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Beta-delayed neutron emission near ^{54}Ca

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Beta-delayed neutron emission (β -n) plays a vital role in shaping the elemental abundance distribution in the r -process via modifying the decay path back to stability and by contributing significantly to the neutron flux after freeze-out [1]. Modeling the β -n process requires a good model of the beta-strength functions and of the neutron emission mechanism. Statistical neutron-emission models assume gamma and neutron decay from a compound nucleus following beta decay and have successfully predicted gross properties of β -n probabilities (P_{1n} , P_{2n} , etc) in some medium- and heavy-mass nuclei. However, recent experimental work found evidence of non-statistical neutron emission after the beta decay in the vicinity of doubly magic ^{132}Sn [2]. Therefore, it is of great importance to study β -n spectroscopy in a broader area of the nuclear chart to provide stringent experimental constraints to the theories, which in astrophysical applications predict those properties for many more r -process nuclei currently out of experimental reach.

We expanded our study onto the nuclei near ^{54}Ca , which is thought to be doubly magic. An experiment studying β -n spectroscopy of $^{52,53}\text{K}$ was carried out at the ISOLDE Decay Station (IDS). These isotopes have large Q_β values (energy window for β decay) and low neutron-separation energies in their daughters ($^{52,53}\text{Ca}$ respectively), making them ideal for the studies of the β -n process. In coincidence with the beta decay of $^{52,53}\text{K}$, gamma-ray and neutron-time-of-flight (TOF) spectra were measured using HPGe clover detectors and VANDLE [3]. In this contribution, I will present the latest results from the experiment, including the reconstructed excitation energies (E_x) and apparent beta feeding (I_β) of the neutron unbound states in $^{52,53}\text{Ca}$, together with their exclusive neutron-emission branching ratios to the states in $^{51,52}\text{Ca}$, respectively. The experimental findings were compared with the shell-model calculations and Hauser-Feshbach statistical model. These comparisons provide valuable insights into the β -n process and its connection with the nuclear structure far from the stability line.

[1] R. Yokoyama et al, Phys. Rev. C 100, 031302(R) (2019).

[2] J. Heideman et al, being reviewed by Phys. Rev. Lett.

[3] M. Madurga et al, Phys. Rev. Lett 117, 092502 (2016).

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