,278
<i>H</i>	1,222,576	111,574	18,396	5,557	194,216	906,130	8,712	12,434
<i>C</i>	1,547,011	67,294	22,799	3,458	241,710	520,287	10,800	7,460
<i>Ar</i>	3,114,331	121,506	46,384	6,503	480,862	936,489	21,932	13,867
<i>Pb</i>	62,127,600	2,507,940	923,012	130,680	10,375,400	18,222,200	437,284	265,304

NOTE: 100 kt-MW-years in Phase I FD corresponds to about 2y FHC + 2y RHC with 1.2 MW beam

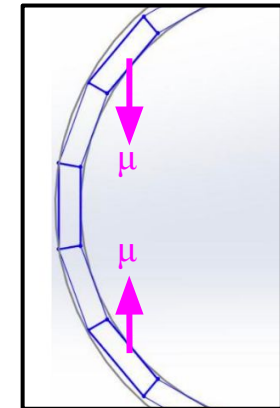
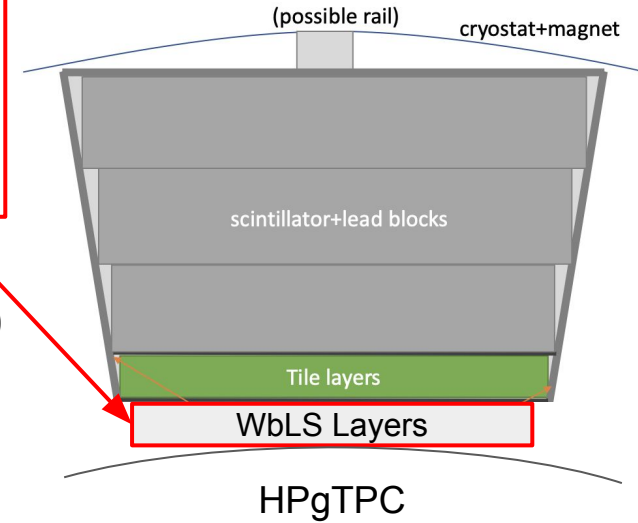
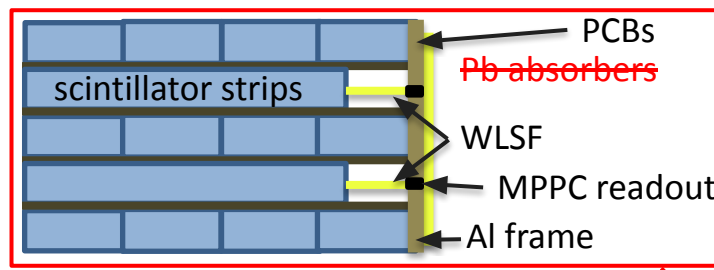
ND-GAr “Thin” Upstream ECAL

- The ND-GAr ECAL is most needed in the downstream direction
- The upstream portion has been redesigned to be “thin”
 - The thickness of the ECAL as a function of angle has not yet been optimized
- WbLS layers could be placed in the downstream portion of the upstream ECAL
 - <1 radiation length (i.e. preshower tracking layers)



WbLS Inside ND-GAr ECAL

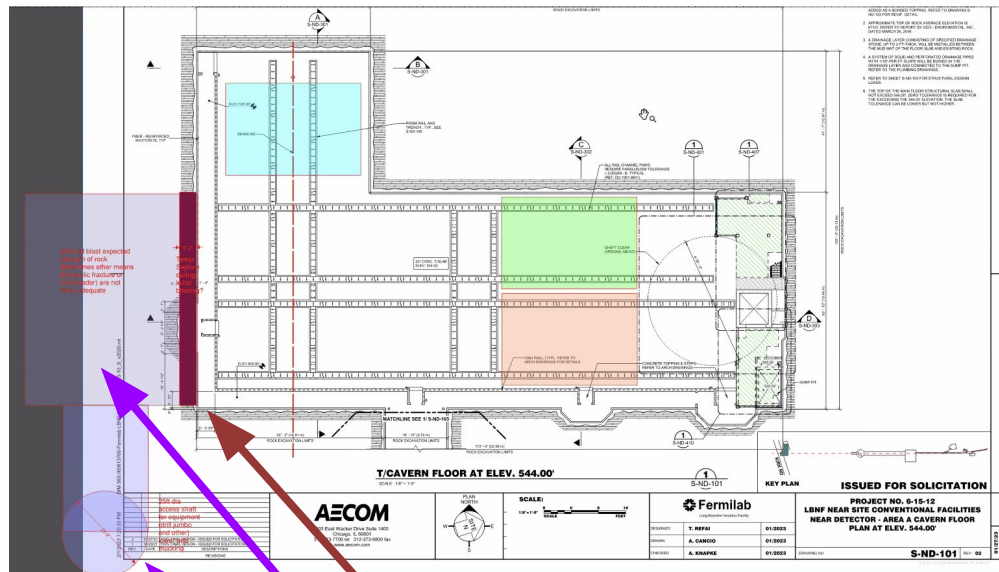
- WbLS layers would need to track X & Y positions
 - Optically segmented X & Y bars (NOvA ND) or 3D cubes (T2K sFGD)
 - Or perhaps a non-segmented LiquidO detector with X & Y fibers
- A few cm WbLS layers provides ~1 ton of target mass
 - A few tons of WbLS in a 2.4 MW beam would produce:
 - ~1M ν_μ -CC events per year on-axis (14 week run)
 - ~100k ν_μ -CC events per year 8m off-axis (2 week run)
 - ~10k ν_μ -CC events per year 28m off-axis (2 week run)
- Additional benefit: variation in detector configurations allows for sampling all of the muon angle phase space
 - The lack of muon acceptance near 90° was an important limitation of the T2K FGD+TPC configuration



Possible ND Hall Extensions

- The Fermilab engineers leading the development of the Phase 1 ND Hall (Tom Hamernik & Kennedy Hartsfield) took a preliminary look at technical feasibility of extending the ND Hall
 - How do we protect detectors that are already installed?
 - Is there enough space to deploy excavation equipment?
- Proposed solution is to create a small 3rd shaft and excavate the additional cavern space (last step: break barrier)
 - Initial look from Fermilab site rock experts revealed no show stoppers
 - Additional study on the impact of vibrations on detectors is needed
- The PRISM range can then be extended, enabling additional detectors

DUNE Near Detector Hall



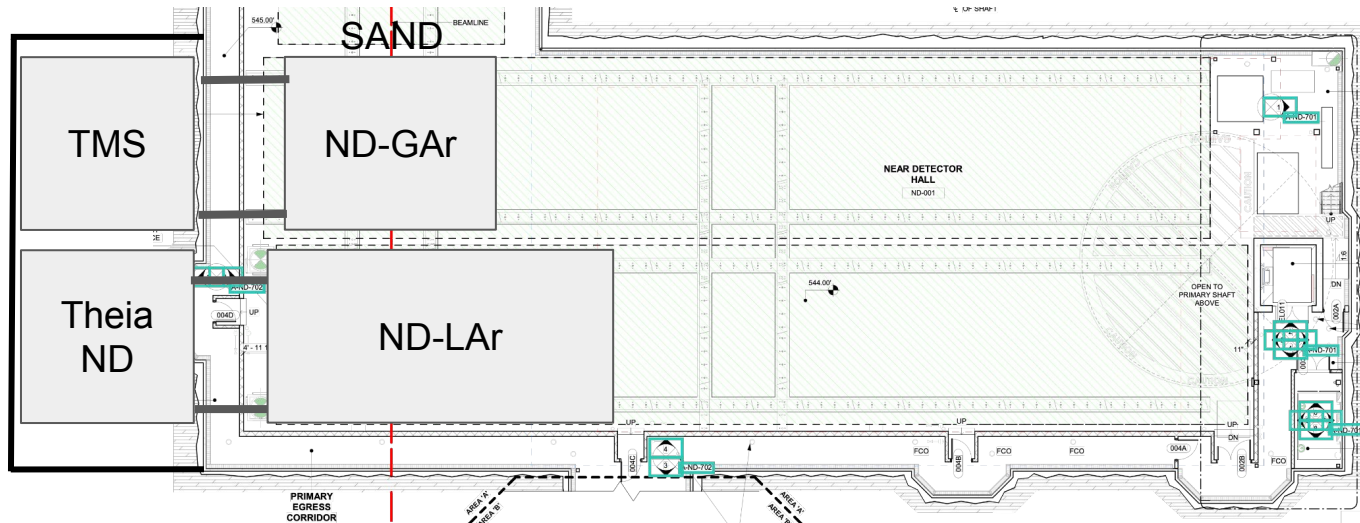
Rock septum (break last)

Potential Extension

Small 3rd Shaft

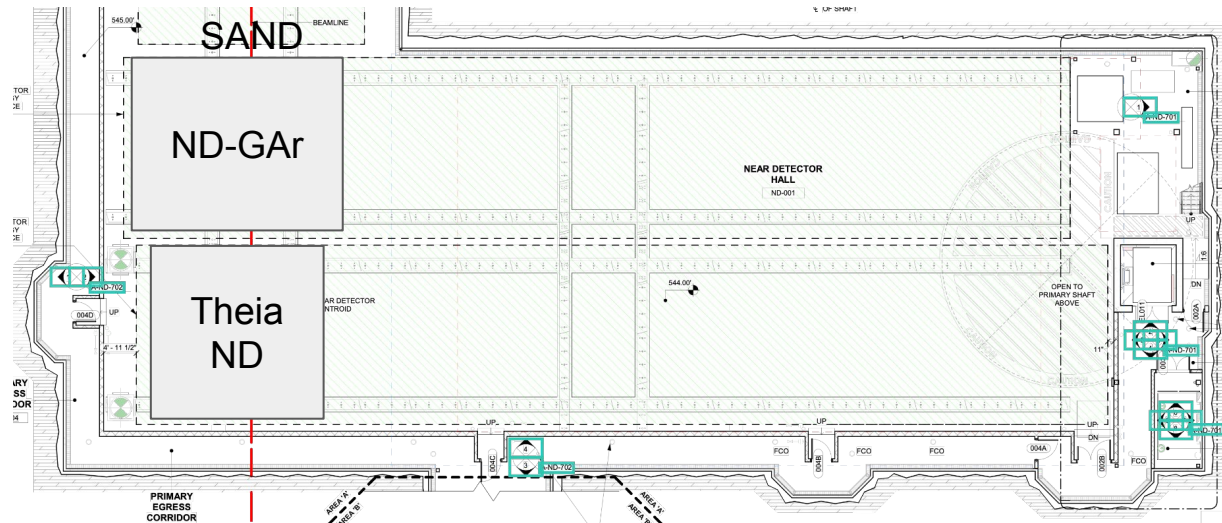
A Dedicated Theia Near Detector

- If the long dimension of the ND Hall can be slightly extended, it may be possible to install a dedicated Theia ND
- The disassembly of TMS is very time consuming and may delay the installation of ND-GAr
 - Instead, TMS could be reused as the muon catcher for a Theia ND
- All 4 detectors could then be moved to different locations along the PRISM rails



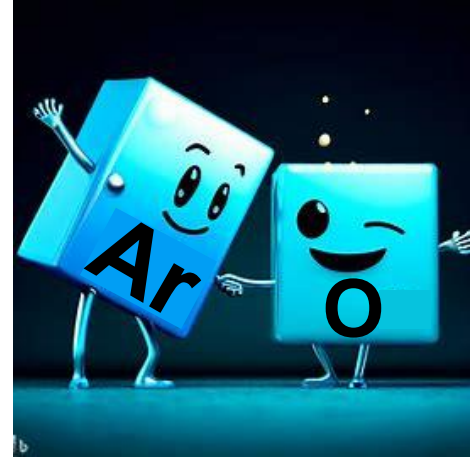
A Dedicated Theia Near Detector II

- After many years of Phase 1 running and producing several rounds of DUNE oscillation results, we will likely collect sufficient statistics with ND-LAr for the full duration of DUNE
 - See my DUNE-PRISM talk from earlier today for caveats regarding beam stability
 - It is possible that ND-LAr may not be needed for DUNE Phase 2
- In this scenario, a Theia ND could be placed in front of ND-GAr
 - Any additional measurements required on Ar could potentially be made by ND-GAr



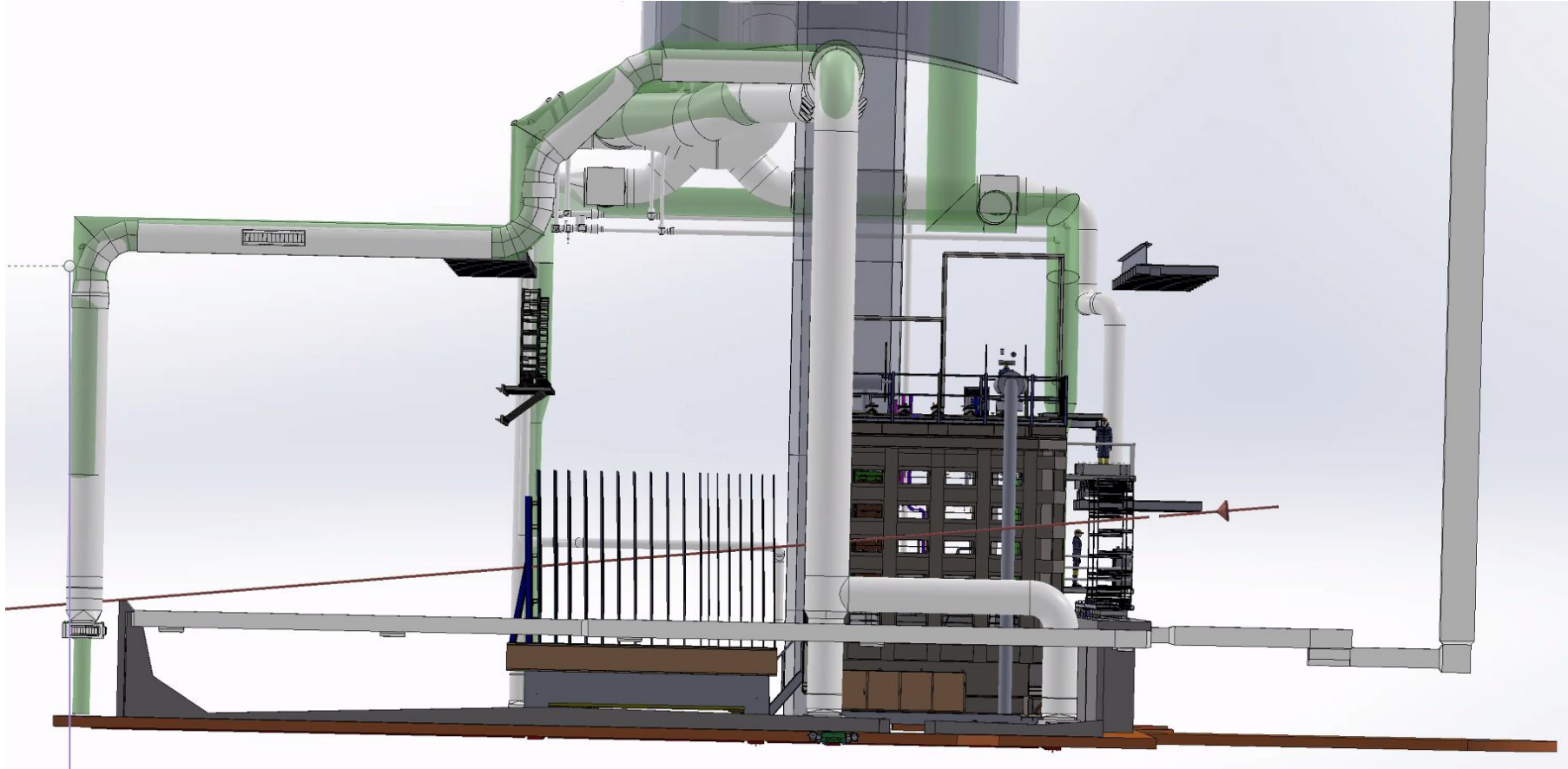
Summary

- Several near detector options exist for a non-Ar 4th FD module
 - and they are all compatible with, and closely coupled to, ND-GAr
- Least complex: nuclear targets in SAND
 - Pros: straightforward to implement C/O/H targets and/or water targets
 - Cons: no additional handle on Erec vs Etrue from off-axis measurements
- More complex: WbLS layers in the ND-GAr ECAL
 - Pros: provides off-axis data and excellent tracking
 - Cons: must balance with ECAL performance; lower event rates must be studied
- Most complex: A new, dedicated Theia ND (e.g. Liquid-O or NOvA ND)
 - Pros: can be designed for high statistics measurements off-axis; possible reuse of TMS when ND-GAr is installed
 - Cons: Requires additional hall excavation, or removal of ND-LAr after multiple years of Phase 1 running
- If the LBL sensitivity can be demonstrated, a Theia far detector would broaden DUNE's physics program (after many years of Phase 1 running) to include a variety of interesting low-E physics phenomena



Backup

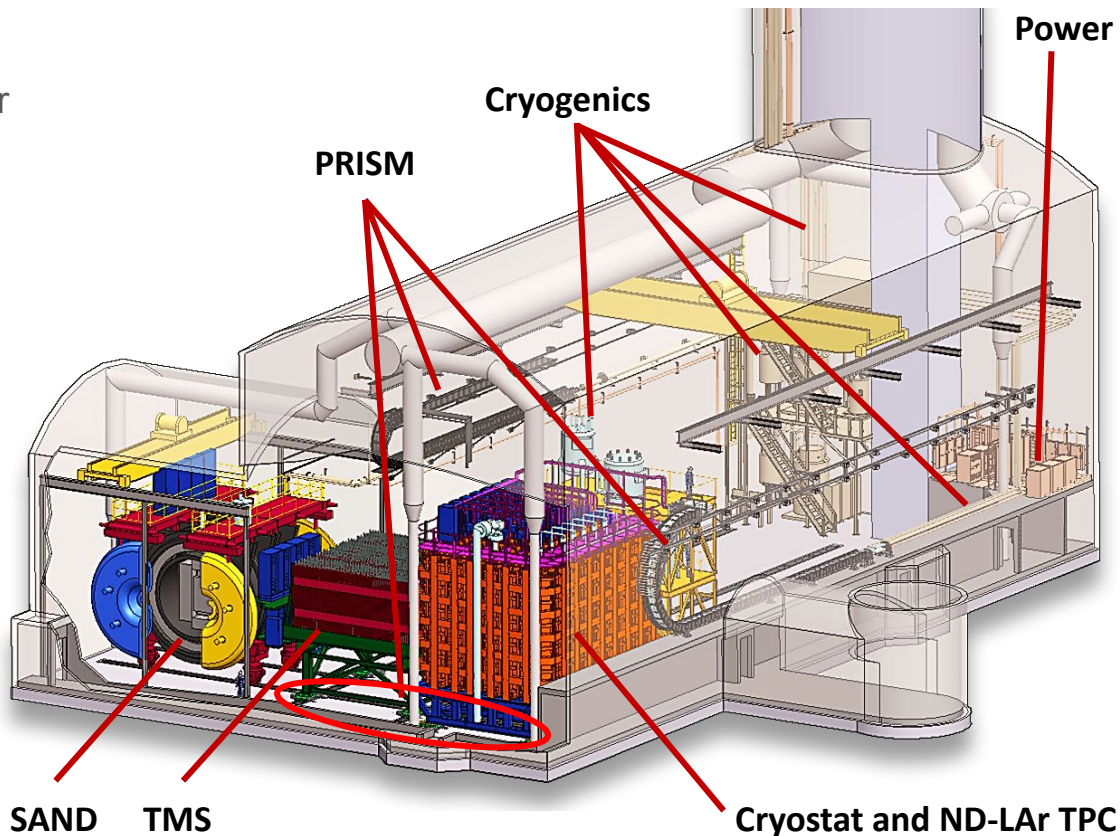
TMS View From Cavern Endwall



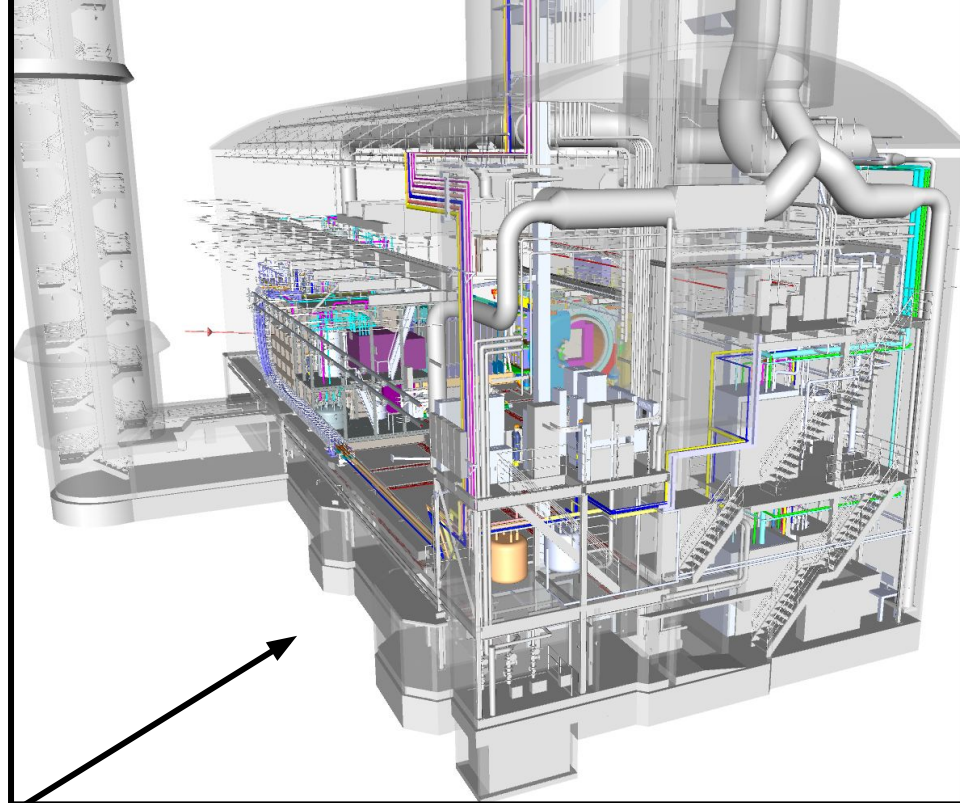
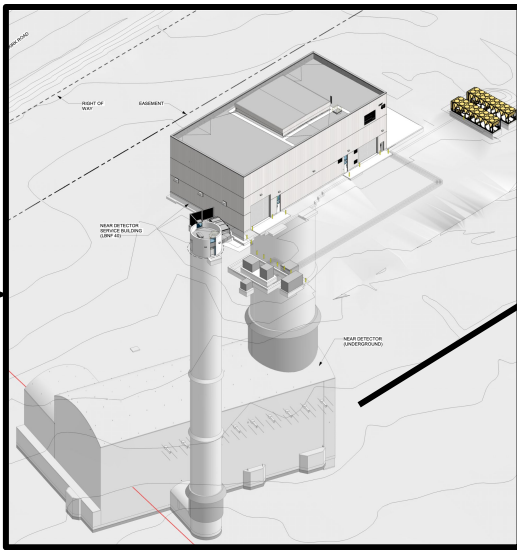
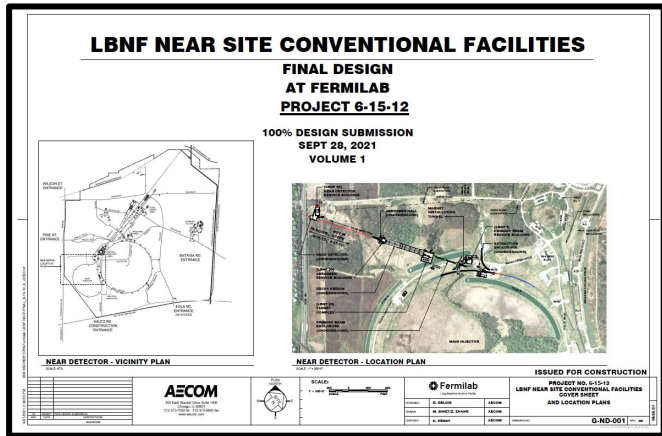
Near Detector Considerations

- Near detectors are an essential element of any LBL analysis
 - Measurements on the same nuclear target(s) as the far detector are required
- DUNE ND is currently designed around Ar
 - ND-LAr TPC: ν -Target w/ similar technology to LAr far detectors
 - TMS: Spectrometer for muons escaping ND-LAr
 - PRISM: ND-LAr + TMS move off-axis to sample a variety of E_ν
 - SAND: Beam monitor

DUNE Near Detector Hall



DUNE Near Detector Hall Design

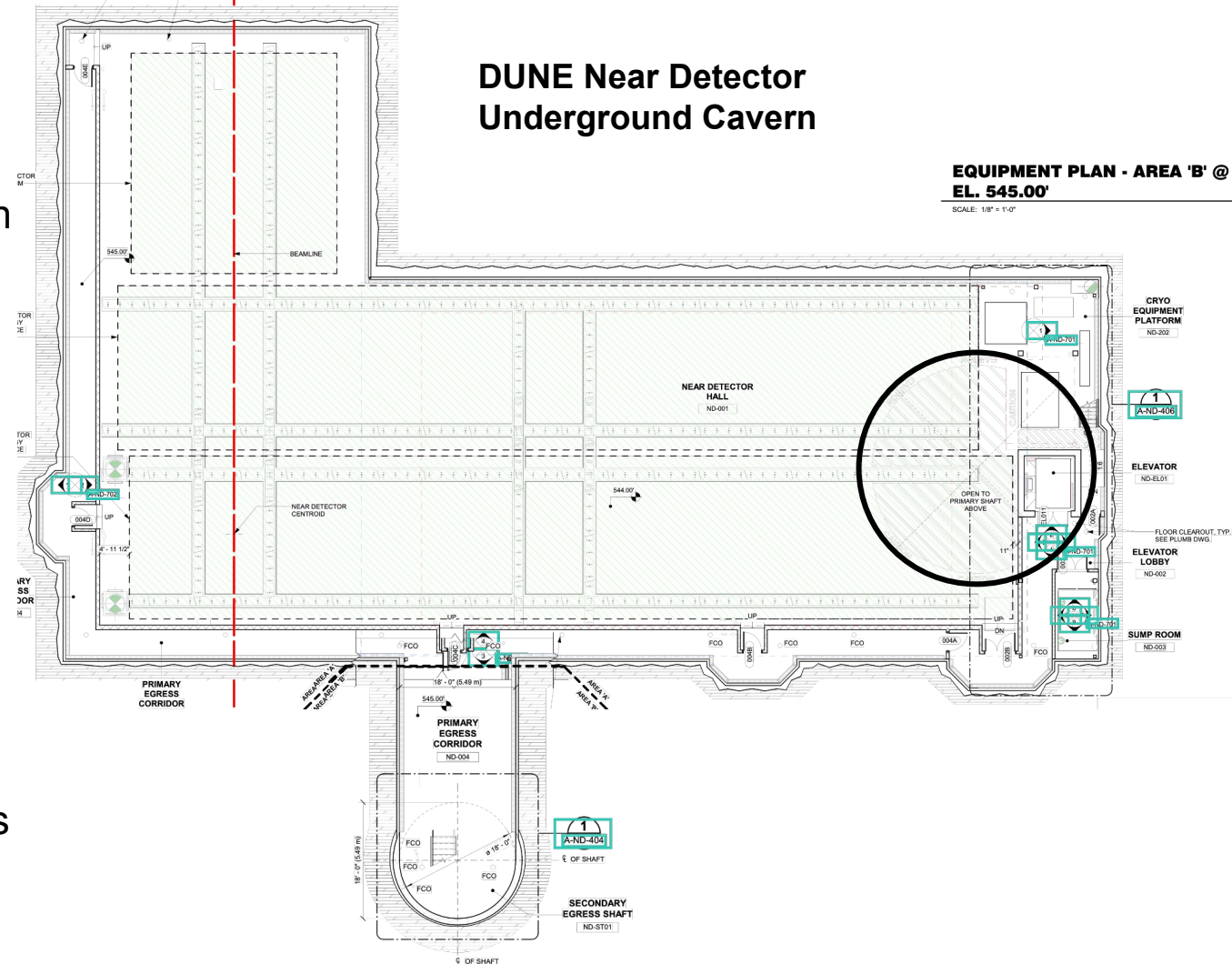


The near detector hall for DUNE Phase 1 is at “100% final design” (i.e. changes to the hall at this point would be very difficult)

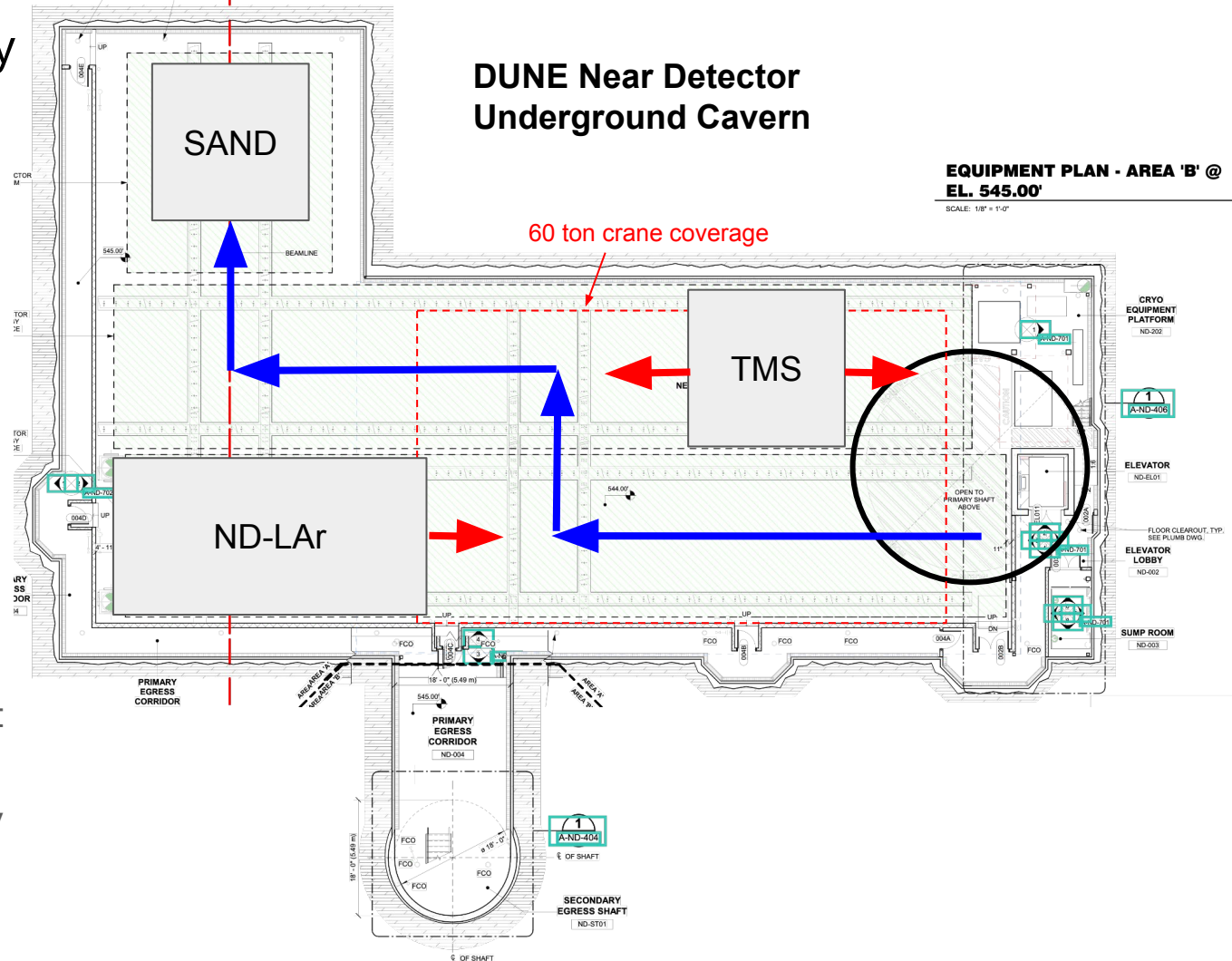
Cavern Layout

- **Red line:** beam axis
- Off-axis travel direction is left-to-right
- 2 pairs of rails for ND-LAr and TMS
- 2 sets of “cross rails” (along the beam direction) for SAND
 - SAND can move between the TMS and ND-LAr rails
- The main shaft, elevator, and equipment mezzanines are on the right side

DUNE Near Detector Underground Cavern

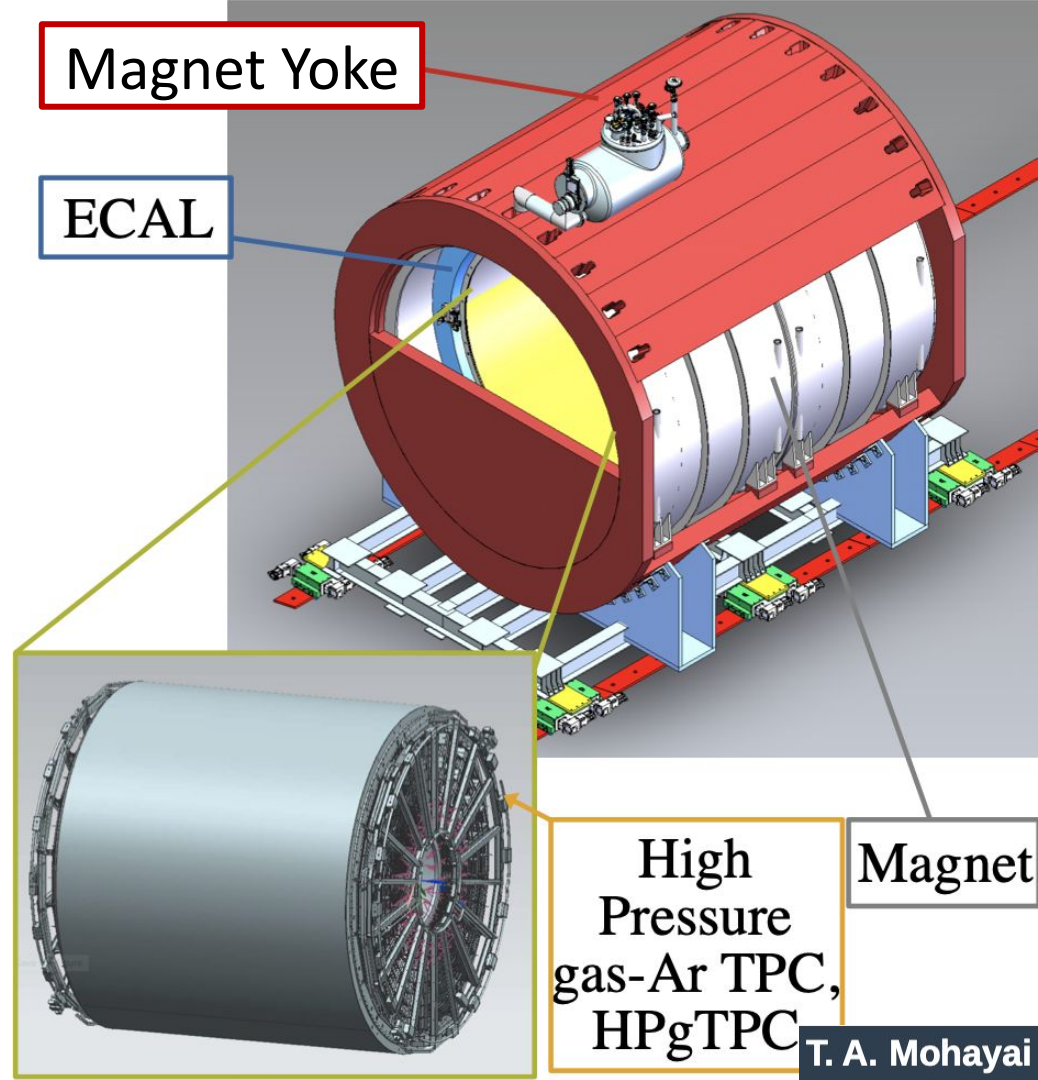


- The rail structure is designed to allow SAND to be installed at almost any time
- TMS and ND-LAr can move (via the PRISM system)
 - ND-LAr can temporarily move under the 60 ton crane coverage
 - TMS can temporarily move under the shaft
- Significant flexibility to accommodate a variety of installation scenarios



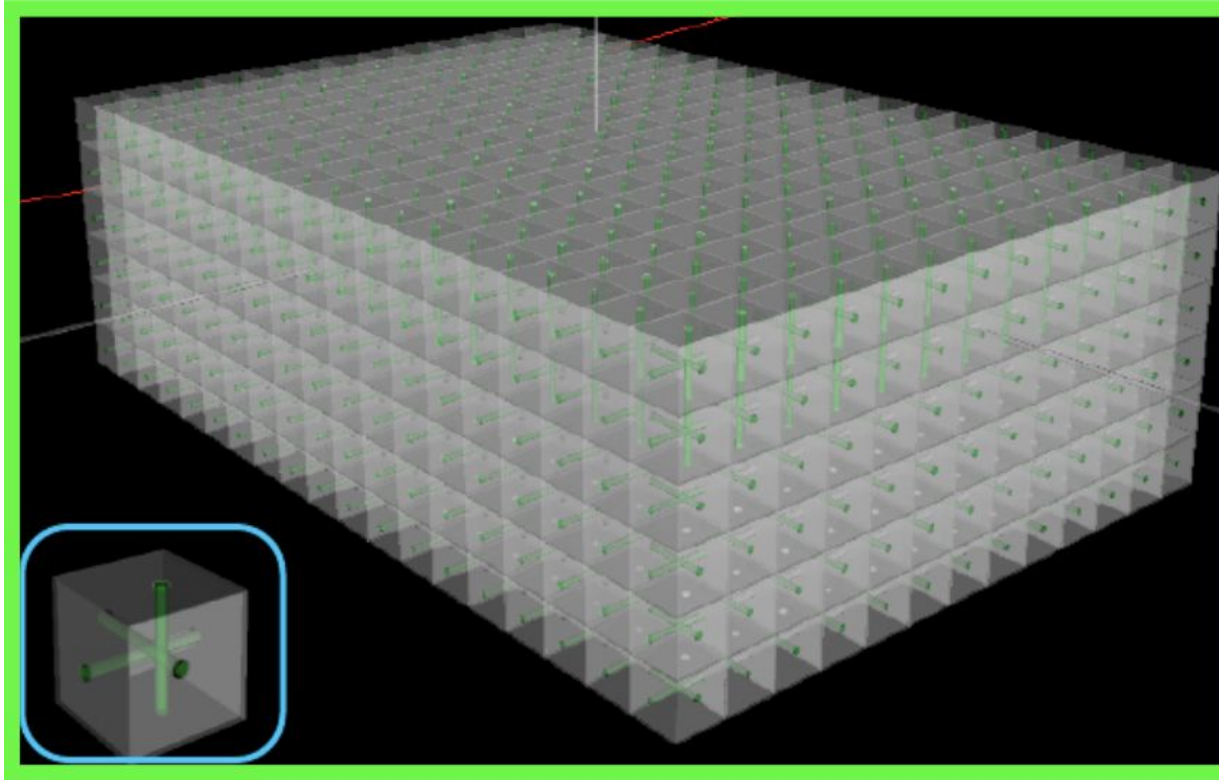
ND-GAr

- DUNE Phase 2 includes plans for an upgraded near detector
- The main option discussed so far is a high-pressure Ar gas TPC in place of TMS
 - Lowers the momentum threshold for detecting particles escaping the Ar nucleus
 - Cleaner measurements of multi-particle final states (e.g. reduces π^+ scattering, γ -conversions, etc.)
- This detector still must function as a muon catcher for ND-LAr
 - Goal is to minimize dead material between ND-LAr and ND-GAr



Super-FGD -> WbLS cubes?

- The Super-FGD currently being constructed for T2K consists of 1 cm^3 scintillator cubes
- Can we incorporate WbLS cubes into a Super-FGD structure for DUNE?
 - For either a dedicated detector or for embedding into ND-GAr?



DUNE FD-TDR Cross Section Model

- Uncertainties included for:
 - Exclusive interactions (QE, Res, SIS/DIS)
 - Final state interactions (FSI)
 - Nuclear effects (RPA, 2p2h)
 - Flavor ratios (ν_e /anti- ν_e)
- A similar set of uncertainties will be needed for C/O/H
 - Fortunately, these nuclei have been studied more extensively
 - Specific expertise in ν -N modeling and GENIE is needed (& communication with DUNE DIRT2/NIUWG)

GENIE Xsec Parameters

Description	1 σ
Quasielastic	
M_A^{QE} , Axial mass for CCQE	$+0.25_{-0.15}$ GeV
QE FF, CCQE vector form factor shape	N/A
p_F Fermi surface momentum for Pauli blocking	$\pm 30\%$
Low W	
M_A^{RES} , Axial mass for CC resonance	± 0.05 GeV
M_V^{RES} Vector mass for CC resonance	$\pm 10\%$
Δ -decay ang., θ_π from Δ decay (isotropic \rightarrow R-S)	N/A
High W (BY model)	
A_{HT} , higher-twist in scaling variable ξ_w	$\pm 25\%$
B_{HT} , higher-twist in scaling variable ξ_w	$\pm 25\%$
C_{V1u} , valence GRV98 PDF correction	$\pm 30\%$
C_{V2u} , valence GRV98 PDF correction	$\pm 40\%$
Other neutral current	
M_A^{NCRES} , Axial mass for NC resonance	$\pm 10\%$
M_V^{NCRES} , Vector mass for NC resonance	$\pm 5\%$

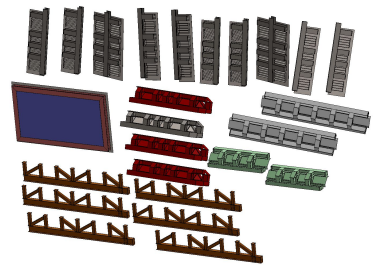
GENIE FSI Parameters

Description	1 σ
N. CEX, Nucleon charge exchange probability	$\pm 50\%$
N. EL, Nucleon elastic reaction probability	$\pm 30\%$
N. INEL, Nucleon inelastic reaction probability	$\pm 40\%$
N. ABS, Nucleon absorption probability	$\pm 20\%$
N. PROD, Nucleon π -production probability	$\pm 20\%$
π CEX, π charge exchange probability	$\pm 50\%$
π EL, π elastic reaction probability	$\pm 10\%$
π INEL, π inelastic reaction probability	$\pm 40\%$
π ABS, π absorption probability	$\pm 20\%$
π PROD, π π -production probability	$\pm 20\%$

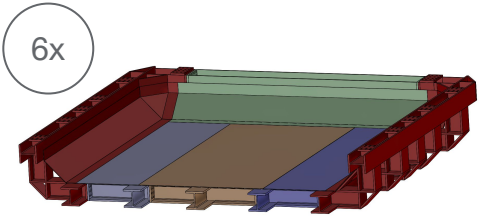
Additional Xsec Parameters

Uncertainty	Mode
BeRPA $[A, B, D]$	$1p1h/\text{QE}$
ArC2p2h $[\nu, \bar{\nu}]$	$2p2h$
E_{2p2h} $[A, B]$ $[\nu, \bar{\nu}]$	$2p2h$
NR $[\nu, \bar{\nu}]$ $[\text{CC}, \text{NC}]$ $[\text{n}, \text{p}]$ $[1\pi, 2\pi, 3\pi]$	Non-res. pion
ν_e PS	$\nu_e, \bar{\nu}_e$ inclusive
$\nu_e/\bar{\nu}_e$ norm	$\nu_e, \bar{\nu}_e$ inclusive
NC norm.	NC

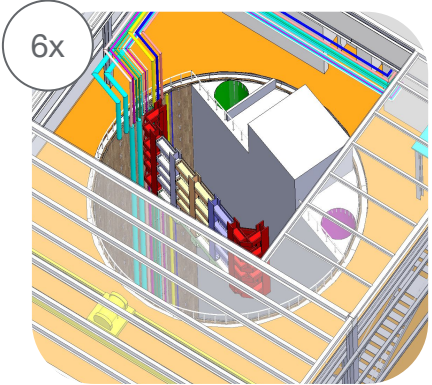
Cryostat Assembly Process Overview



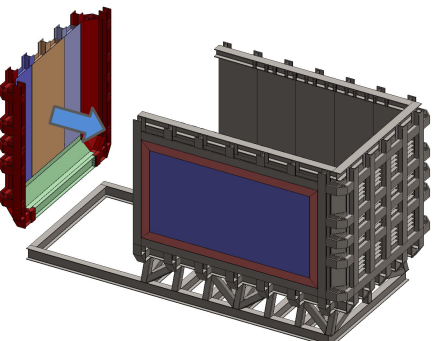
1. Receive 160 metric tons of warm structure parts in 40 pieces



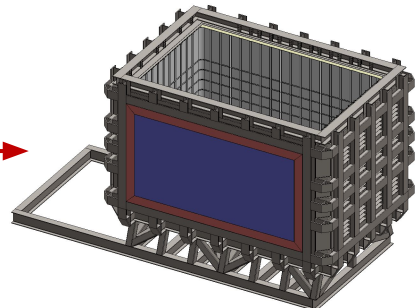
2. Align and bolt pieces into subassemblies



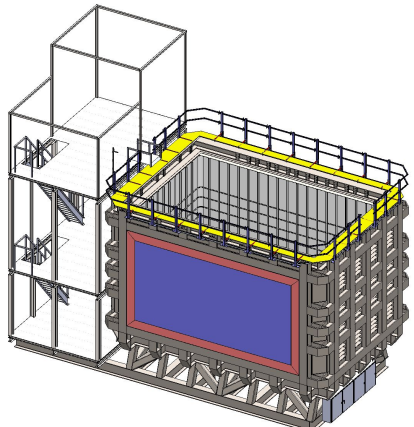
3. Lower 15-50 metric ton subassemblies into cavern



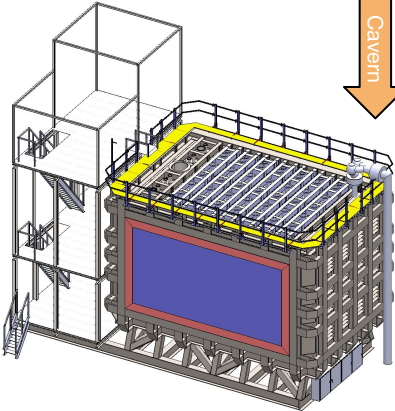
4. Align, bolt and weld subassemblies to form warm structure



5. Install cold membrane



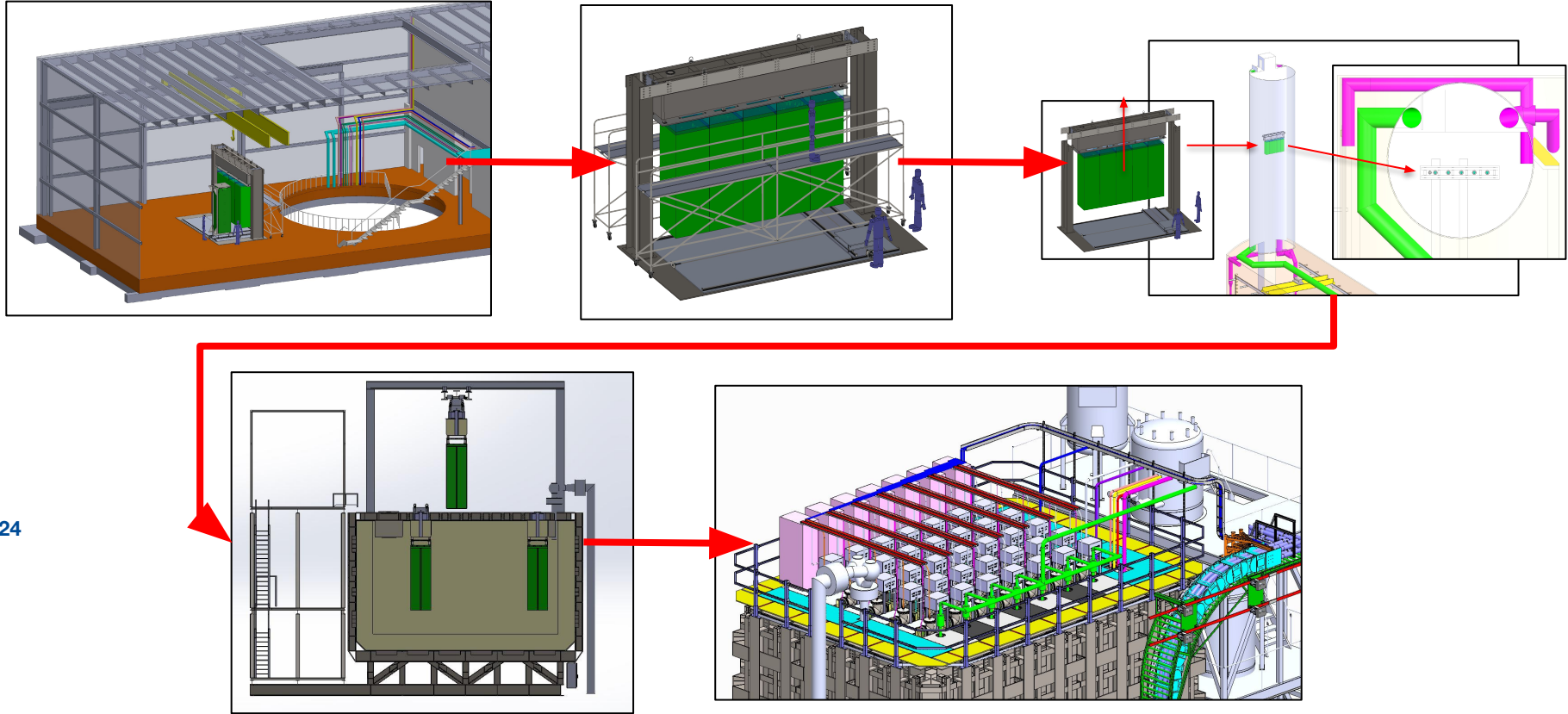
6. Install mezzanines



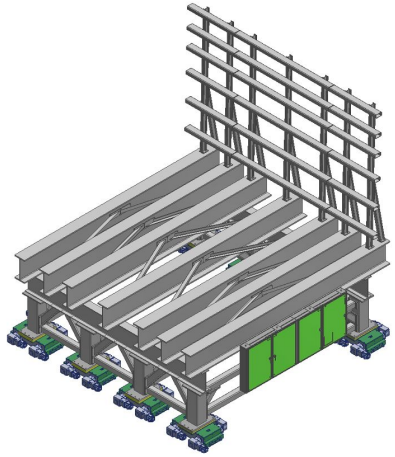
7. Install TPC rows and lid sections



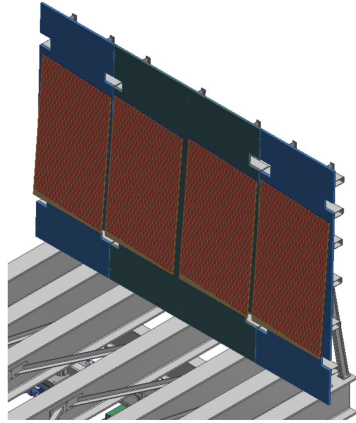
ND-LAr Assembly Process Overview



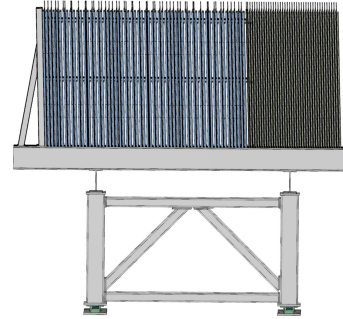
TMS Components and Assembly Overview



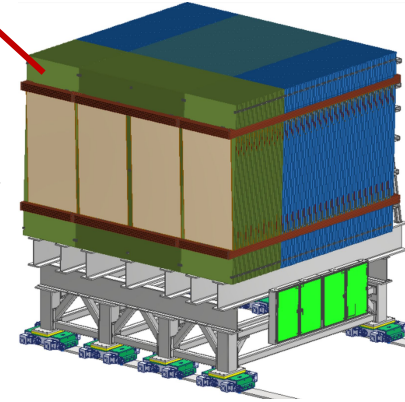
Iron Throne



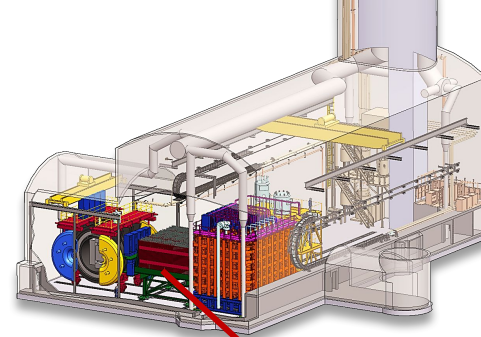
1st TMS layer

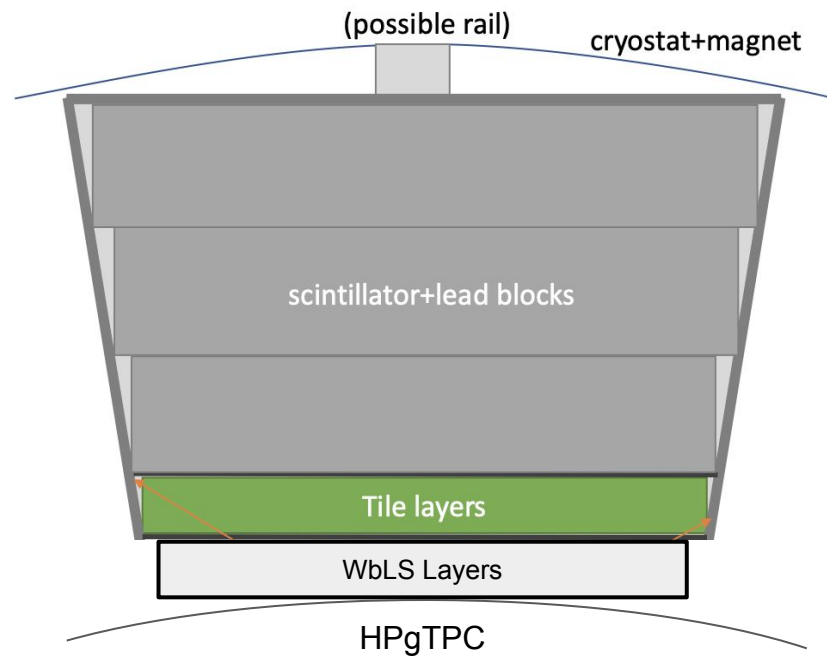


100 TMS layers

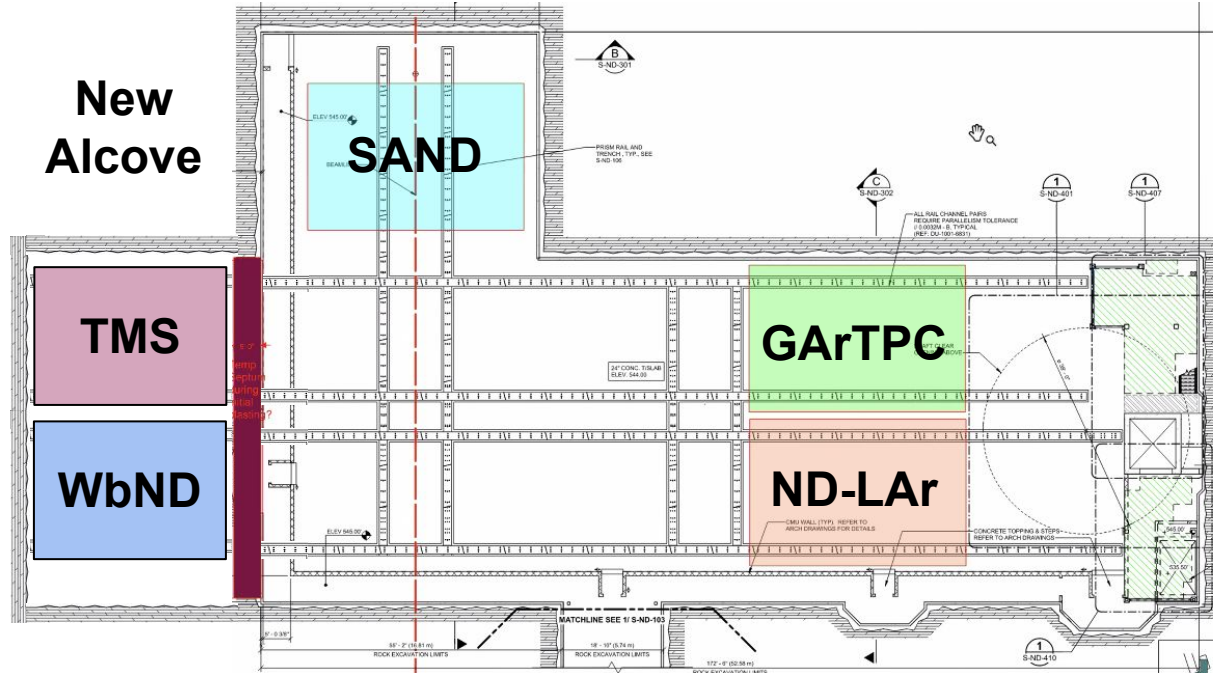


Magnet Coils

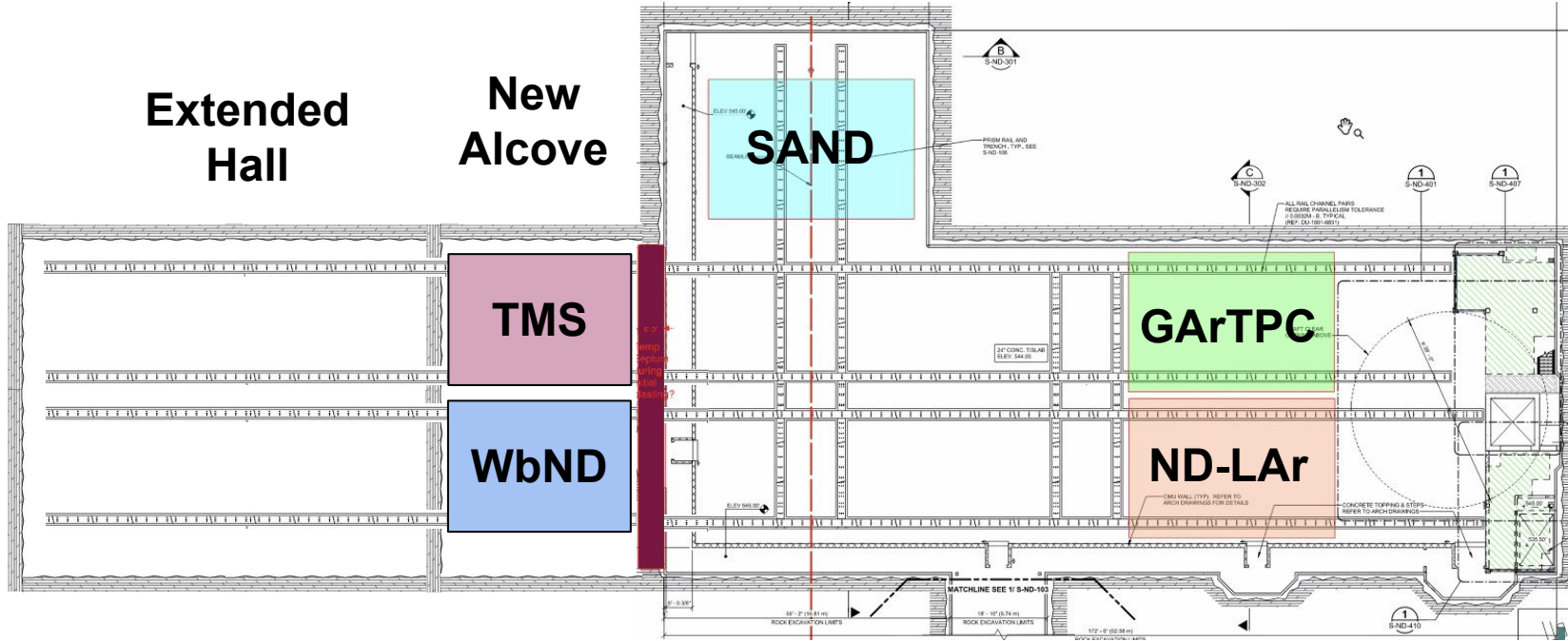




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