

An Assessment of Nuclear Data for Fusion and the Need for New Experiments

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U.S. DEPARTMENT OF
ENERGY

Office of
Science

Fission and Fusion from a Nuclear Scientist/Data Person's Perspective

Fission

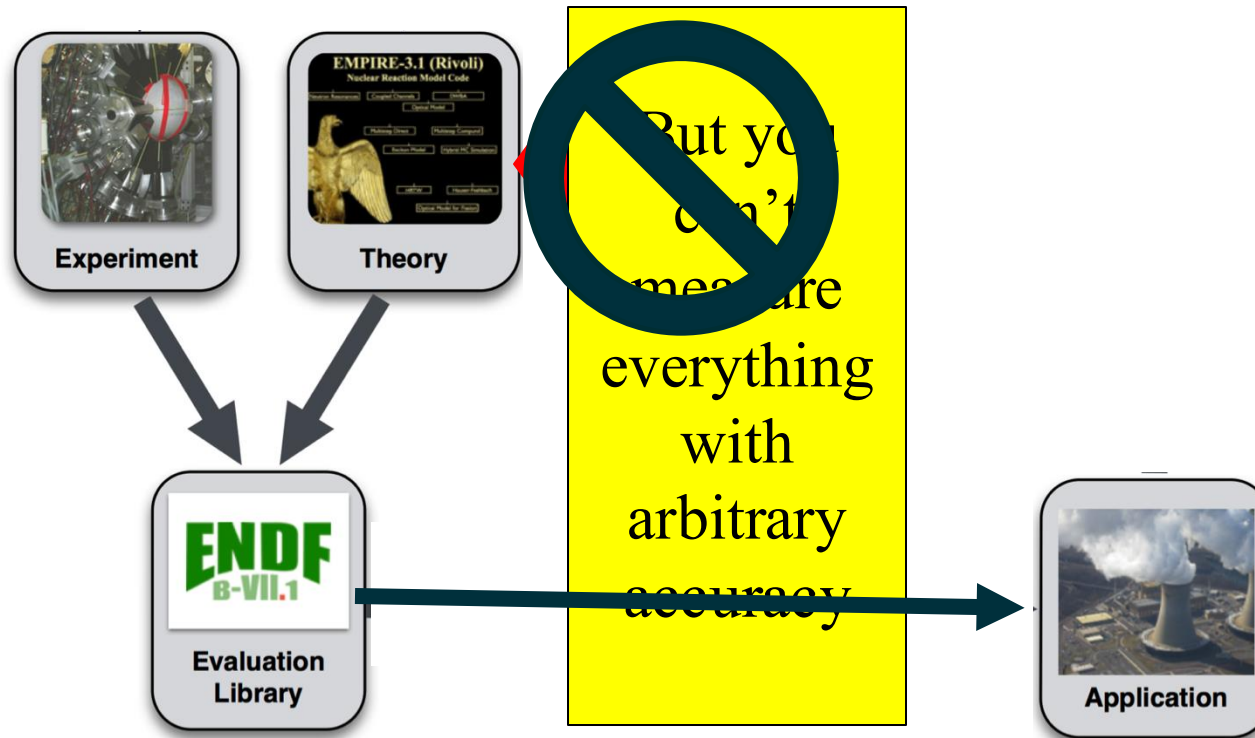
1. 85% of the energy is in the fission fragments, which can be stopped in \approx a few μm via *electronic stopping*.
2. Fuel needs to be dug up from the earth.
3. Reactors *MUST lose most of the* **neutrons** *made per fission*.
4. Generates some long-lived fission fragments *and actinides with proliferation concerns (Pu)*

Fusion

1. A 14 MeV neutron contains 80% of the energy and needs 10's of cm to be stopped *via nuclear scattering*
2. $\frac{1}{2}$ of the fuel needs to be made
3. Reactor *MUST use* **neutrons** *to create 1.1-12 tritons per fusion.*
4. Generates *some long-lived activation products and material with proliferation concerns (T)*

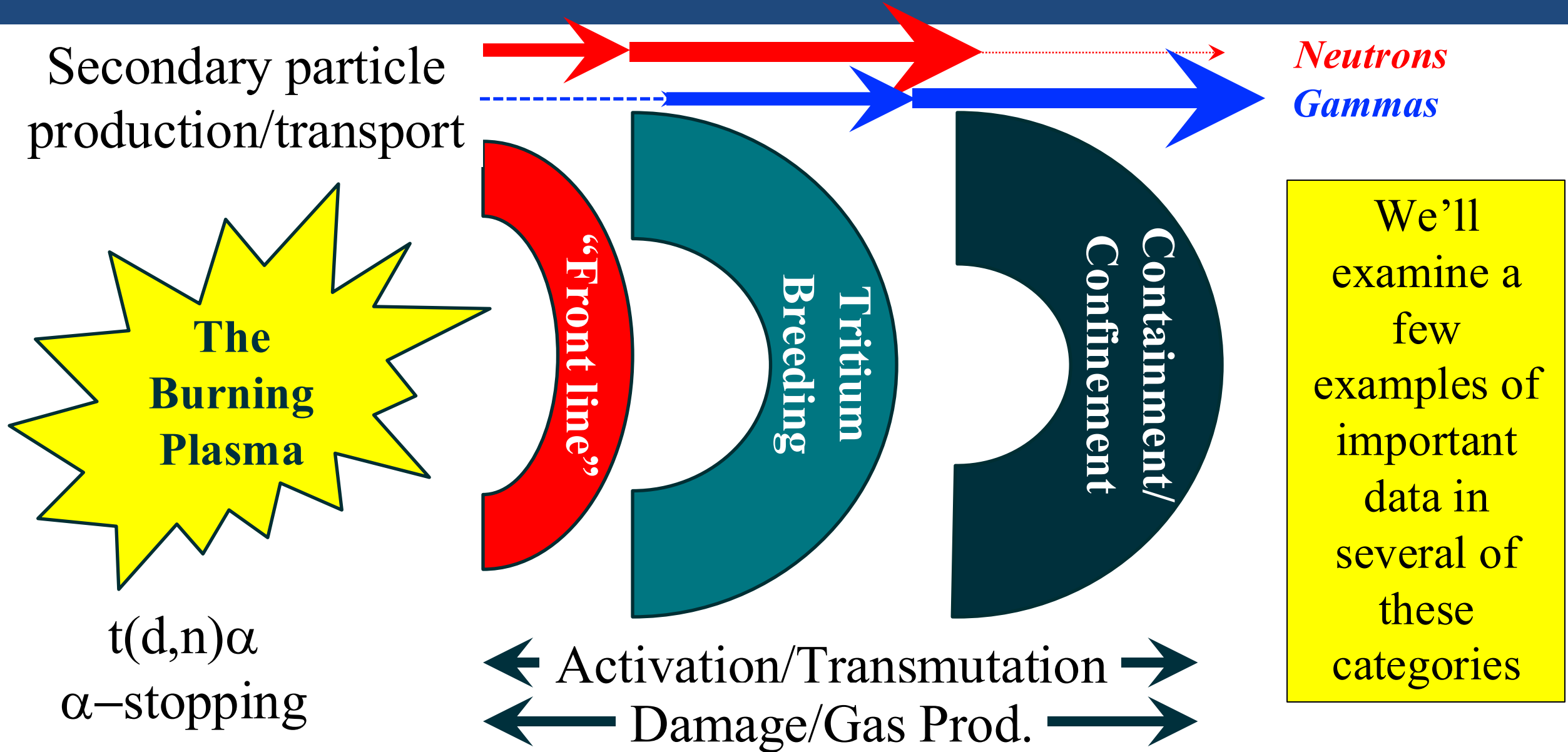
Neutron transport/capture are important to both

Reaction evaluation in a nutshell (Thanks to D. Brown!)



You shouldn't use integral benchmarks alone to gain confidence over your predictive capabilities if there is a lack of differential data to guide evaluation

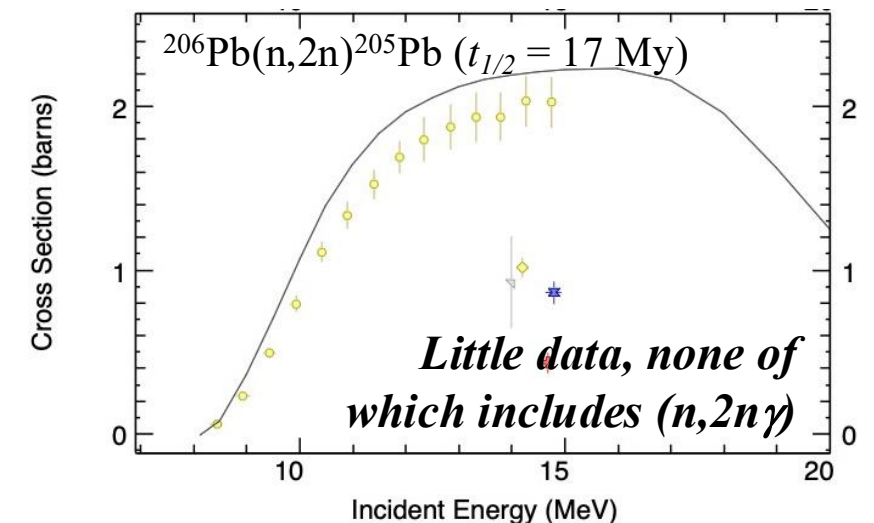
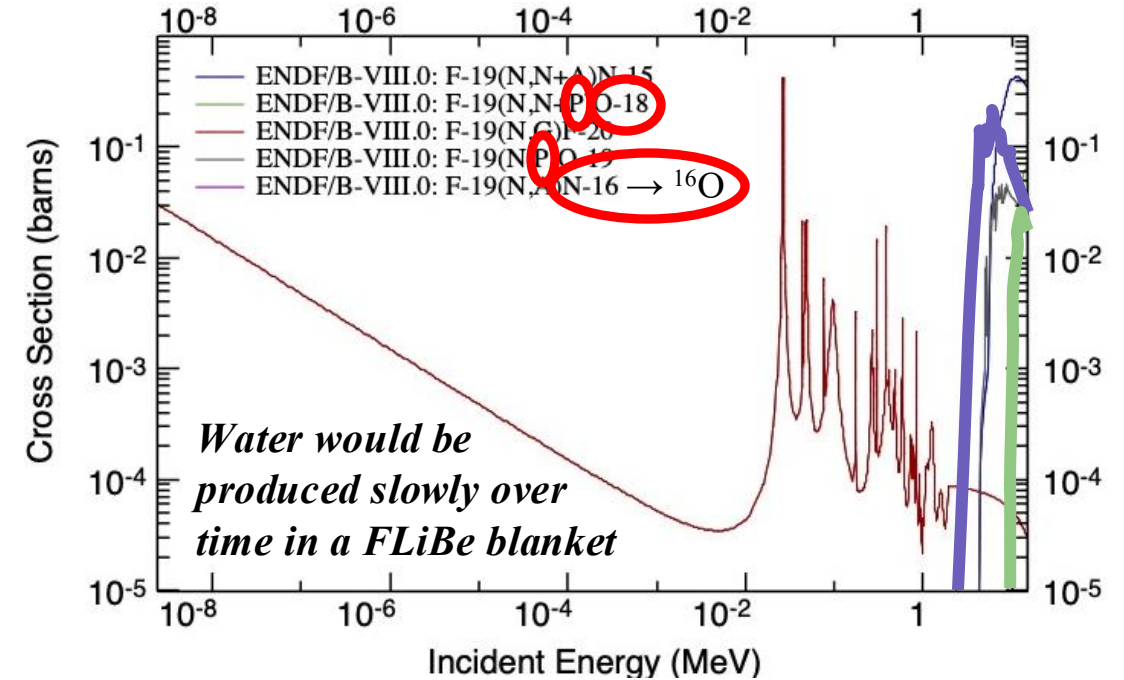
Where does Nuclear Data Play a role in a fusion energy system?



Transmutation/Activation & Gamma production is important through the fusion energy system

- Important Low-Z Nuclei: F, Li, Be, B, O
 - Channels: (n, γ), (n,p), (n,pn), (n, α), (n, α n), (n,2n)
 - BUT: Most don't lead to long-lived radioactive nuclei
 - ***Chemical transformation could pose operational risks***
- Important High-Z Nuclei: Pb, W, Ta, Bi
 - Channels: (n, γ), (n,2n), (n,3n)

Each High-Z (n,2n) reaction results in 1-3 γ -rays with a $\sum E_\gamma = 3-8$ MeV

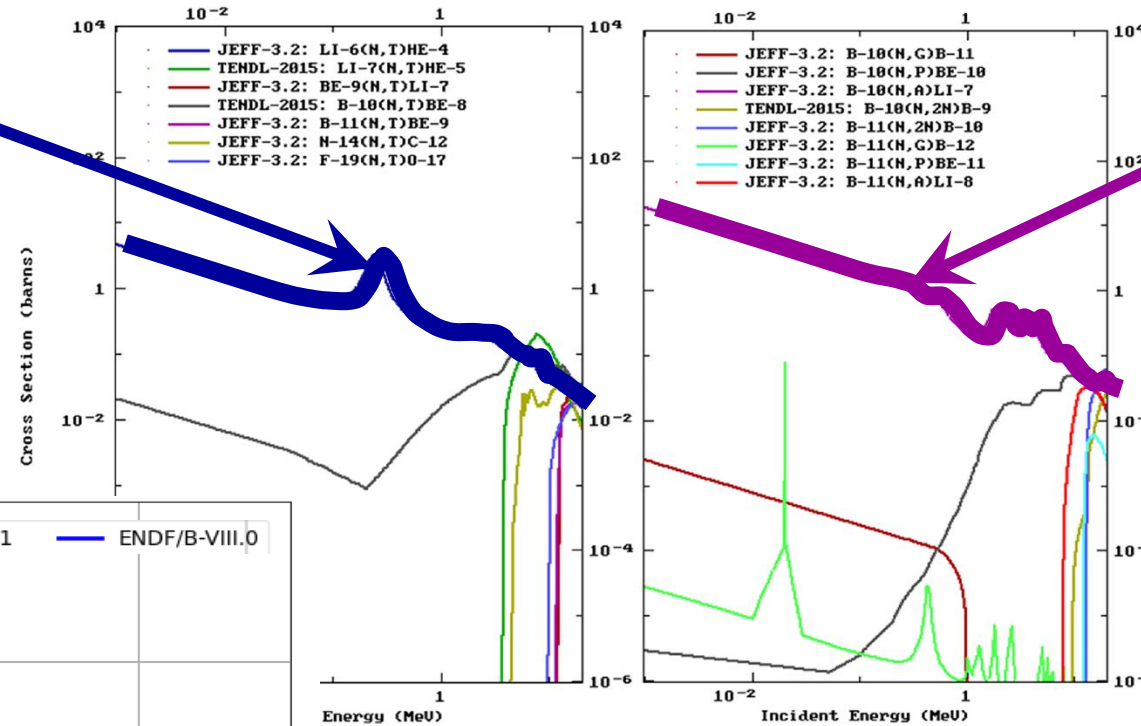


Tritium Breeding has focused on ${}^6\text{Li}(n,\alpha)\text{t}$ but ${}^{10}\text{B}$ might be of interest too

First principles review of options for tritium breeder and neutron multiplier materials for breeding blankets in fusion reactors

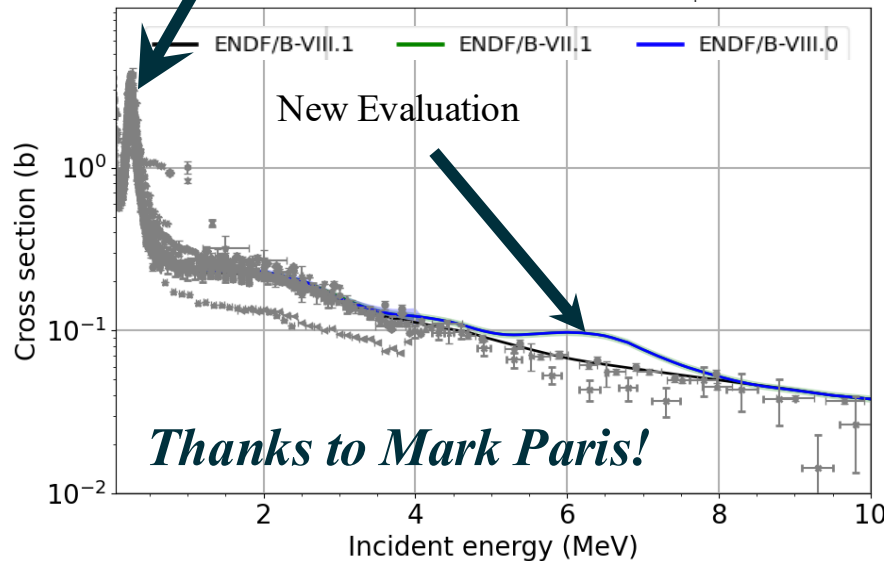
F.A. Hernandez*, P. Pereslavytsev

This peak in ${}^6\text{Li}(n,\alpha)\text{t}$ makes “fusion-eers” happy 😊

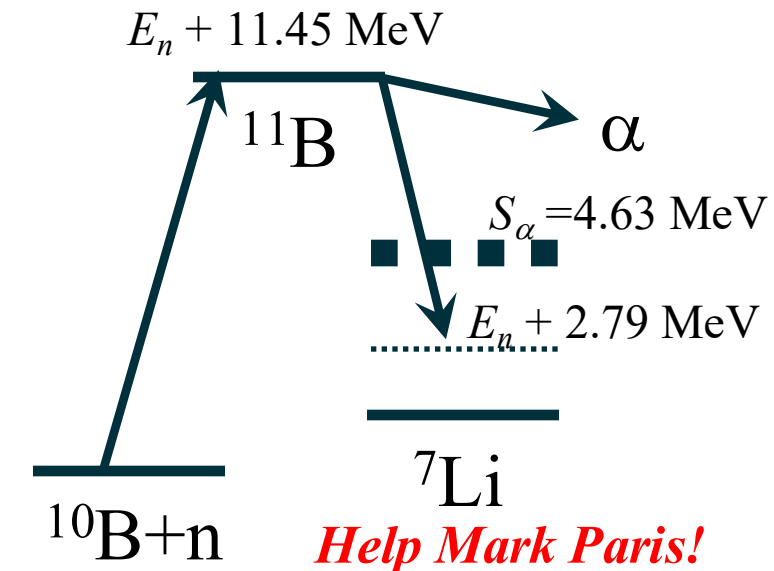


Is ${}^{11}\text{B}(n,\alpha){}^7\text{Li}$ REALLY that big?

Or does this curve have contributions from ${}^{11}\text{B}(n,\alpha){}^7\text{Li}^* \rightarrow \text{t} + \alpha$?



An activation experiment at 14 MeV and a coincident LENZ-style experiment might help tritium production



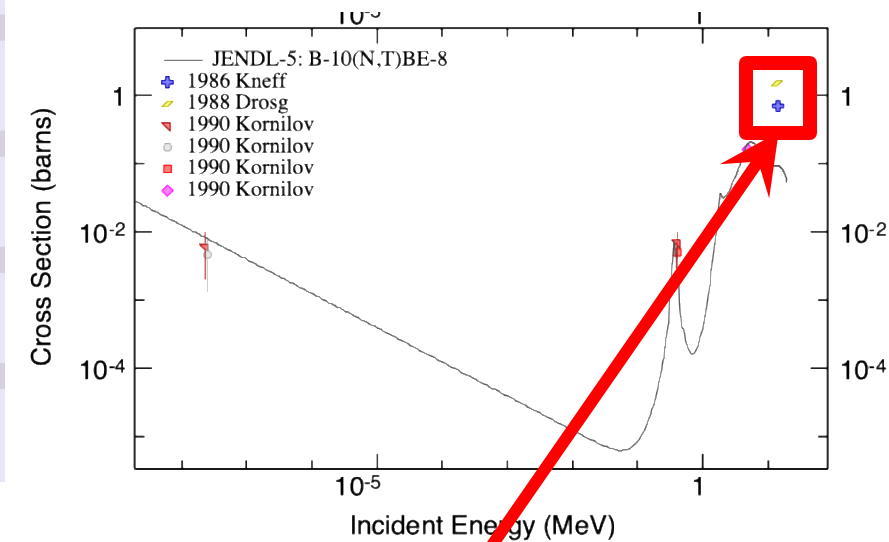
Tritium Breeding has focused on ${}^6\text{Li}(n,\alpha)t$ but ${}^{10}\text{B}$ might be of interest too

Here's all the available EXFOR data for ${}^{10}\text{B}(n,\alpha)$

n	Display	Year	Author-1	Energy range,eV	Points	Reference
1)	5-B-10(N,A)3-LI-7,,SIG	Q(keV)=2757.222	C4: MF=3	MT=107	Op=0	
Quantity: [CS] Cross section						
1	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2019	Haoyu Jiang+	1.00e0	2.40e6	57 + J,CPH/C,43,124,02,2019
2	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2017	R.Bevilacqua+	1.60e5	3.12e6	57 + J,EPJ/CS,143,11010,2017
3	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2017	Zhimin Wang+	4.00e6	5.00e6	8 + J,PR/C,88,044621,2017
4	<input checked="" type="checkbox"/> A <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2011	Zhang Guohui+	4.00e6	5.00e6	2 + J,CPH,28,(8),082801,2011
5	<input checked="" type="checkbox"/> A <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2008	Guohui Zhang+	4.00e6	5.00e6	2 + J,ARI,66,1427,2008
6	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	2002	Guohui Zhang+	4.17e6	6.52e6	4 + J,NSE,142,203,2002
7	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1975	S.J.Friesenhahn+	2.35e3	1.72e6	52 + C,75WASH,1,232,197503
* 8	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1969	D.Bogart+	2.90e4	8.17e5	27 + J,NP/A,125,463,69
9	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1968	R.L.Macklin+	9.80e4	5.05e5	5 + J,PR,165,1147,1968
10	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1967	S.A.Cox+	1.07e4	2.47e5	42 + J,JNE,21,271,1967
11	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1964	F.P.Mooring+	1.15e4	7.71e4	8 + P,ANL-6877,5,64
12	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov			1.15e4	7.71e4	8
13	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1963	A.Prosdocimi+	6.01e-3	8.20e-2	50 + J,JNE/AB,17,83,1963
14	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov			2.53e-2		2
* 15	<input checked="" type="checkbox"/> <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1960	E.G.Bilpuch+	3.00e3	6.80e4	32 + J,AP,10,455,60
16	<input checked="" type="checkbox"/> A <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1957	H.Bichsel+	2.00e4	4.78e6	98 + J,PR,108,1025,57
17	<input checked="" type="checkbox"/> A <input type="checkbox"/> i <input type="checkbox"/> X4 <input type="checkbox"/> X4+± <input type="checkbox"/> CSV+ <input type="checkbox"/> T4 <input type="checkbox"/> Cov	1944	C.L.Bailey+	5.80e5		1 + R,LA-46,1944

No measurements
above 6.52 MeV!

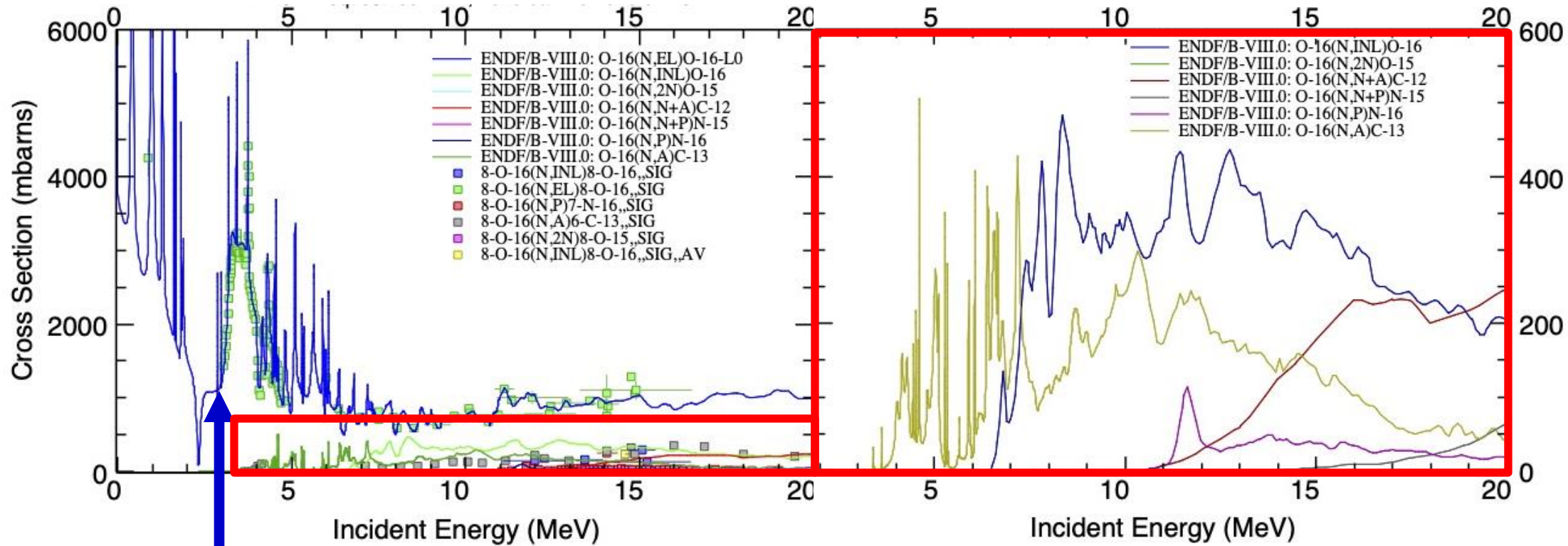
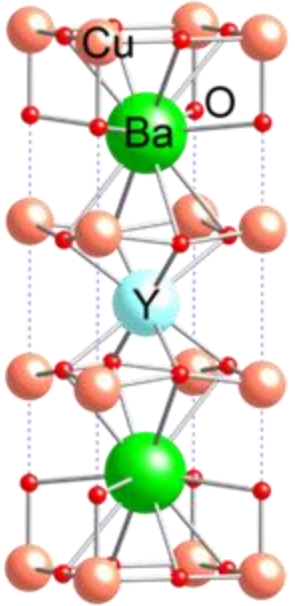
There's virtually no
data for ${}^{10}\text{B}(n,t)$ itself



*Additional measurements, especially
between 6.5 & 14 MeV are needed*

But 14 MeV looks
promising

Let's look at an important materials damage data issue for MCF: ^{16}O for REBCO (MCF magnets)

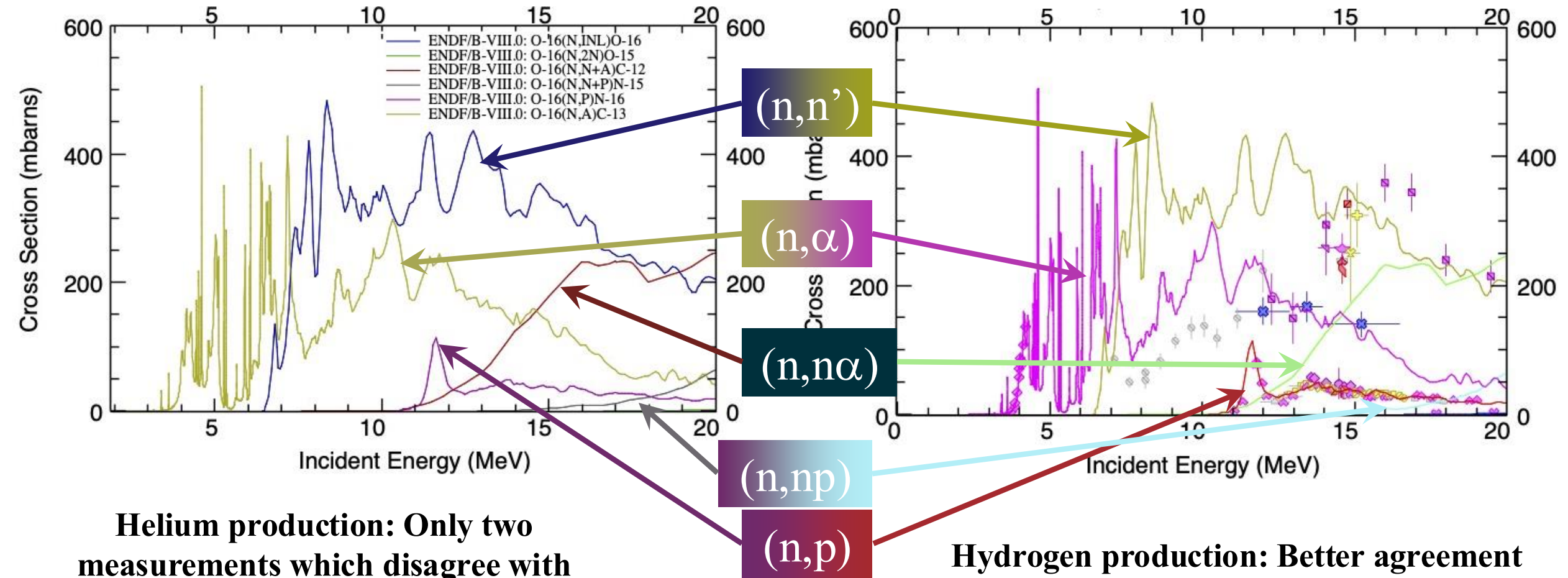


Elastic scatter (main constituent of σ_{DPA} at fast energies) has been measured for $E_n > 3$ MeV

The non-elastic channels are about as large at fast energies ($E_n > 3$ MeV) and they contribute to both σ_{DPA} AND gas production

What differential data exists for the ^{16}O non-elastic channels?

Inelastic scatter, including γ -ray production, is almost completely unmeasured)



Helium production: Only two measurements which disagree with each other and the evaluation

Hydrogen production: Better agreement between data sets and evaluation

Fortunately, parts of DOE have been investing in improving fast neutron scattering data, offering fusion a chance to address fusion needs incrementally

First Priority	Follow-up	Remaining	
H	He	F	Gd
C	Li	Mg	Bi
N	Be	P	Np
O	B	S	Am
Na	Cl	Ar	NA-22 has funded C & Na
Al	Ca	K	
Si	Mn	Ca	
Fe	Ni	Ti	
Cu	Ge	As	NA-113 has funded Bi & Nb
Pb	Br	Kr	
W	Cd	Mo	
U	I	Sn	
Pu	Cs	Sb	
	La	Xe	

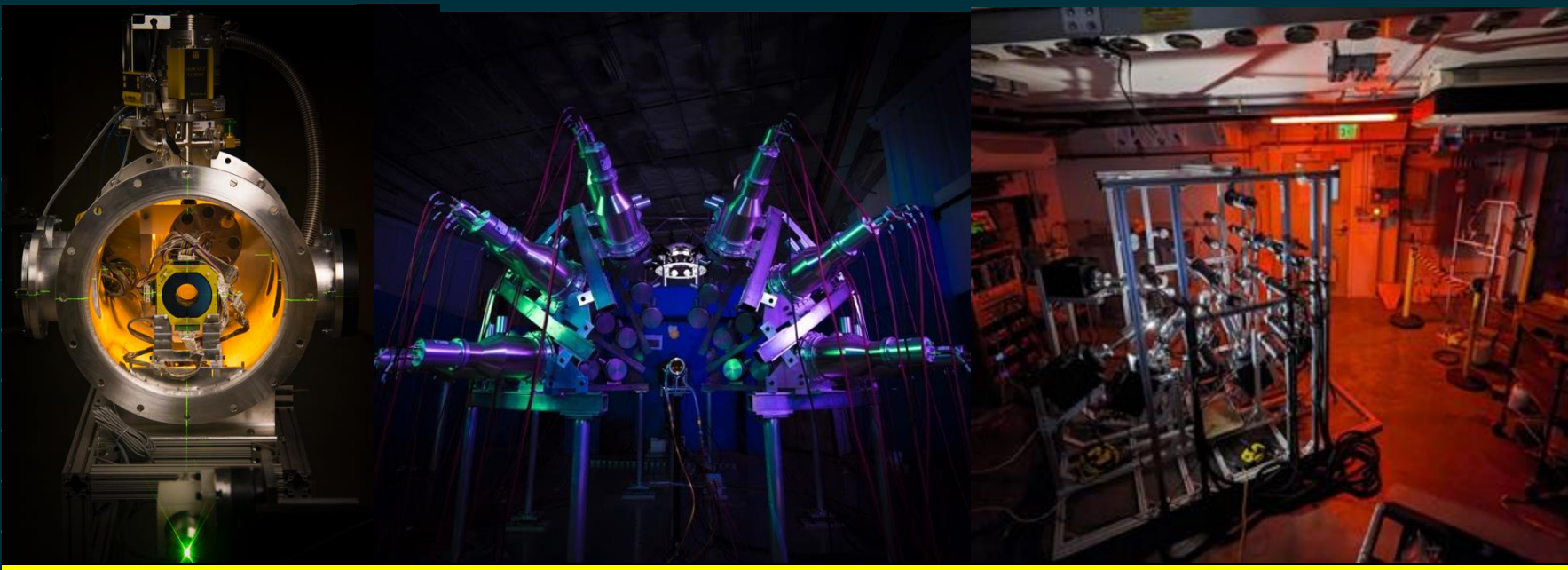
DOE-NE has funded Fe, Cl & U

Table 2 from the FY21 NA-22 portion of the Nuclear Data Interagency Working Group FOA

LENZ @ LANL
Gas Production
Cross Sections

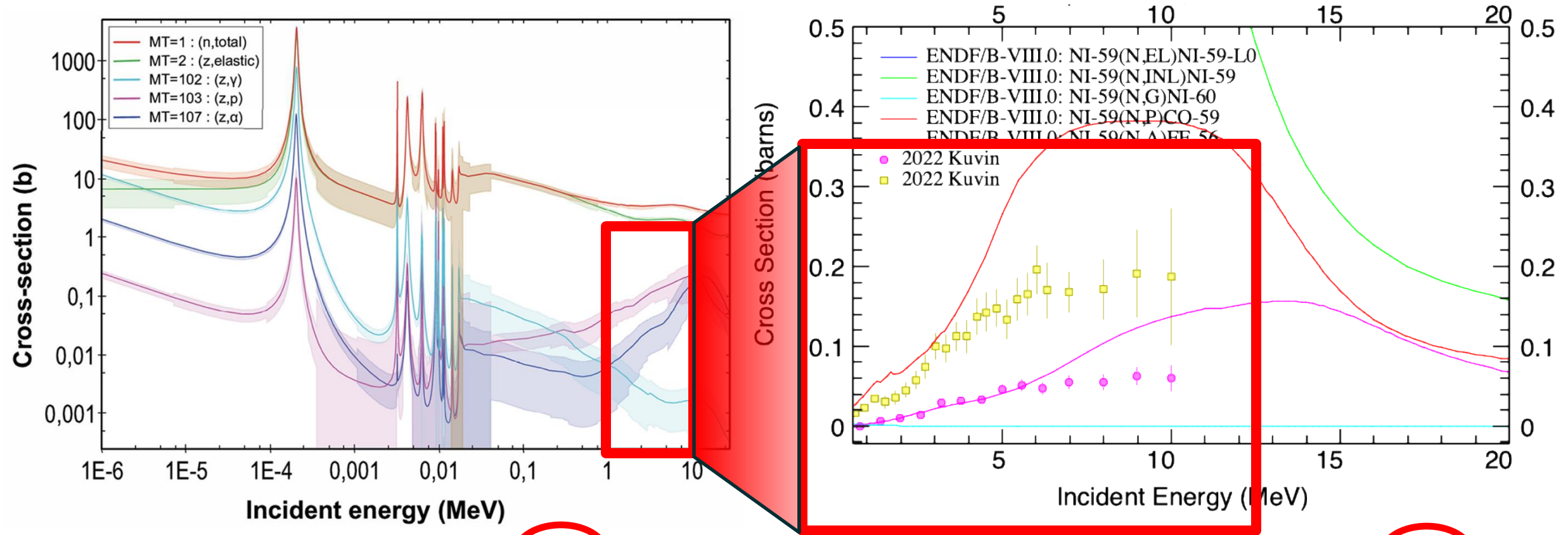
Chi-nu → CoGNAC @
LANL - Neutron Scatter
Cross Sections

GENESIS @ LBNL
Neutron Scatter and γ-ray
production Cross Sections



BUT: the measurement → evaluation time scale takes years, so advance planning using the WANDA/NDIAWG process is *essential*

Some nice work from LANL are highlighting these uncertainties, but absolute normalization is tricky and it often doesn't cover the entire energy range

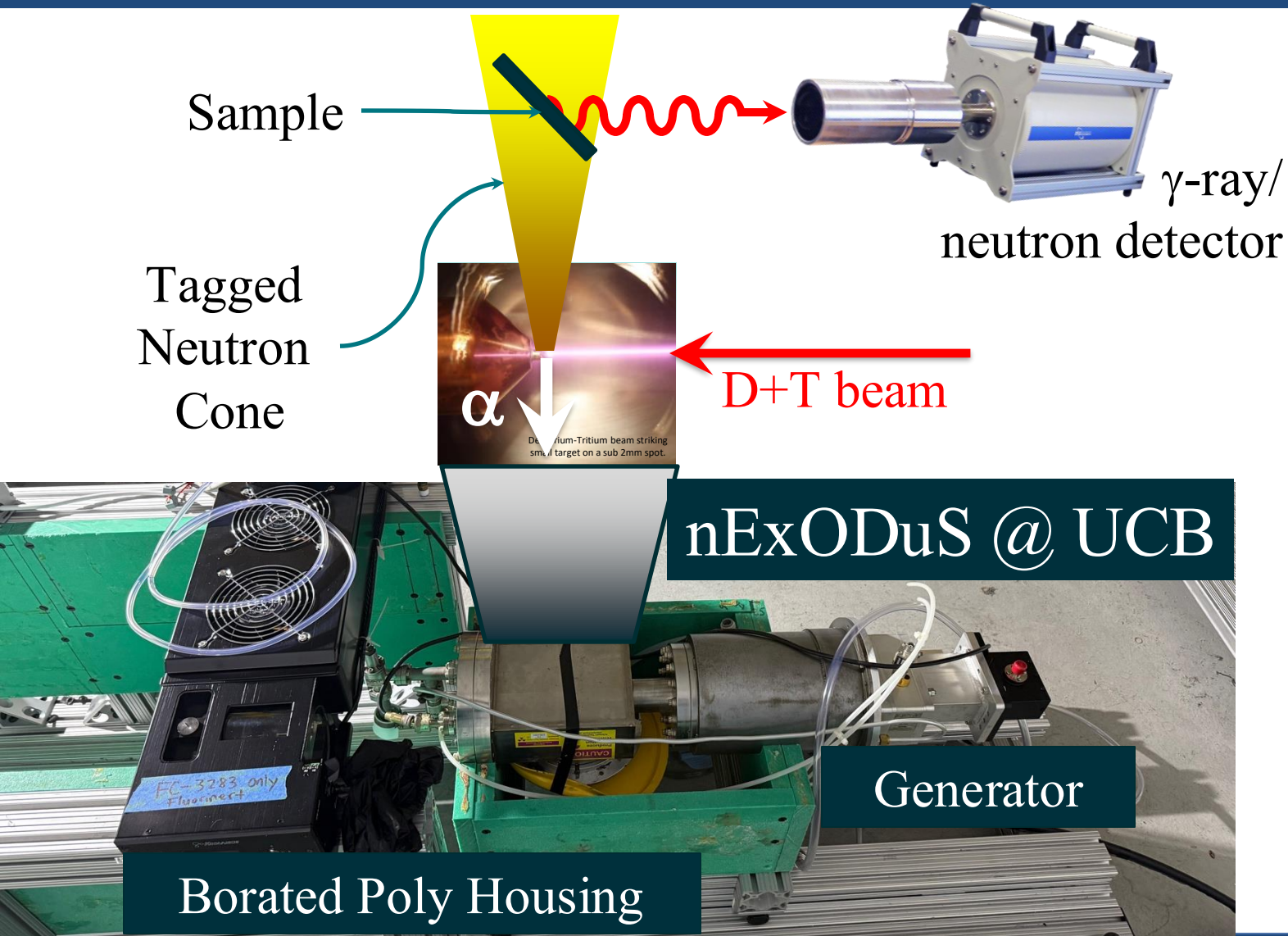


[J.-Ch. Sublet *et al.*, Eur. Phys. J. Plus \(2019\) 134: 350](#)

[S.A. Kuvin *et al.*, Phys. Rev. C 105, 044608 \(2022\)](#)

Absolute measurements at 14 MeV are needed

A joint JHUAPL/NNDC/Berkeley team* are using a DT-API system to measure $(n_{14}, x n \gamma)$ cross sections (*Thanks Keegan for the advance advertisement!*)



Simple Count Rate Estimate

$$R_{\alpha-\gamma/n} = \sigma(N)_{sample} \Phi_n$$

For a 10 g sample 1 m away from
a 1 cm scintillator @ 5 cm:

$$N_{sample} \approx 10^{23} \text{ atoms}$$

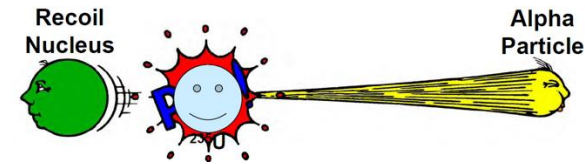
$$\Phi_n \approx 10^8 \frac{n}{s} \times \frac{1}{100} \rightarrow \approx 10^9 \frac{n}{day}$$

$$\rightarrow R_{\alpha-\gamma/n} = \frac{10^8}{\sigma (\text{barn}) \cdot day}$$

**10^5 peak counts/day for a
0.1% detection efficiency**

For Materials Damage let's look WAY BACK to WANDA 2019: Materials Damage Session Findings/Recommendations (C. Romano)

- ENDF is missing data for recoils and (n, α)
 - Inaccurate (n, α) data caused serious miscalculation of materials lifetimes
- Need to understand how changed materials properties change the neutronics
 - Transmutation
 - Porosity
 - Chemical bonds
- Processing of the data in NJOY needs modernization to meet the needs of the fusion community
- Current models do not determine the size of vacancies
- Stopping Powers are not well understood
- Radionuclides produced under transmutation might create further Primary Knock-on Atoms (PKAs) as they decay, and so the rate of these must also be quantified.
 - Gilbert, Mark & Sublet, Jean-Christophe (2016). PKA distributions: Contributions from transmutation products and from radioactive decay. Nuclear Materials and Energy. 9. 10.1016/j.nme.2016.02.006
- Inelastic Scatter cross sections need to be improved for fast neutron transport calculations
- Secondary particle production is not well known and requires measurement and theory development



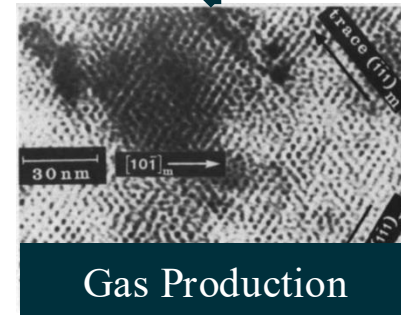
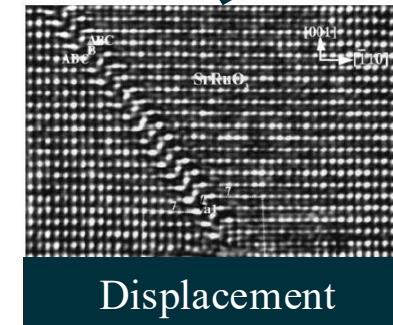
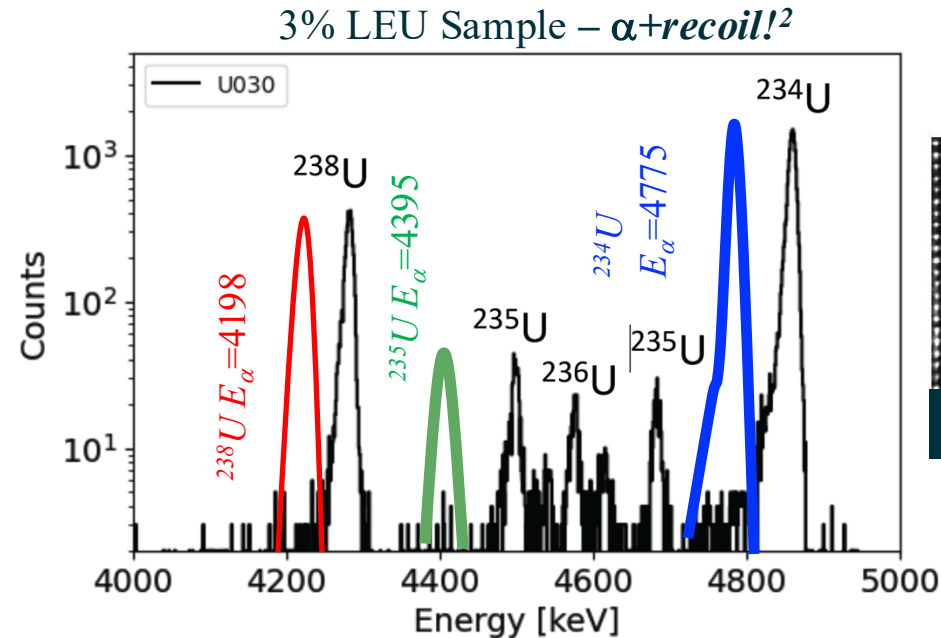
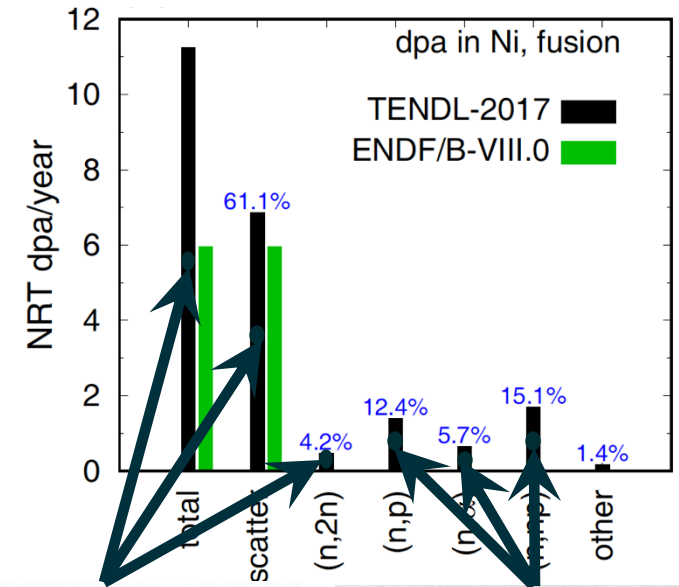
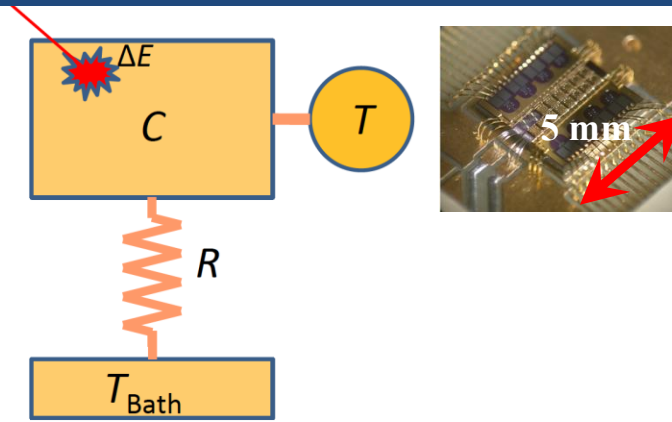
For Materials Damage let's look WAY BACK to WANDA 2019: Materials Damage Session Findings/Recommendations (C. Romano)

- **There should there be a coordinated and comprehensive materials damage database for validation of calculations**
- Idaho National Laboratory has a database of irradiation materials available to borrow
 - Includes irradiation data but no post irradiation testing data
- Improved dosimetry standards will be necessary for 14 MeV neutrons
- Post Irradiation Testing
 - There are several capable laboratories with full suite of testing
 - Impossible to measure sigma-dpa due to annealing: Can look at crystalline structure, but not accurate measure of dpa
- Need to standardize materials analysis methods and format for irradiated materials data

This may no longer be true!!!

Differential σ_{DPA} and gas production measurements using a microcalorimeter coupled to a DT-API neutron generator

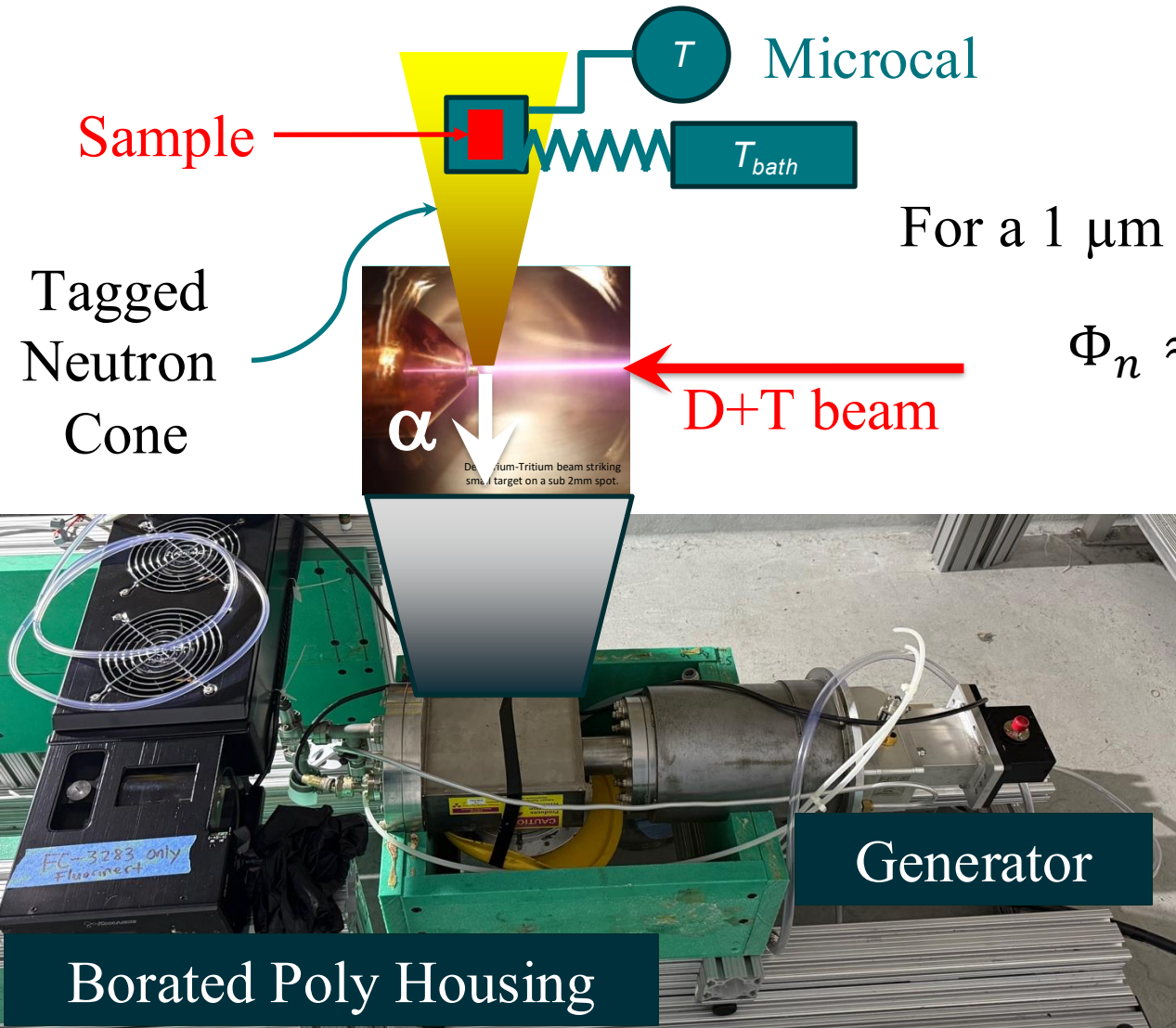
- Radiation is deposited in an absorber material and read out through a thermal resistance at low temperature.
- Source can be the absorber or embedded in it.



Radiation Threshold	
Silicon	3.6 eV
Germanium	2.9 eV
μCal	$\approx 30 \mu\text{eV}$

See Keegan Kelly's talk from yesterday

Differential σ_{DPA} and gas production measurements using a microcalorimeter coupled to a DT-API neutron generator



Simple Count Rate Estimate

$$R_{\alpha-\gamma/n} = \sigma(N)_{sample} \Phi_n$$

For a 1 μm thick sample ($N_{sample} \approx 10^{20} \text{ atoms}$) @ 10 cm:

$$\Phi_n \approx 10^8 \text{ n/s} \times 1/10^2 \rightarrow \Phi_n \approx 10^{11} \text{ n/day}$$

$$\rightarrow R_{\alpha-\gamma/n} = 10^7 / \sigma (\text{barn}) \bullet \text{day}$$

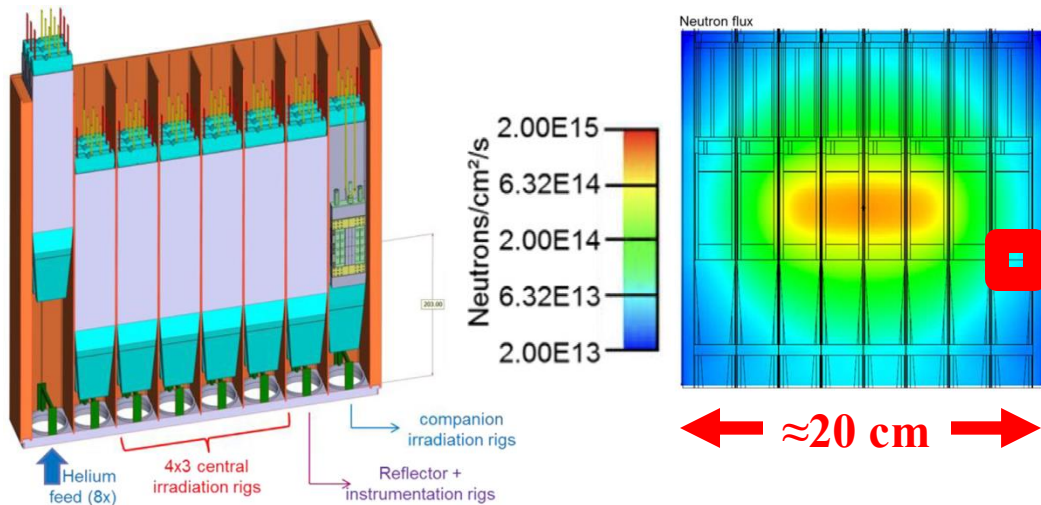
Energy deposited in ^{56}Fe by a 14.1 MeV neutron

(n,el) or (n,inl)	$\leq 247 \text{ keV}$	$\sigma_{14}^{max} \approx 0.6 \text{ b}$
(n,p) (Q=-3.0 MeV)	$\leq 11,38 \text{ keV}$	$\sigma_{14}^{max} \approx 0.1 \text{ b}$
(n, α) (Q=+0.32 MeV)	$\leq 14.7 \text{ MeV}$	$\sigma_{14}^{max} \approx 0.1 \text{ mb}$
(n,pn) (Q=-10.2 MeV)	$\leq 4.2 \text{ MeV}$	$\sigma_{14}^{max} \approx 0.1 \text{ b}$

All major parts of σ_{DPA} are resolvable

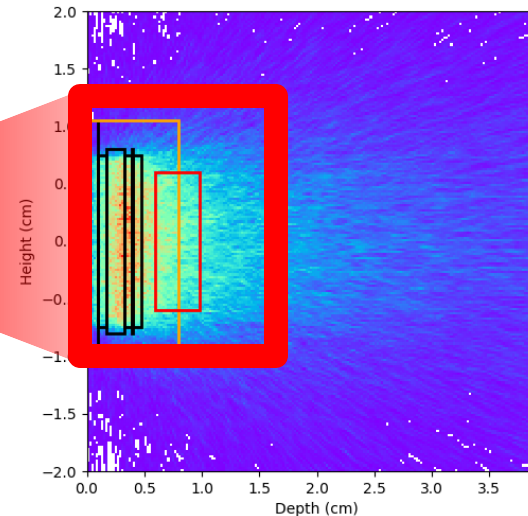
How about integral measurements using high-flux neutron beams from thick target deuteron breakup (IFMIF etc.)?

Future Facilities (IFMIF¹/FPNS)

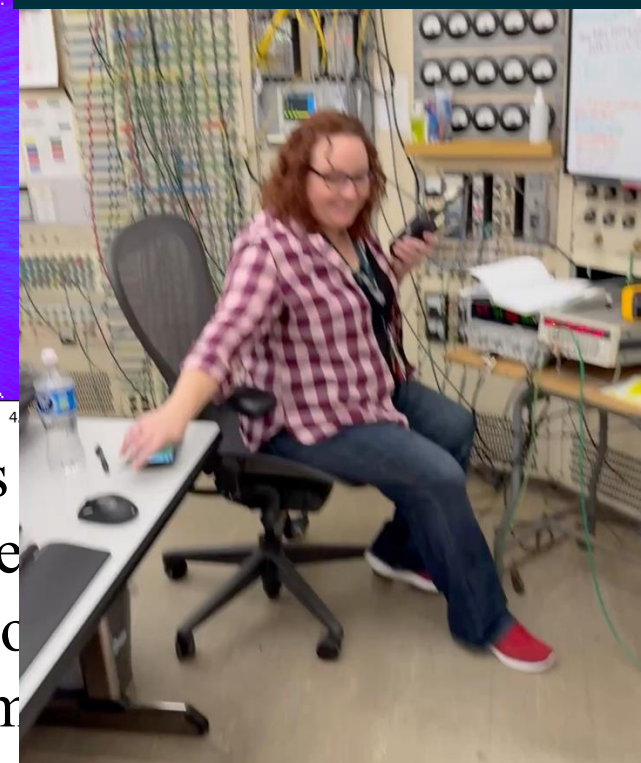


- IFMIF/FPNS plan to achieve multi-DPA damage in 100+ cm³ volumes using thick target deuteron breakup
Cost: $\approx \$10^{8-9}$ + years to design & build

Current Facilities



Countney Matzkind (MDA) this past Friday at the 88-Inch



- The same spectrum is neutron single-event and radioisotope production % of the flux in ≈ 1 cm

Integral tests can be performed now, but they are NOT a substitute for differential data

Conclusions

- Many nuclear data needs for fusion energy overlap with other programs, allowing them to benefit from work being done by DOE/NP, the NNSA, DOE-NE, BUT numerous gaps exist
 - Differential data is lacking for tritium breeding, materials damage and gamma-ray production
- The combination of DT neutron generators with Associated Particle Imaging and Microcalorimetry offers a chance to address nuclear data gaps relevant to fusion:
- Opportunities exist for fusion energy to partner with other nuclear data users to accelerate progress and mitigate risk.

Thanks for your attention!