

Past and Present Experiments of Geo Neutrinos

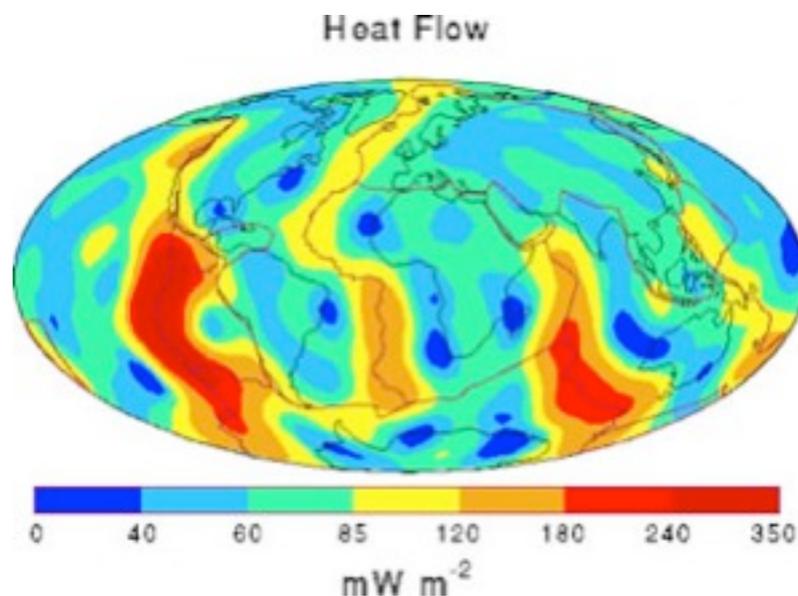
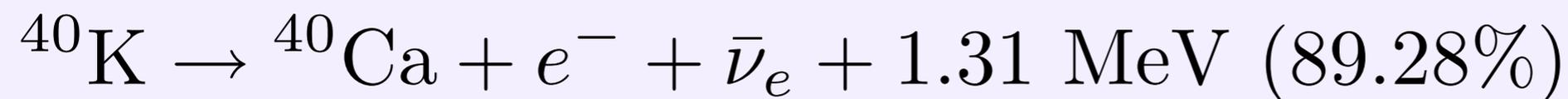
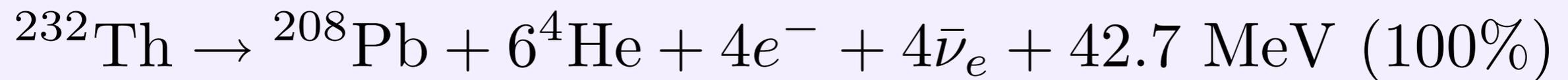
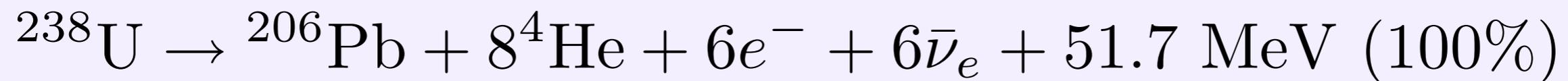
The 13th International Conference on Topics
in Astroparticle and Underground Physics

Sep. 9, 2013

Itaru Shimizu (Tohoku Univ.)

Geologically Produced Anti-Neutrino

Beta-decay of radioactivities (U, Th, K) in the Earth



Surface heat flow
44 TW

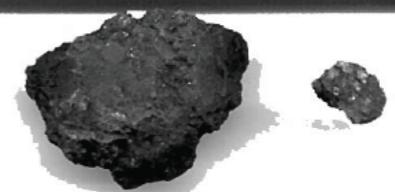
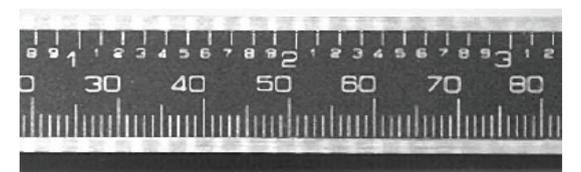
Bulk Silicate Earth (BSE) model

chondrite meteorite

U : 8 TW

Th : 8 TW

K : 4 TW



Radiogenic heat
19 TW



Geo Neutrino

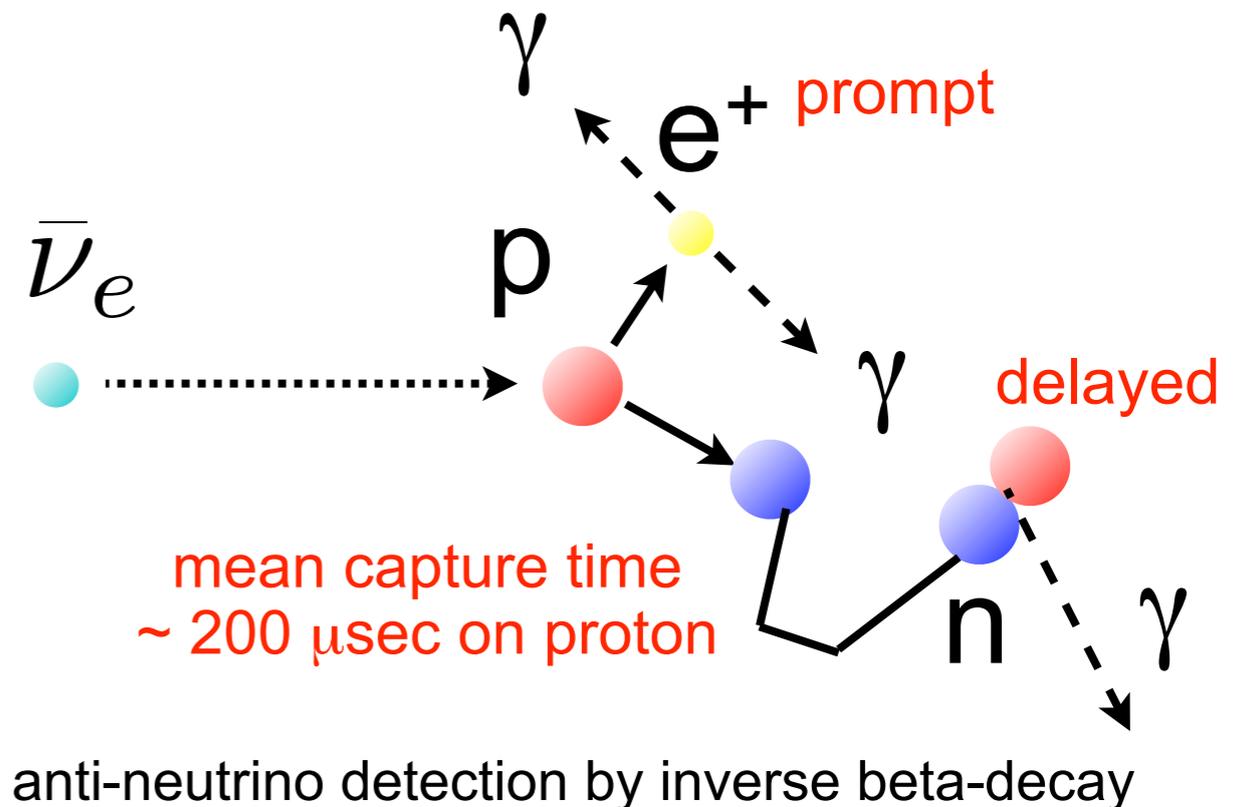
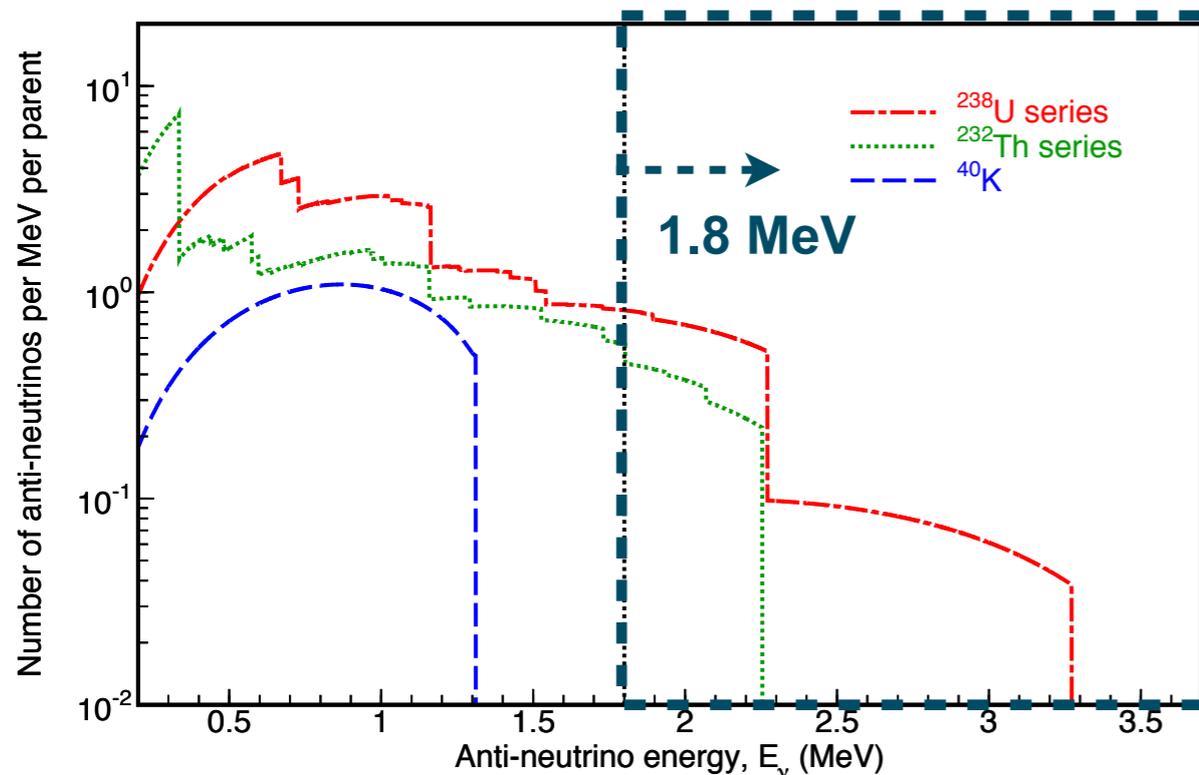
- G. Eder (1966)
- G. Marx (1969)
- L. Krauss et al. (1988)
- M. Kobayashi, Y. Fukao (1981)
- R. Raghavan et al. (1998)
- Rothschild et al. (1998)
- G. Fiorentini et al. (2003)

first calculation in science literature

systematic search of target detector material

feasible plan in KamLAND and Borexino

detailed neutrino flux calculations



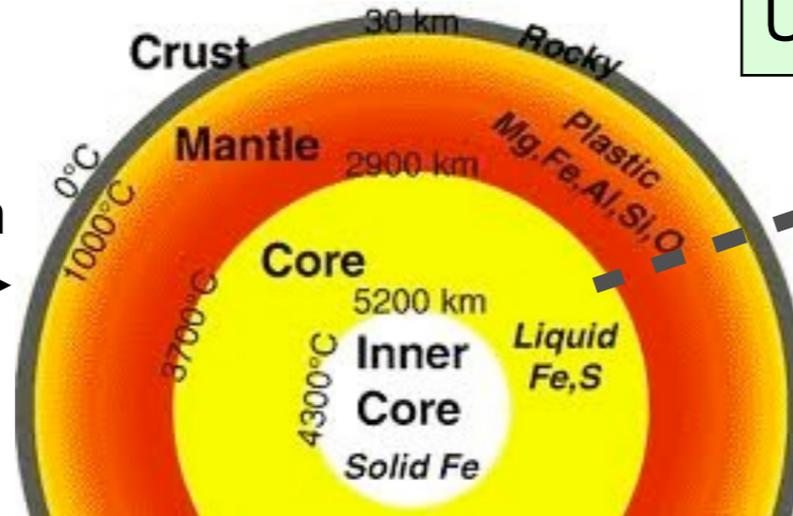
Neutrino Geoscience

Heat sources in the Earth

~ 4 billion years ago



Earth formation



Energy release by radioactive decay of U, Th, K → **radiogenic heat**

Release of gravitational energy through metallic core separation → **primordial heat**

↓
still remain ?

5 Big Questions:

- What are earth's K/U & Th/U ratios?
planetary volatility curve
- Radiogenic contribution to heat flow?
secular cooling
- Distribution of reservoirs in mantle?
whole vs layered convection
- Radiogenic elements in the core??
Earth energy budget
- Nature of the Core-Mantle Boundary?
hidden reservoirs

Experimentally investigated

Geo neutrino detector

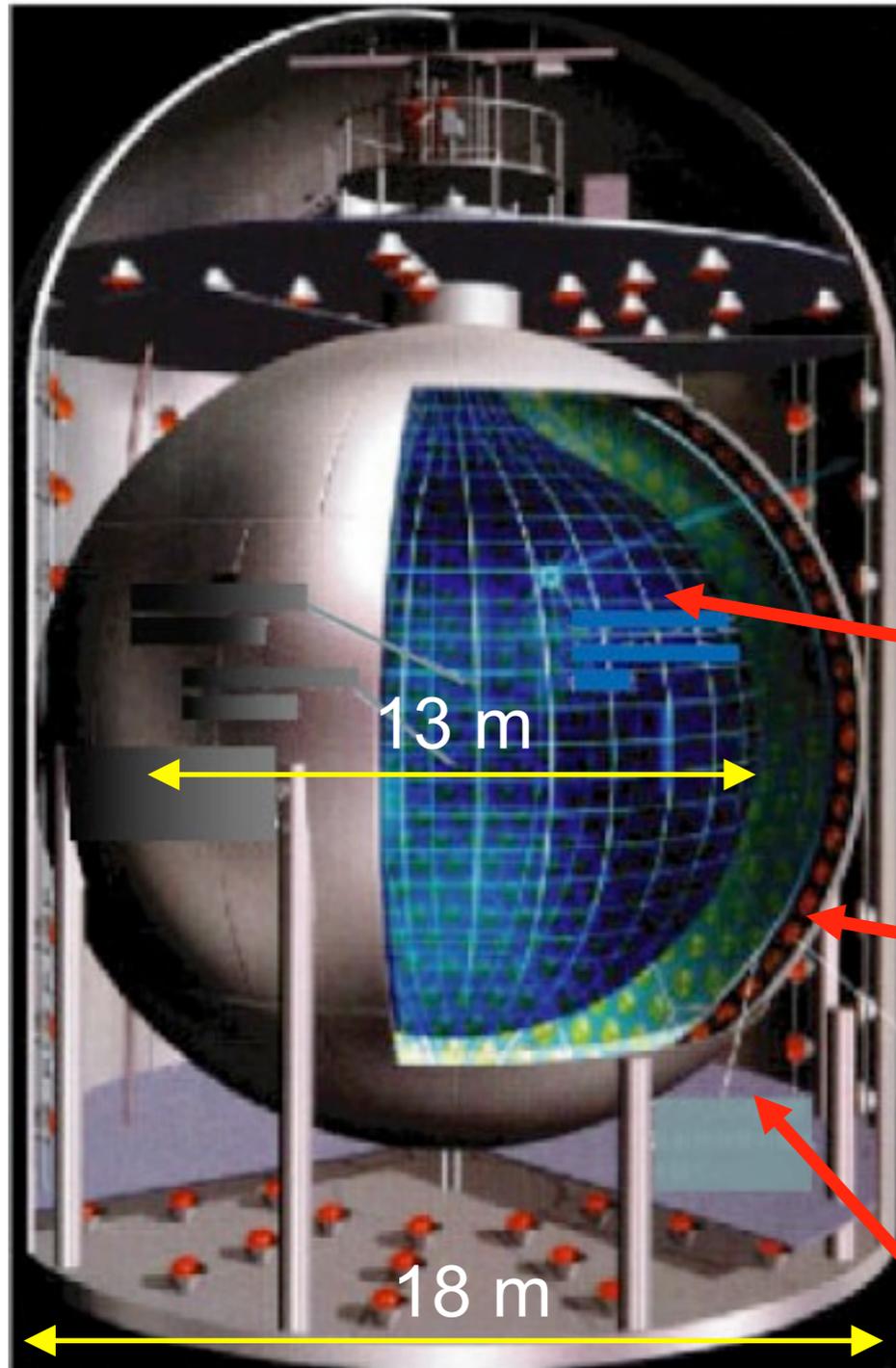
- KamLAND (Japan)
- Borexino (Italy)

Geo neutrino experiment will play a key role in answer all the questions !

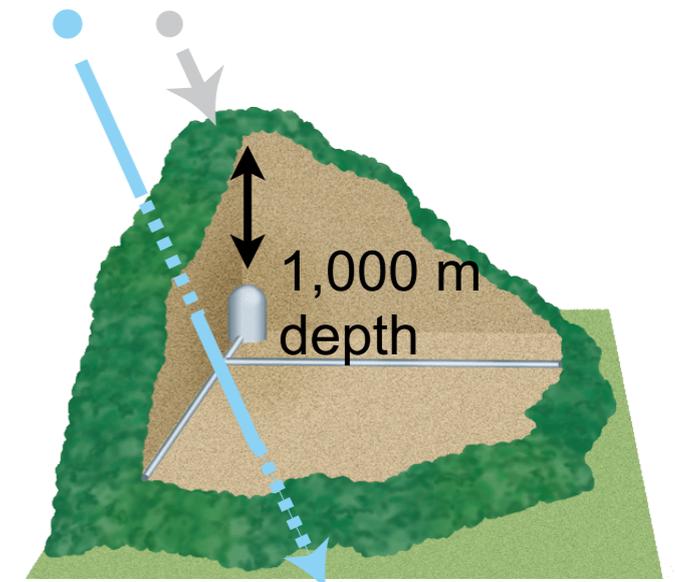
KamLAND

Kamioka Liquid Scintillator Anti-Neutrino Detector

operated since 2002



neutrino, cosmic-ray



1,000 ton Liquid Scintillator

Dodecane (80%) Pseudocumene (20%) PPO (1.36 g/l)

1,325 17 inch + 554 20 inch PMTs

commissioned in February, 2003

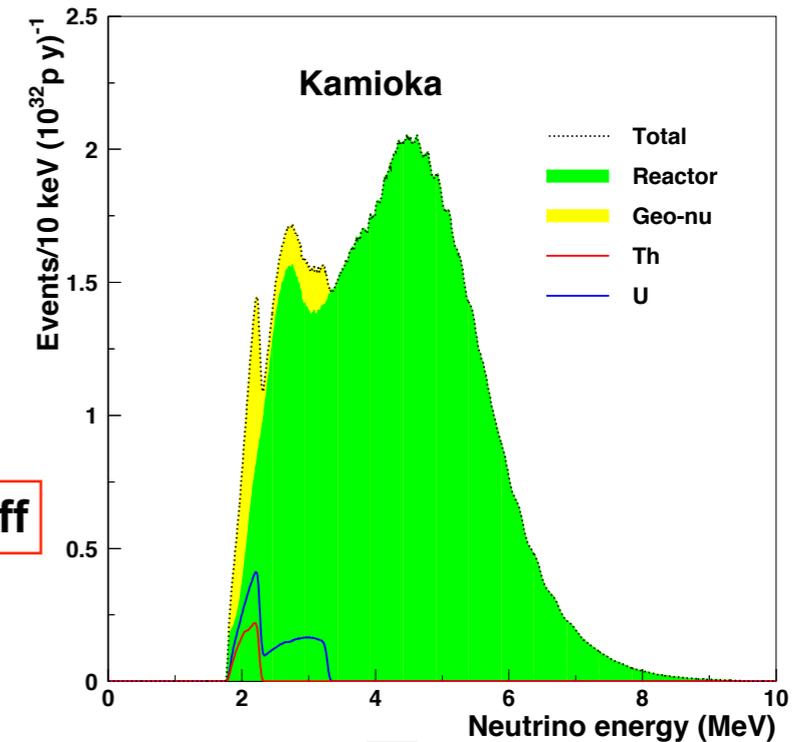
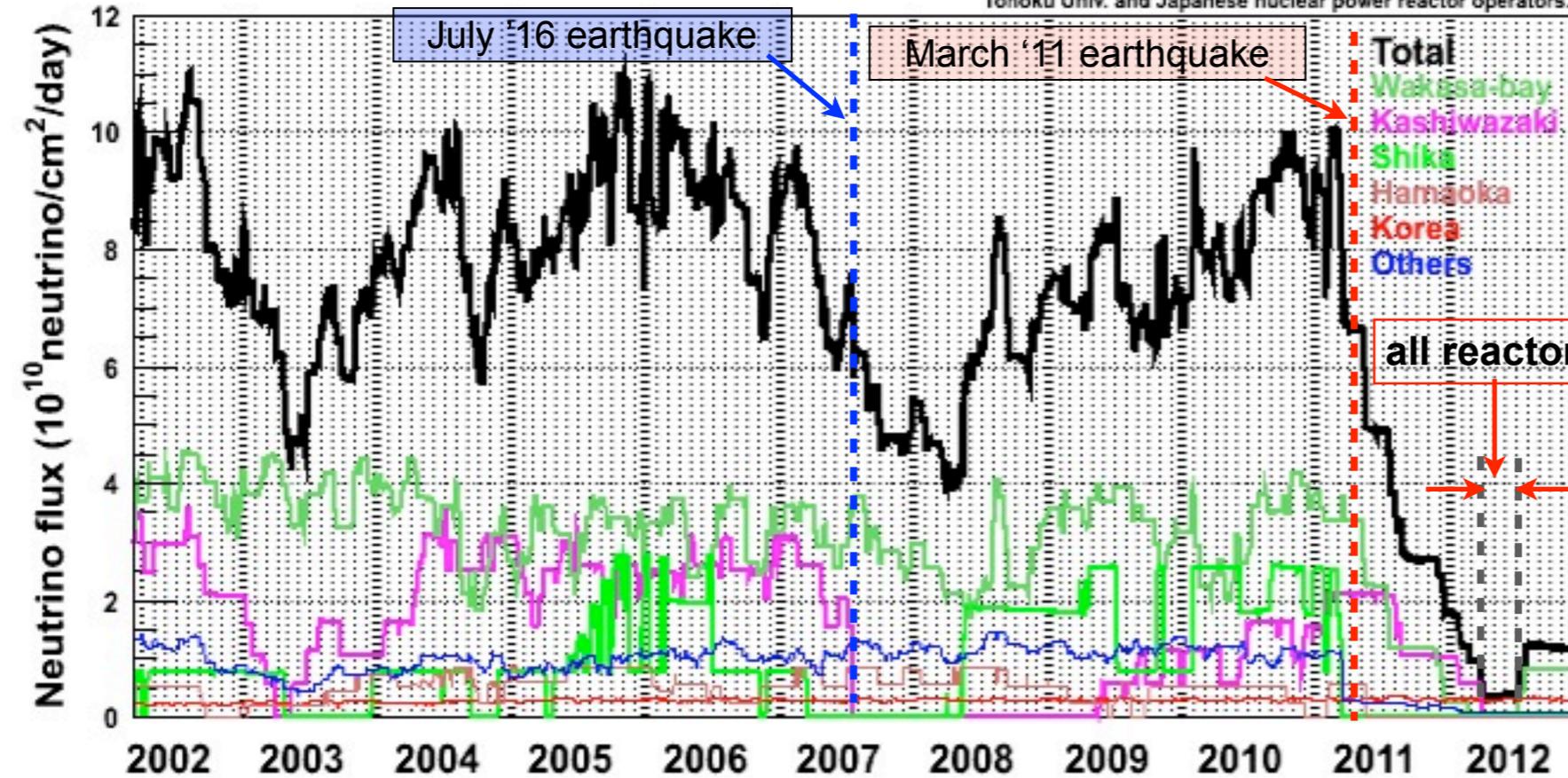
photocathode coverage : 22% → 34%

Water Cherenkov Outer Detector

Anti-Neutrino Flux in Kamioka

time variation of neutrino flux

Data provided according to the special agreements between Tohoku Univ. and Japanese nuclear power reactor operators.

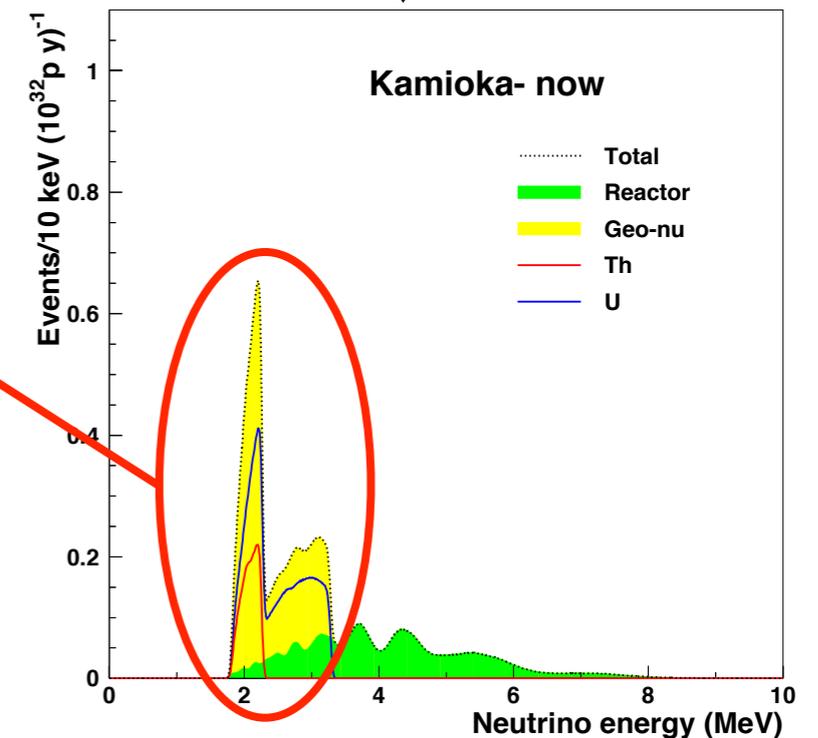


decrease of reactor neutrino

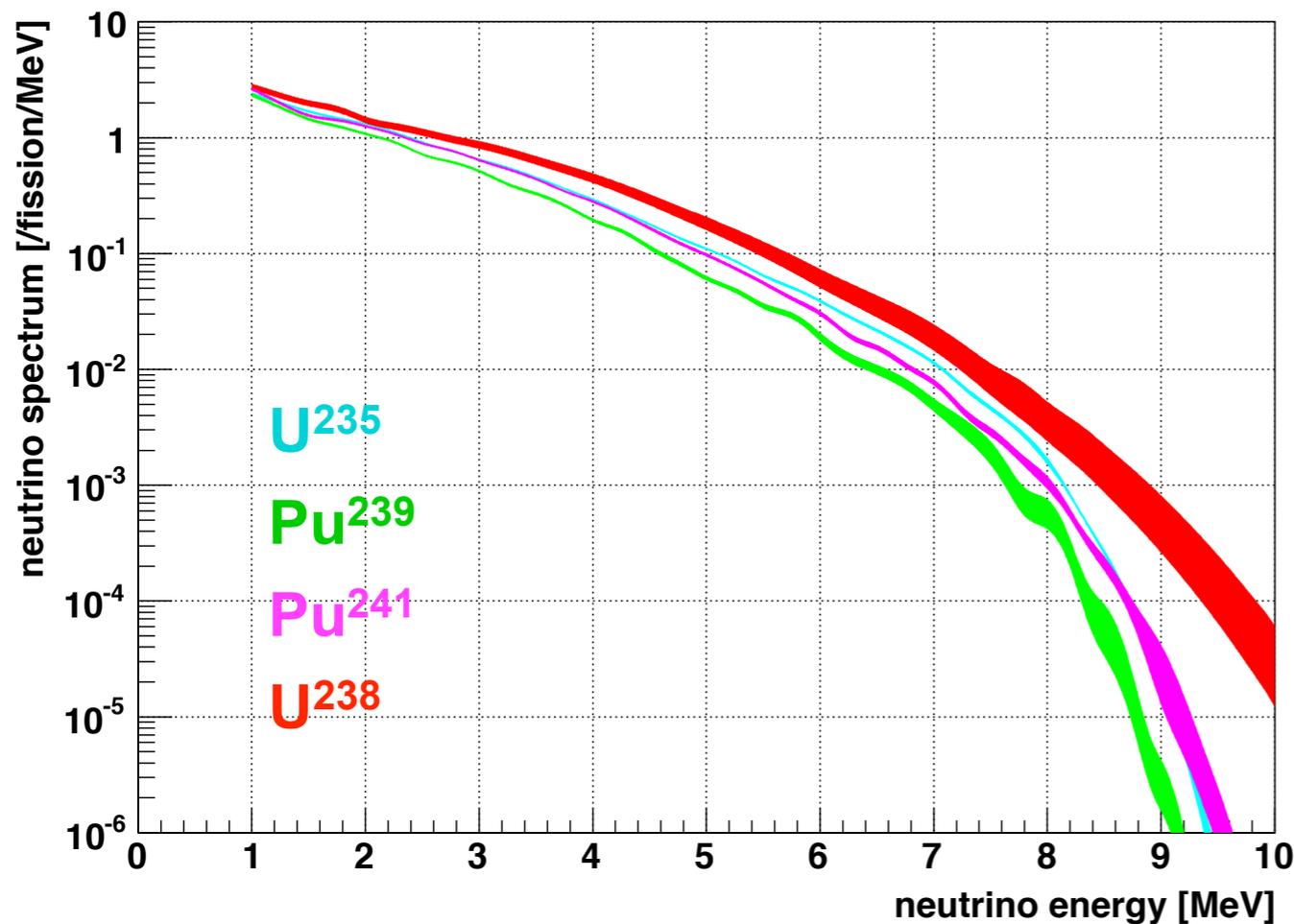
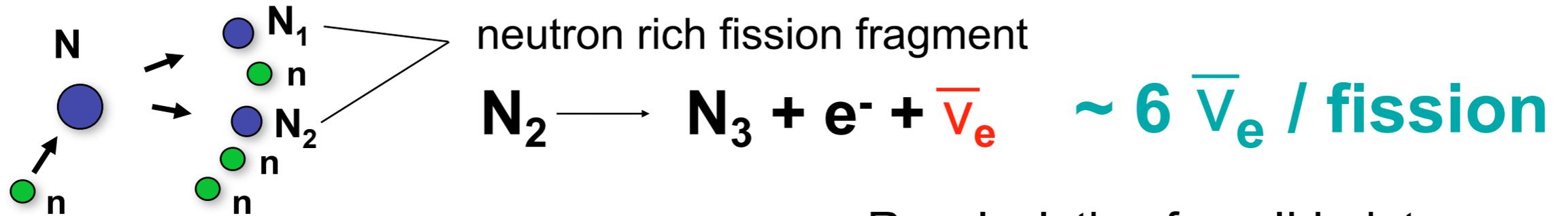
significant reduction of anti-neutrino flux from reactors after Fukushima-I accident

good data for geo neutrino observation

“Reactor on-off” study for neutrino oscillation and geo neutrino analysis



Reactor Neutrino Spectrum

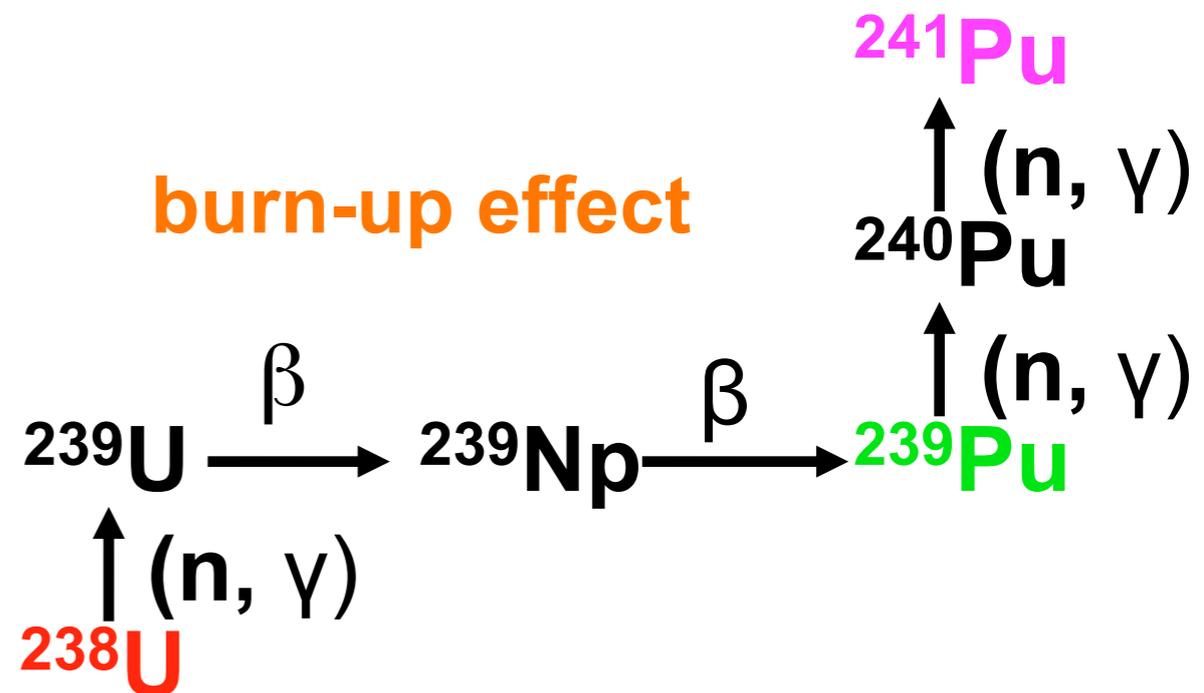


^{235}U , ^{239}Pu , ^{241}Pu : re-evaluation of ILL (P. Huber)
 ^{238}U : theoretical calculation (Th. Mueller et al.)

Recalculation from ILL data

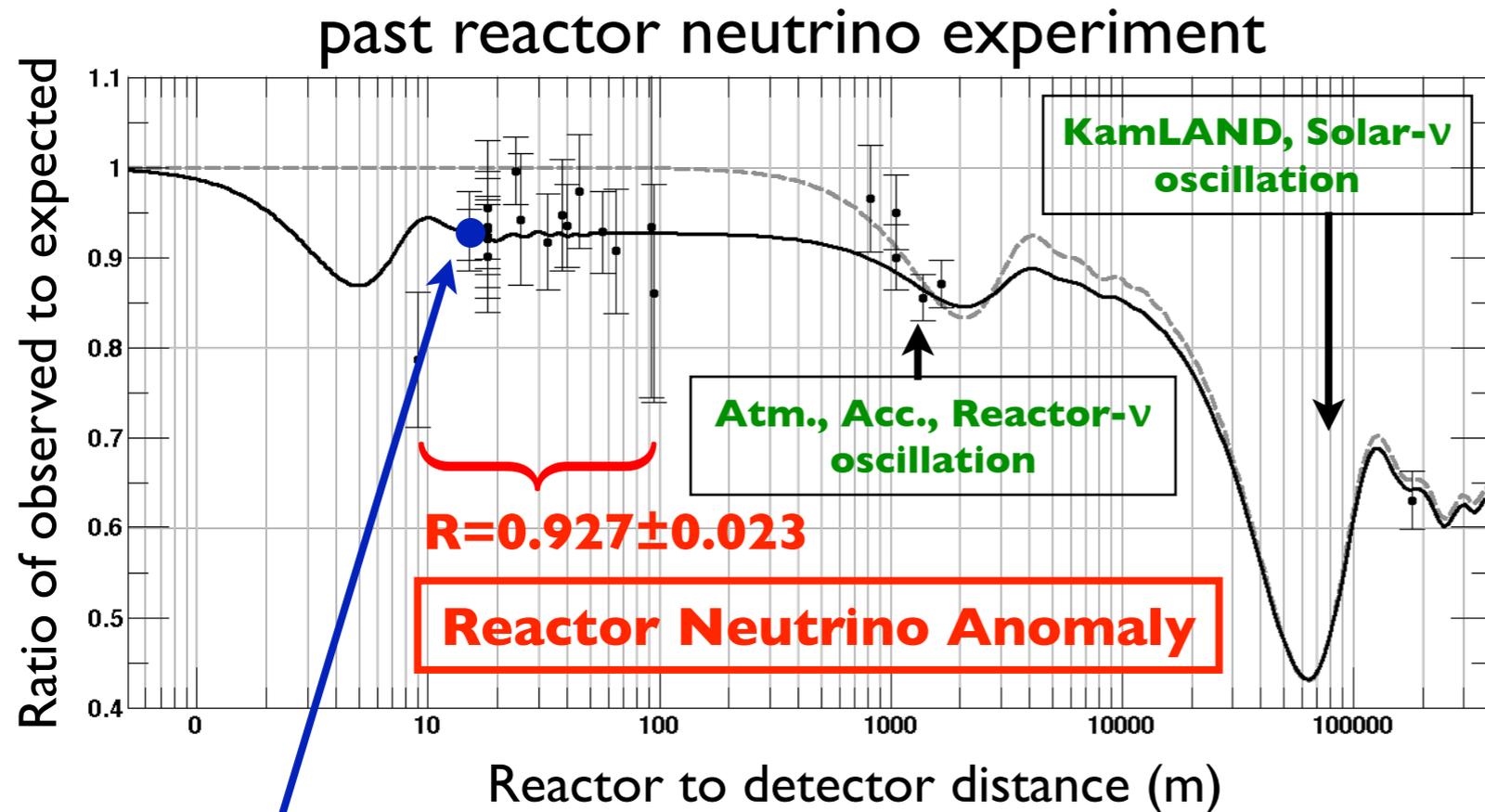
(Lhuillier/Mueller, Saclay)

- built from 845 nuclei and 10000 β -branches
- introduce a new ab-initio conversion method
- includes full error propagation and correlation



$\sim 3\%$ upward-shift by re-evaluation of ILL data

Normalization with Near Data



possibility of systematic bias or sterile- ν oscillation?

Bugey4 accuracy on the neutrino flux is **1.4%** ($L = 15$ m from PWR reactor)

Normalization of cross section per fission

$$\langle \sigma \rangle_{\text{reac.}} = \langle \sigma \rangle_{\text{Bugey4}} + \sum_i (\alpha_i^{\text{reac.}} - \alpha_i^{\text{Bugey4}}) \langle \sigma \rangle_i \quad (i = {}^{235}\text{U}, {}^{238}\text{U}, {}^{239}\text{Pu}, {}^{241}\text{Pu})$$

α_i : fractional fission rate $\langle \sigma \rangle_i = \int dE S_i(E) \sigma(E)$: effective cross section per fission

analysis is insensitive to "Reactor Neutrino Anomaly"

Time Variation of Event Rate

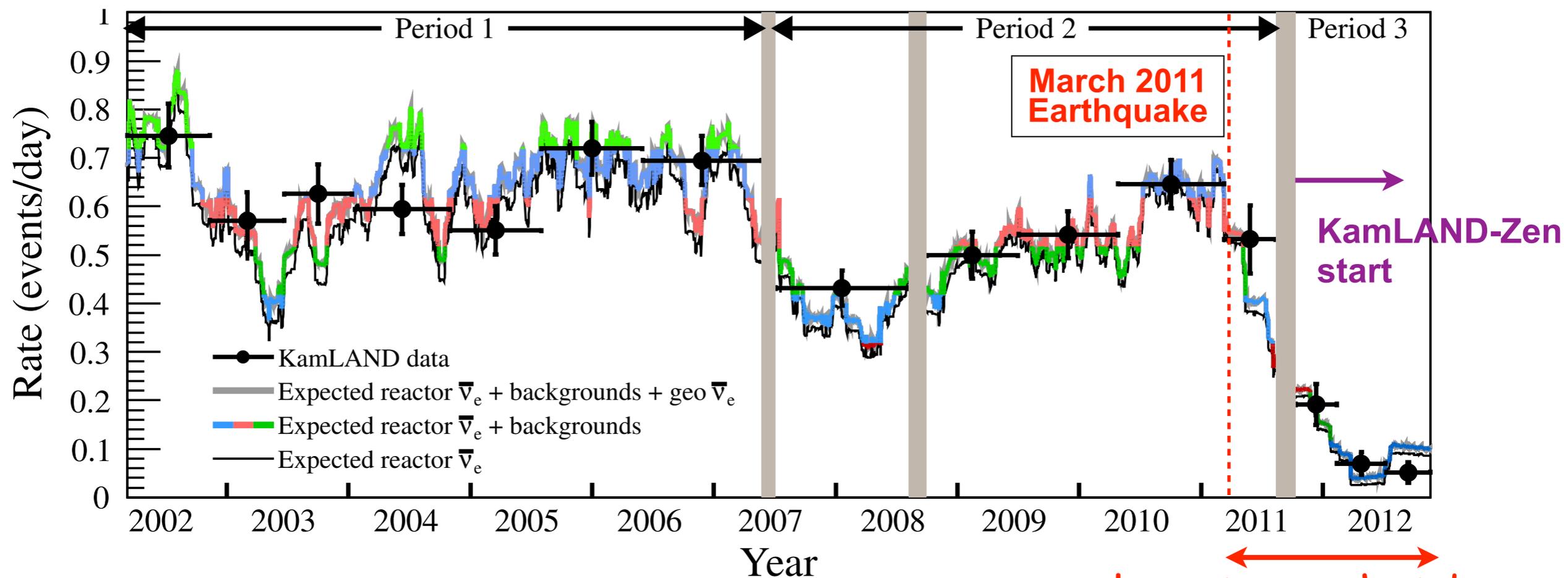
Total livetime
2991 days

Period 1: Mar. 2002 - May 2007

Period 2: May 2007 - Aug. 2011 (after LS purification)

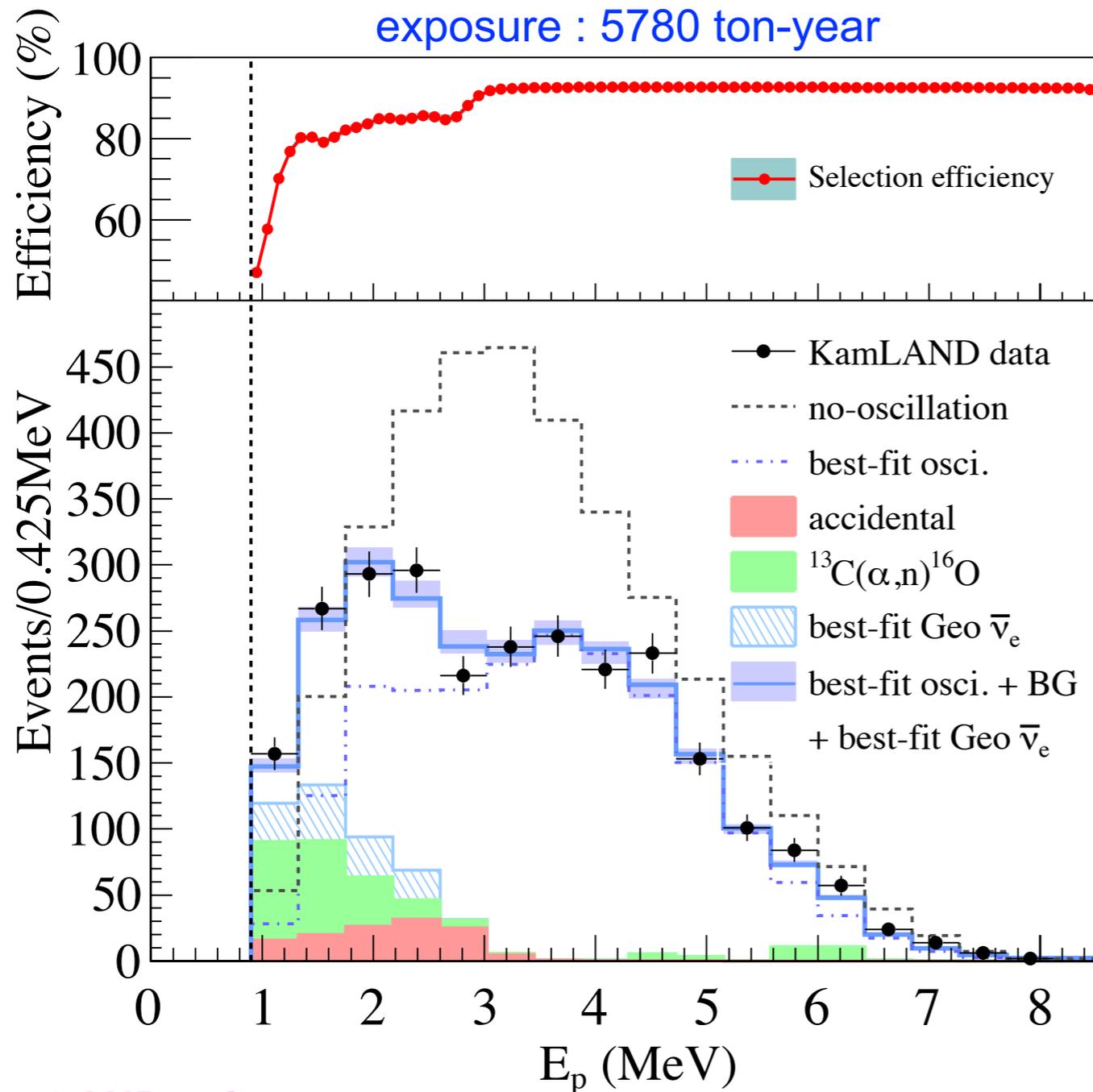
Period 3: Oct. 2011 - Nov. 2012 (after KamLAND-Zen start)

$$2.6 < E_p < 8.5 \text{ MeV}$$

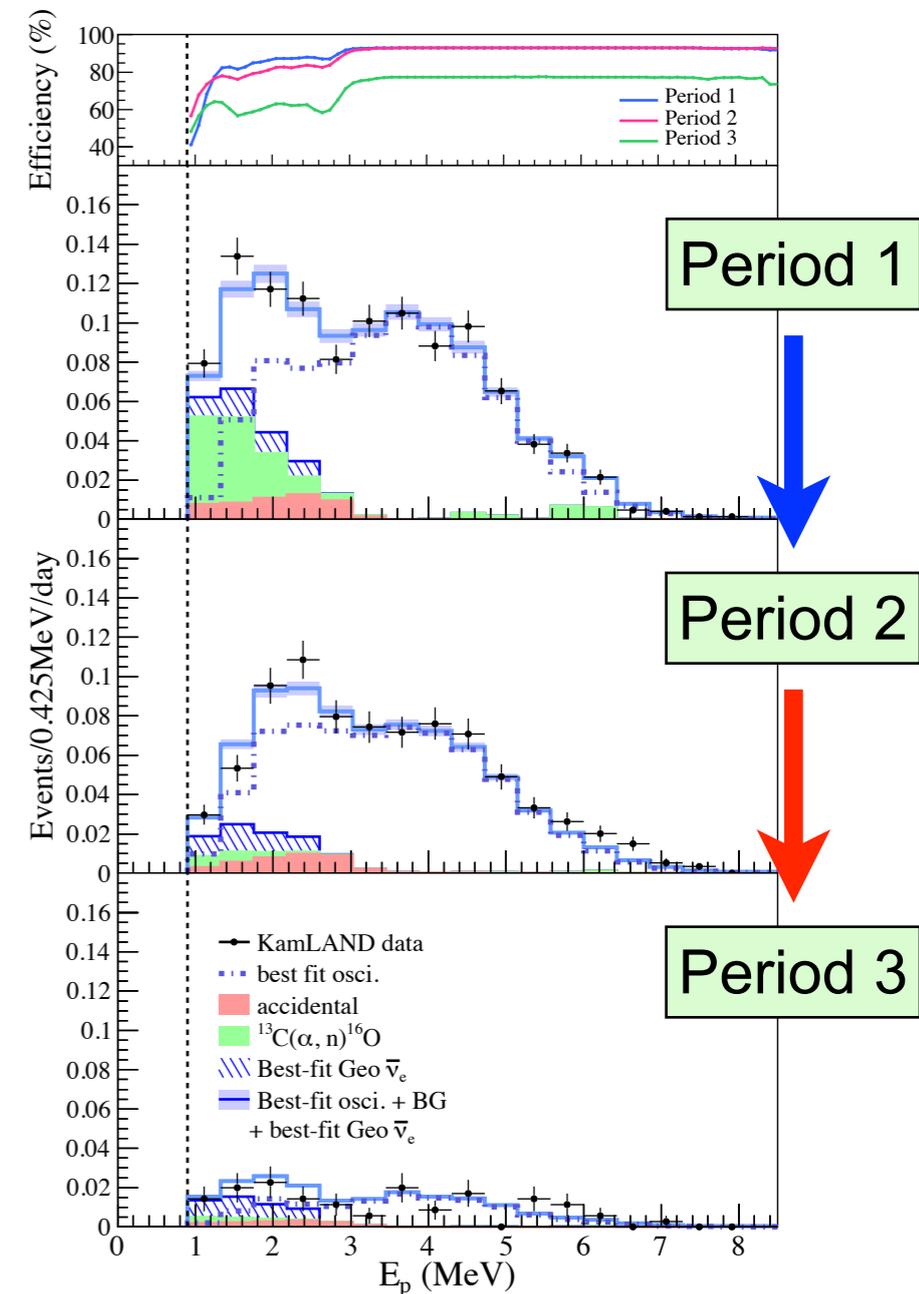


Data have good agreement with expected rate

Observed Energy Spectrum



purification (a,n) ↓
 earthquake reactor ↓



KamLAND only

$$\Delta m_{21}^2 = 7.54^{+0.19}_{-0.18} \times 10^{-5} \text{eV}^2$$

$$\tan^2 \theta_{12} = 0.481^{+0.092}_{-0.080}$$

$$\sin^2 \theta_{13} = 0.010^{+0.033}_{-0.034}$$

No osci. expected 3564 ± 145

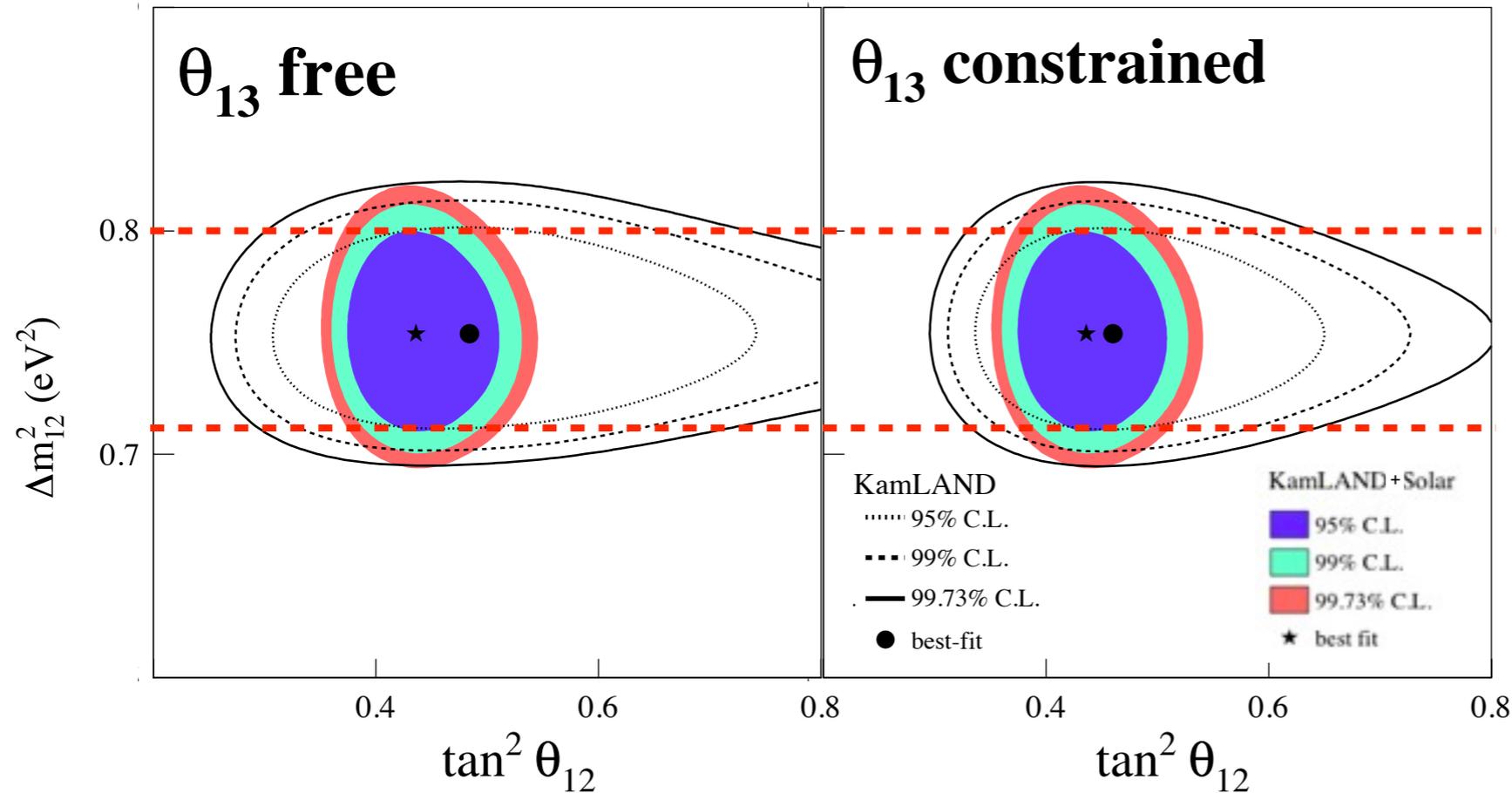
Background (w/o geo neutrino) 364 ± 31

Observed events 2611

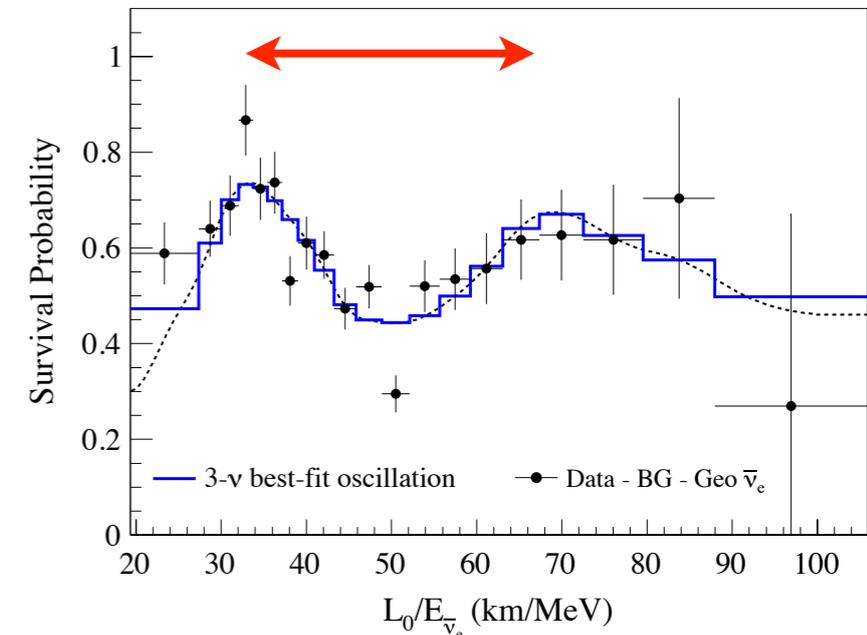
significant reduction

Neutrino Oscillation Parameter

Δm^2 : systematic uncertainty 1.9%
(dominated by linear energy scale uncertainty)



frequency meas.
by KamLAND



KamLAND+Solar

$$\Delta m_{21}^2 = 7.53^{+0.19}_{-0.18} \times 10^{-5} \text{ eV}^2$$

$$\tan^2 \theta_{12} = 0.437^{+0.029}_{-0.026}$$

$$\sin^2 \theta_{13} = 0.023^{+0.015}_{-0.015}$$

KamLAND+Solar+ θ_{13}

$$\Delta m_{21}^2 = 7.53^{+0.18}_{-0.18} \times 10^{-5} \text{ eV}^2$$

$$\tan^2 \theta_{12} = 0.436^{+0.029}_{-0.025}$$

$$\sin^2 \theta_{13} = 0.023^{+0.002}_{-0.002}$$

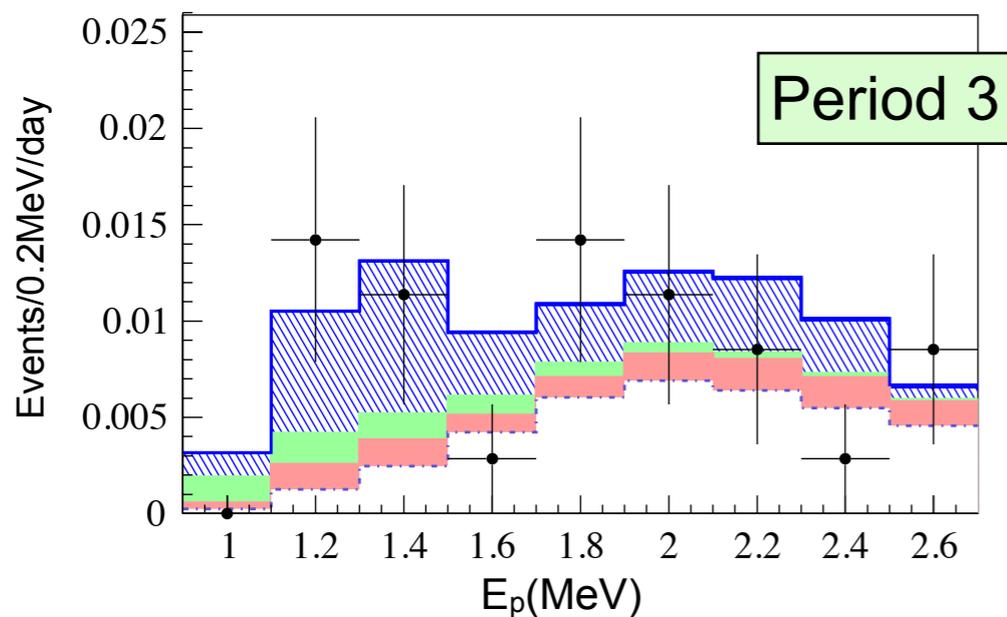
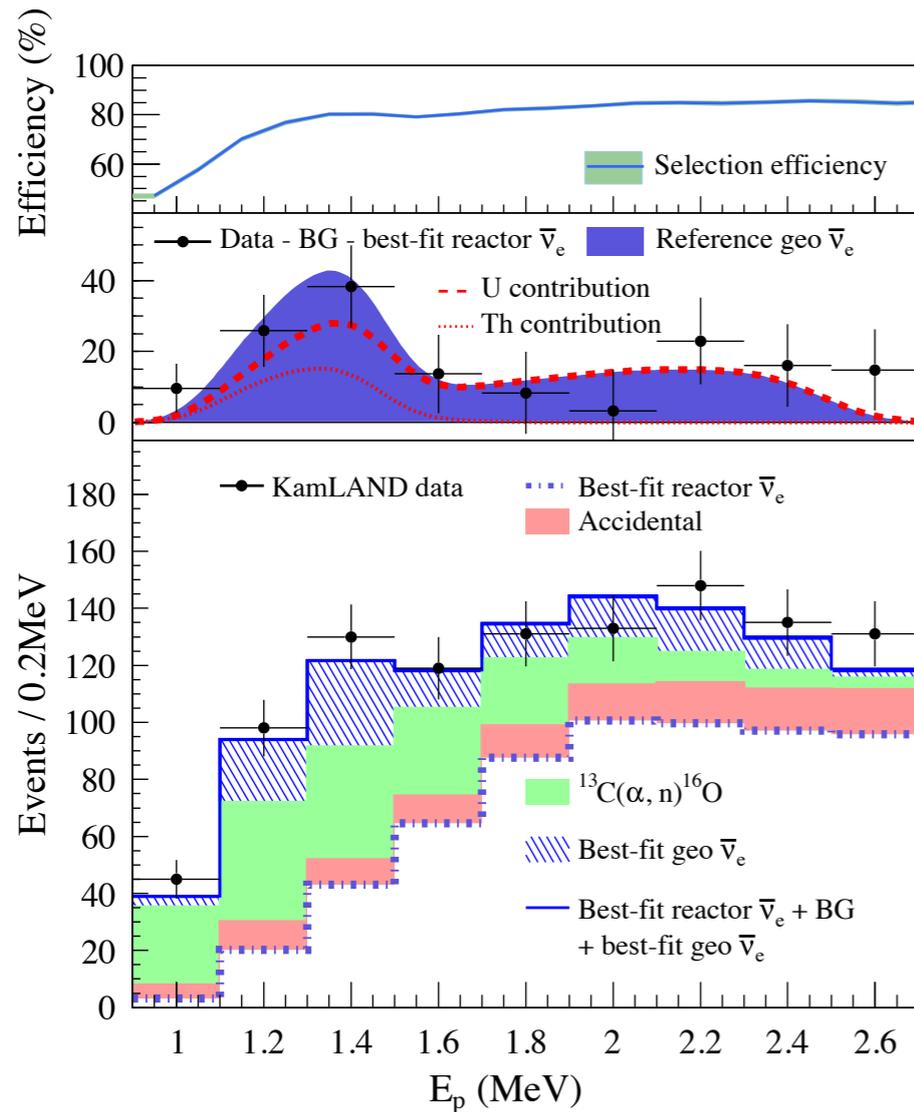
θ_{13} constraint

← insensitive

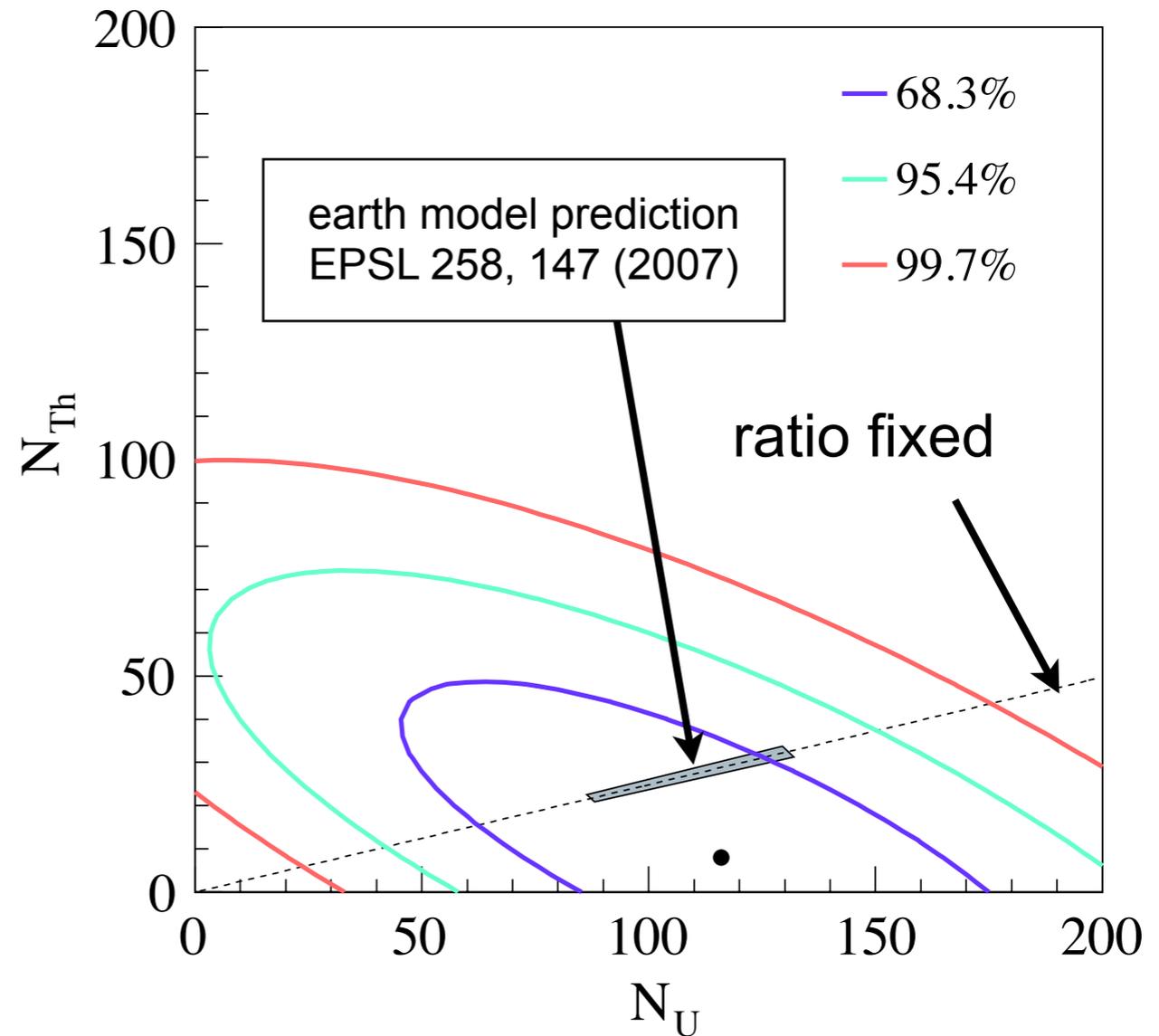
← no strong impact

Survival probability of geo $\bar{\nu}_e = 0.511 \pm 0.014$ (2.7%)

Fit with U/Th components



best-fit $(N_U, N_{Th}) = (116, 8)$



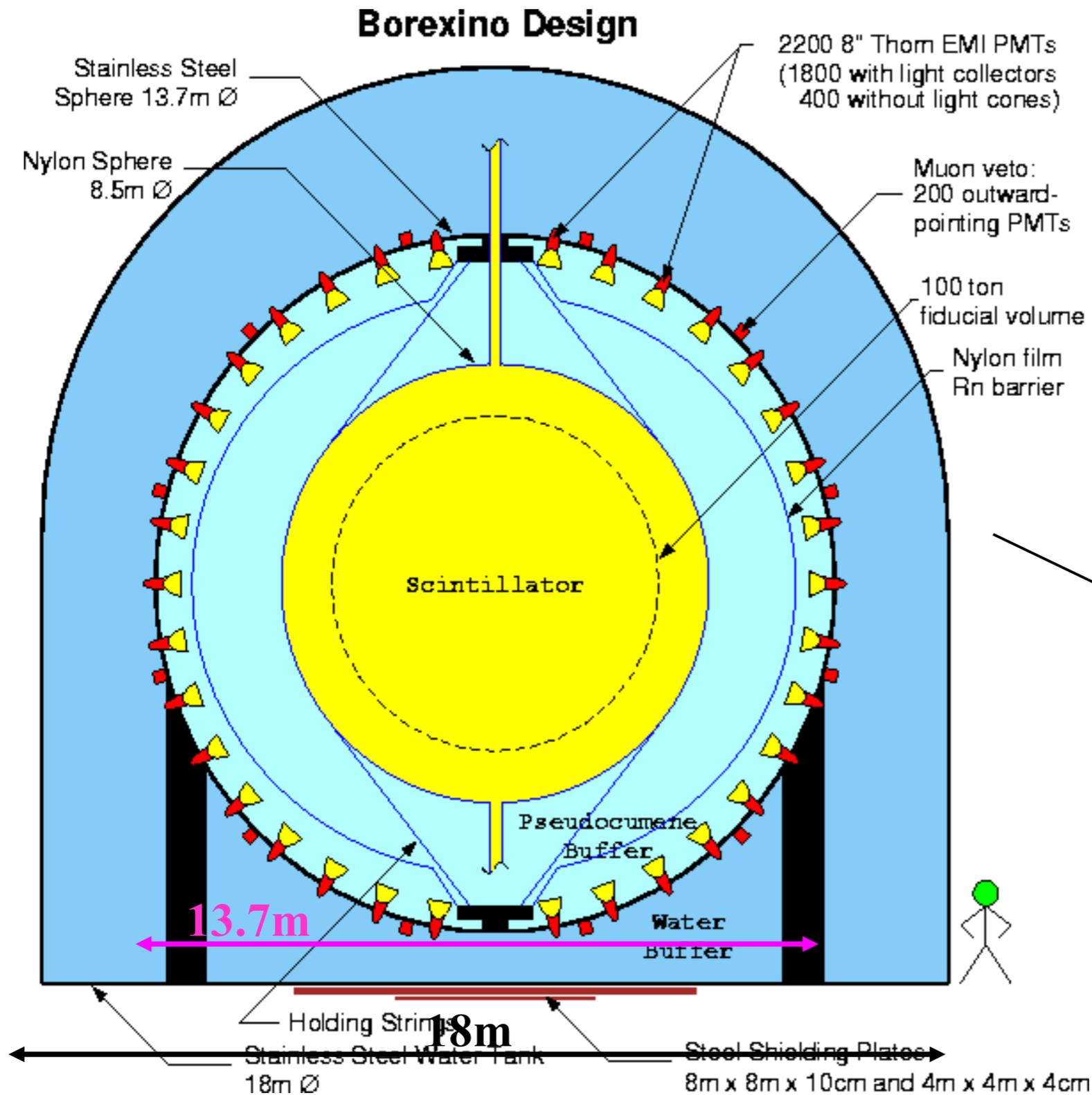
$$N_{\text{geo}} = 116^{+28}_{-27} \text{ events}$$

$$F_{\text{geo}} = 3.4^{+0.8}_{-0.8} \times 10^6 \text{ /cm}^2\text{/sec}$$

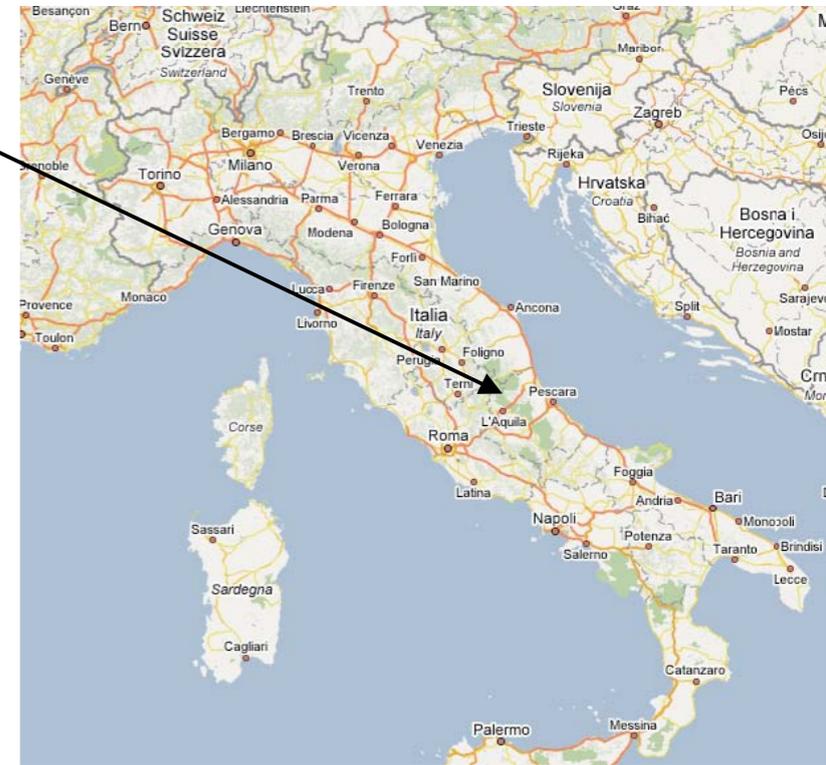
$$(30.7^{+7.5}_{-7.3} \text{ TNU})$$

Borexino

Smirnov, TAUP 2011

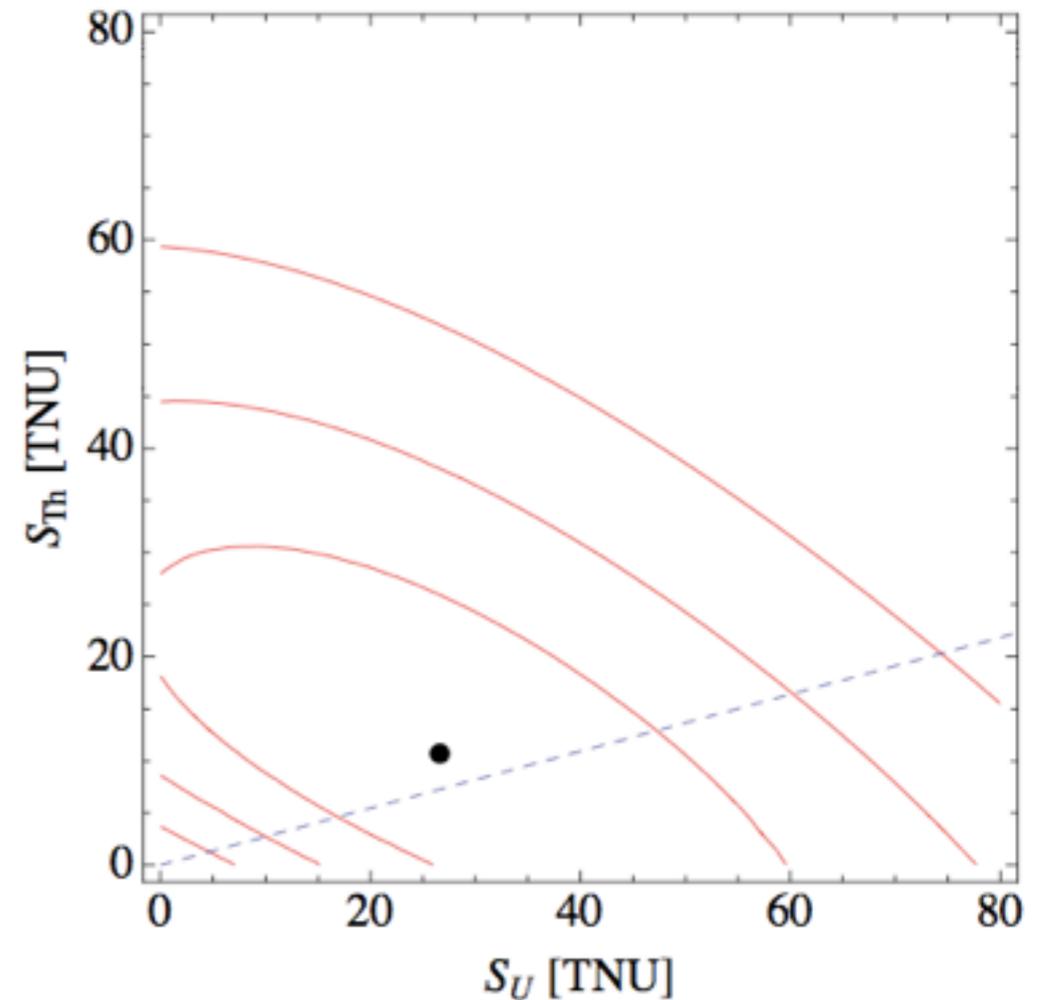
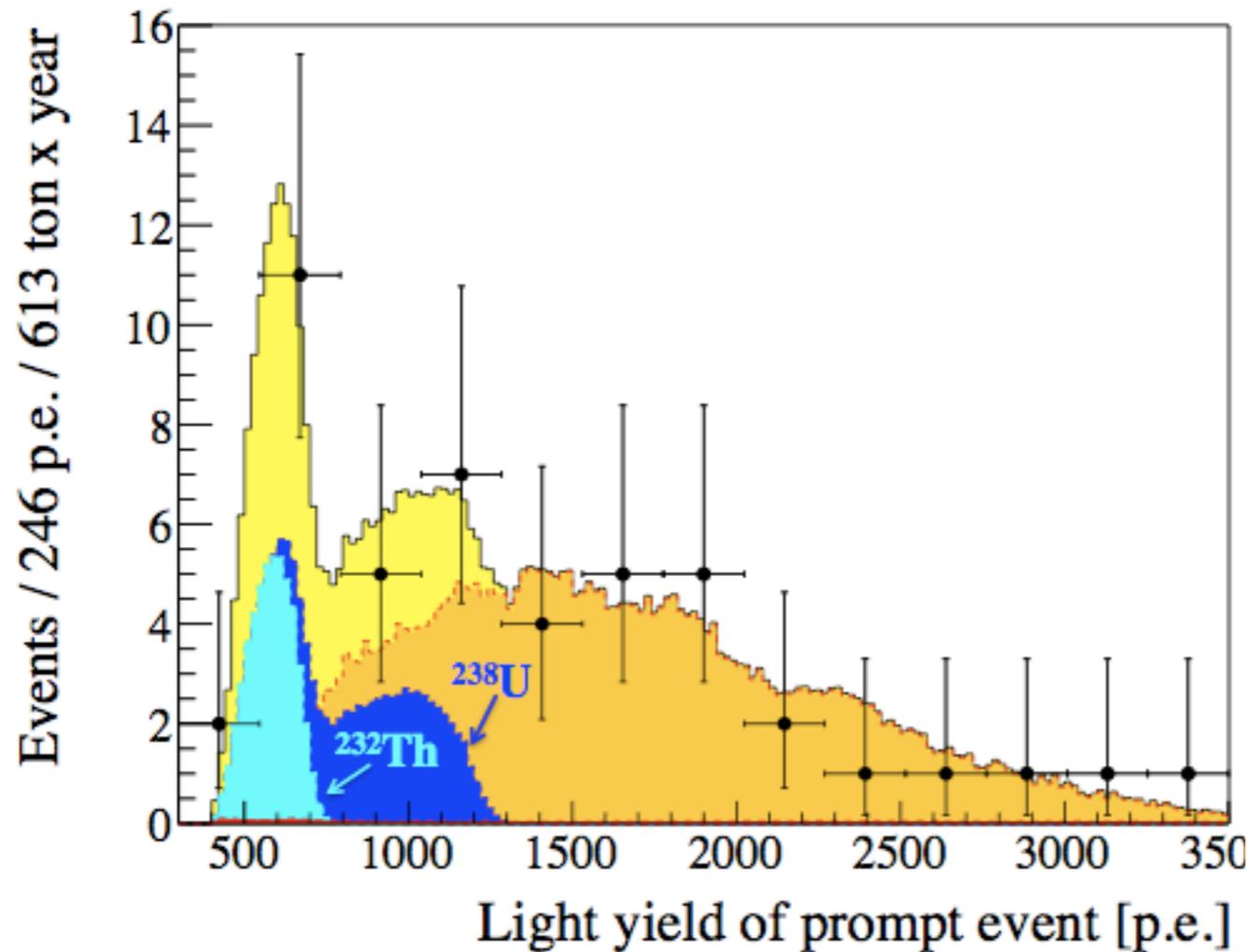


- **278 t of liquid organic scintillator PC + PPO (1.5 g/l)**
- **(ν ,e)-scattering with 200 keV threshold**
- **Outer muon detector**



Fit with U/Th components

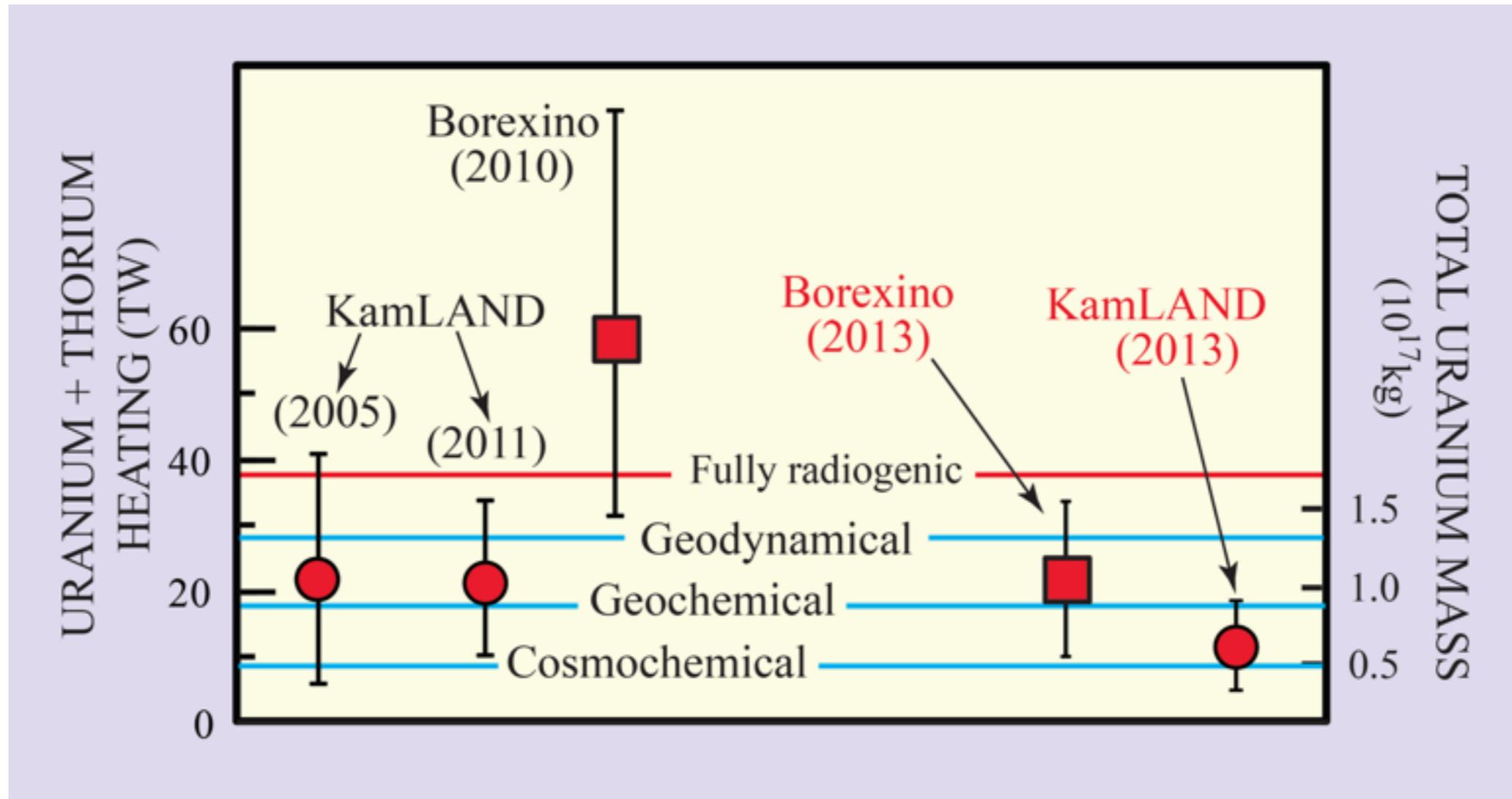
Zavatarelli, Neutrino Geoscience 2013



N_{reactor} Expected with osc.	N_{reactor} Expected no osc.	Others back.	N_{geo} measured	N_{reactor} measured	N_{geo} measured	N_{reactor} measured
events	events	events	events	events	TNU	TNU
33.3 ± 2.4	60.4 ± 4.1	0.70 ± 0.18	14.3 ± 4.4	$31.2_{-6.1}^{+7}$	38.8 ± 12.0	$84.5^{+19.3}_{-16.9}$

Summary of Geo Neutrino Results

McDonough, Neutrino Geoscience 2013



MODELS'

Cosmochemical: 'uses' meteorites - 'O'Neill & Palme' ('08); 'Javoy et al' ('10); 'Warren' ('11)

Geochemical: 'uses' terrestrial rocks - 'McD & Sun' '95; 'Allegre et al' '95; 'Palme O'Neil' '03'

Geodynamical: 'parameterized' convecBon - 'Schubert et al; 'TurcoHe et al; 'Anderson'

Agreement of results at multi-site

Earth Model Comparison

Three classes BSE compositional estimates

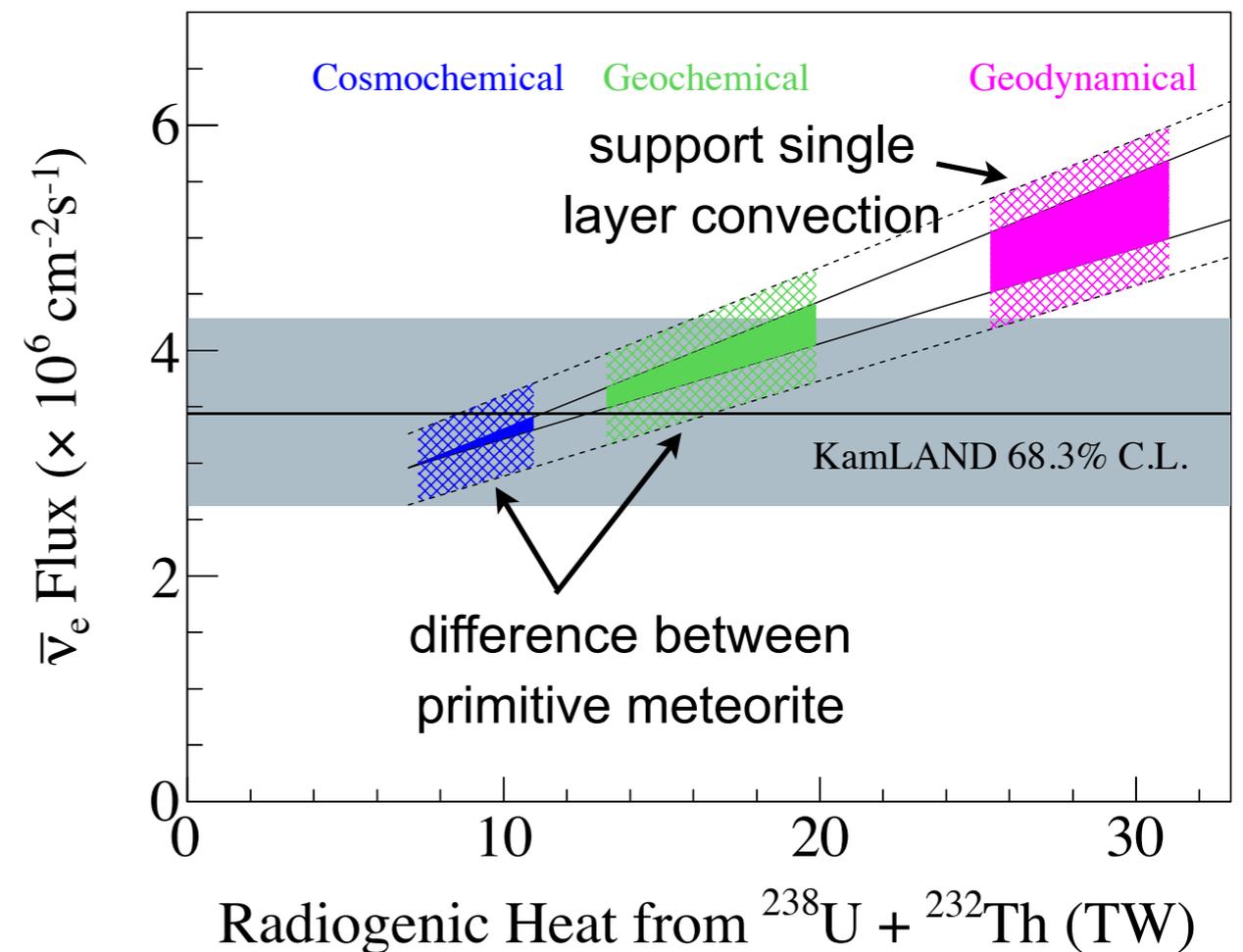
O. Šrámek et al. Earth. Plan. Sci. Letters 361 (2013) 356–366

Model	Cosmochem.	Geochem.	Geodyn.
A_U (ppb)	12 ± 2	20 ± 4	35 ± 4
A_{Th} (ppb)	43 ± 4	80 ± 13	140 ± 14
A_K (ppm)	146 ± 29	280 ± 60	350 ± 35
Th/U	3.5	4.0	4.0
K/U	12000	14000	10000
Tot. Power (TW)	11 ± 2	20 ± 4	33 ± 3
Mantle power (TW)	3.3 ± 2.0	12 ± 4	25 ± 3
Mantle Urey ratio	0.08 ± 0.05	0.3 ± 0.1	0.7 ± 0.1

KamLAND result (2013)

radiogenic **$14.2^{+7.9}_{-5.1}$ TW**

heat low from
Earth's surface **47 ± 2 TW**



Geodynamical prediction with homogeneous hypothesis is disfavored at **89% C.L.**

All composition models are still consistent within $\sim 2\sigma$

Geo-v measurement is in agreement with BSE models

Physicist/Geoscientist Collaborative Work

Neutrino Geoscience 2013 International Workshop (Takayama in Japan)

Nature News in April 2013



<http://www.awa.tohoku.ac.jp/geoscience2013/>

“Neutrino Geoscience” is developing

- Earth model from geophysics and geochemistry
- Multi-site measurement by neutrino detectors



detailed map of neutrino source inside the Earth

International scientific community is organized

A snippet from a Nature News article. It features a 3D diagram of the Earth's interior showing the crust, mantle, and core, with a neutrino detector icon. To the right is a world map with a color scale. The main headline reads "Detectors zero in on Earth's heat". The sub-headline says "Geoneutrinos paint picture of deep-mantle processes." The byline is "BY ALEXANDRA WITTE". The text discusses the use of geoneutrinos to study Earth's internal heat sources and mantle processes. It mentions the KamLAND detector in Japan and the future Super-Kamiokande detector. The article also notes that geoneutrinos are produced by the radioactive decay of elements like uranium and thorium in the mantle. It mentions that the KamLAND detector has detected 116 probable geoneutrinos between December 2007 and August 2012. The article is dated 4 APRIL 2013 | VOL 456 | NATURE | 17.

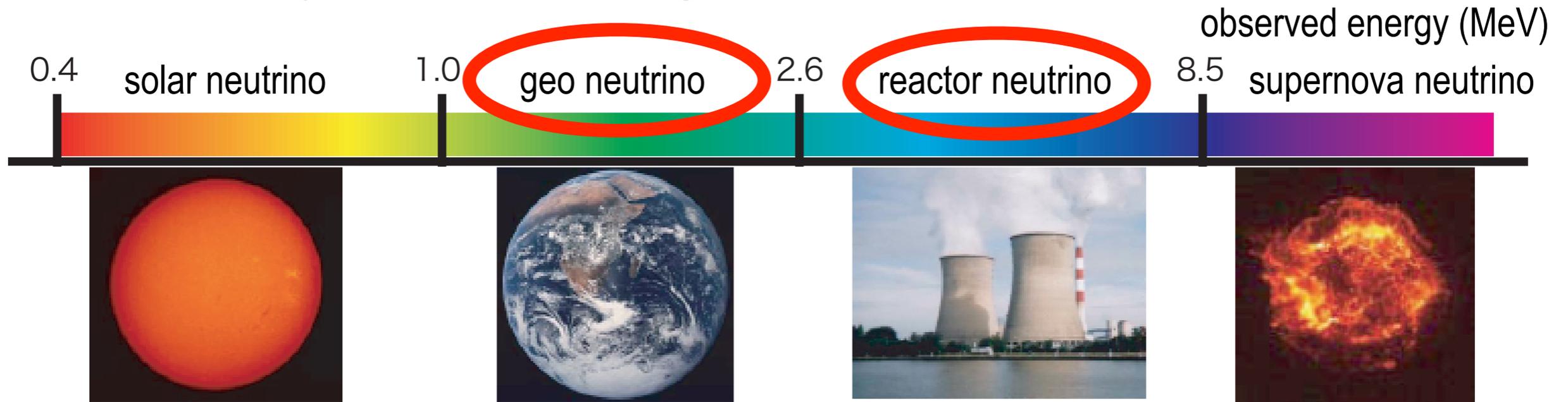
http://www.nature.com/polopoly_fs/1.12707!/menu/main/topColumns/topLeftColumn/pdf/496017a.pdf

Summary

- Multi-site measurements by KamLAND and Borexino have just started.
- Geo neutrino experiments showed
 - Observed flux is still consistent with all BSE models
 - Radiogenic heat contributes only half of Earth's total heat flow → primordial heat still remain
- Tests of primitive meteorite and mantle convection model are the next target.
- Multi-site measurements at geologically different locations will be important for the tests.

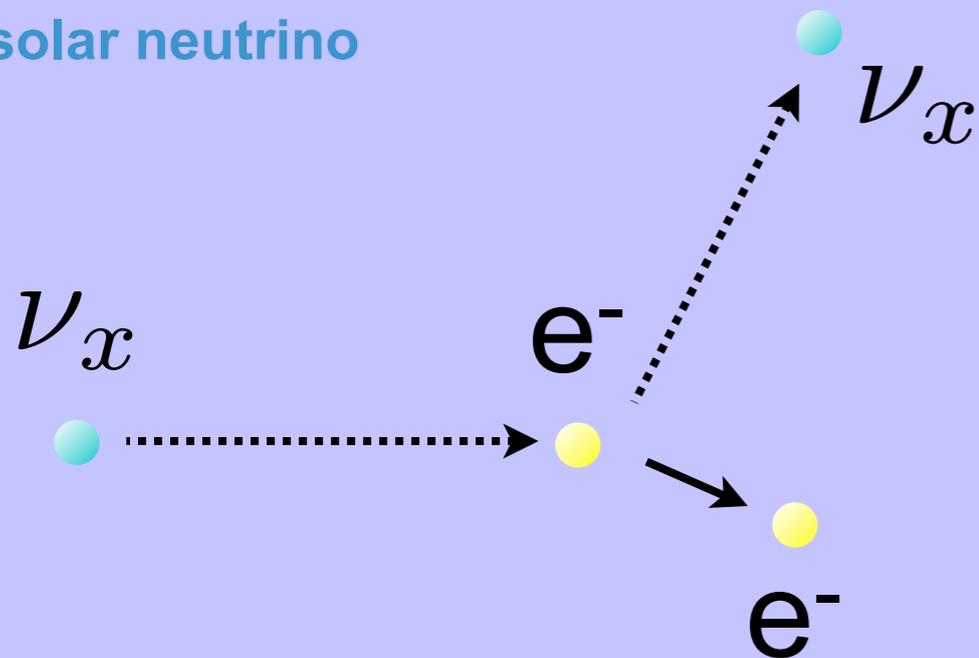
Backup

Physics Target in KamLAND



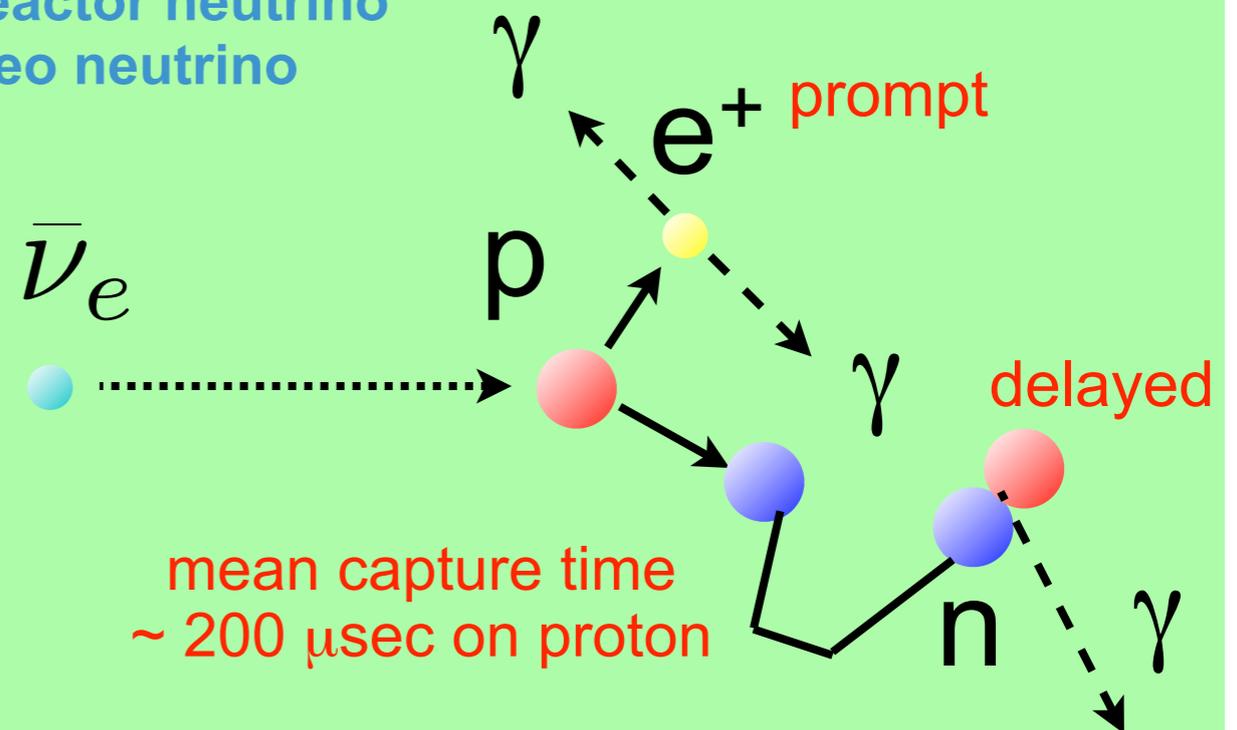
[Mar. 2013 new result \(hep-ex/1303.4667\)](https://arxiv.org/abs/hep-ex/1303.4667)

solar neutrino



neutrino detection by electron scattering

reactor neutrino
geo neutrino



anti-neutrino detection by inverse beta-decay

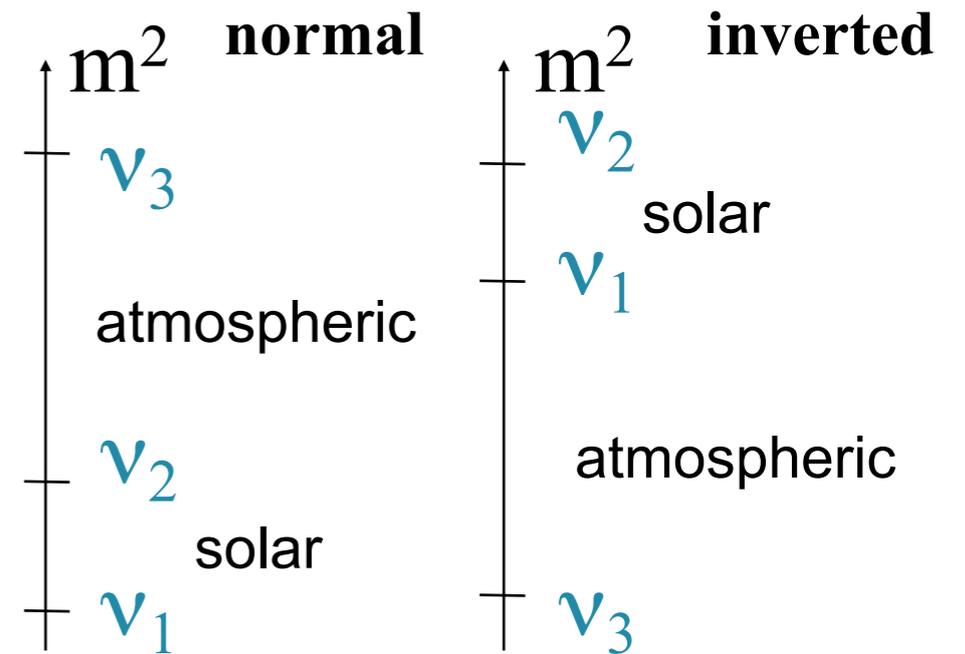
Neutrino Oscillation

MNS (Maki-Nakagawa-Sakata) Matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} V_{e1} & V_{e2} & V_{e3} \\ V_{\mu1} & V_{\mu2} & V_{\mu3} \\ V_{\tau1} & V_{\tau2} & V_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\Delta m_{32}^2$$

$$\Delta m_{32}^2$$



θ_{23}

θ_{13} , CP phase

θ_{12}

Majorana phase

$$V = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} \\ 0 & e^{-i\delta} & 0 \\ -s_{13} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\rho} & 0 & 0 \\ 0 & e^{i\sigma} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

atmospheric

solar

6 parameters : 3 mixing angle, 2 mass difference, 1 CP phase

+ 2 Majorana phase



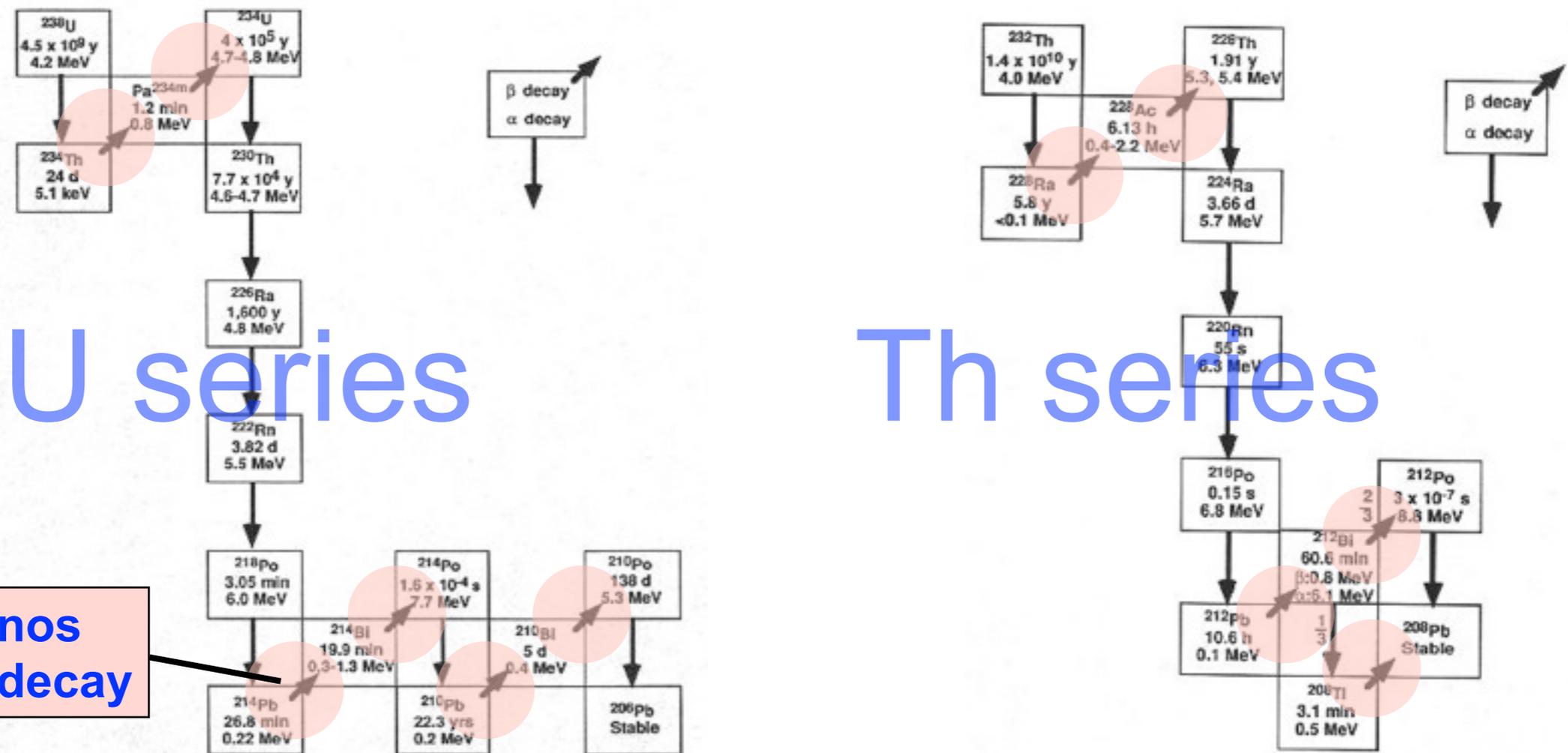
investigated by neutrino oscillation experiments

(solar, atmospheric, accelerator and reactor neutrinos)

Neutrino from the Earth

Heat sources in the Earth

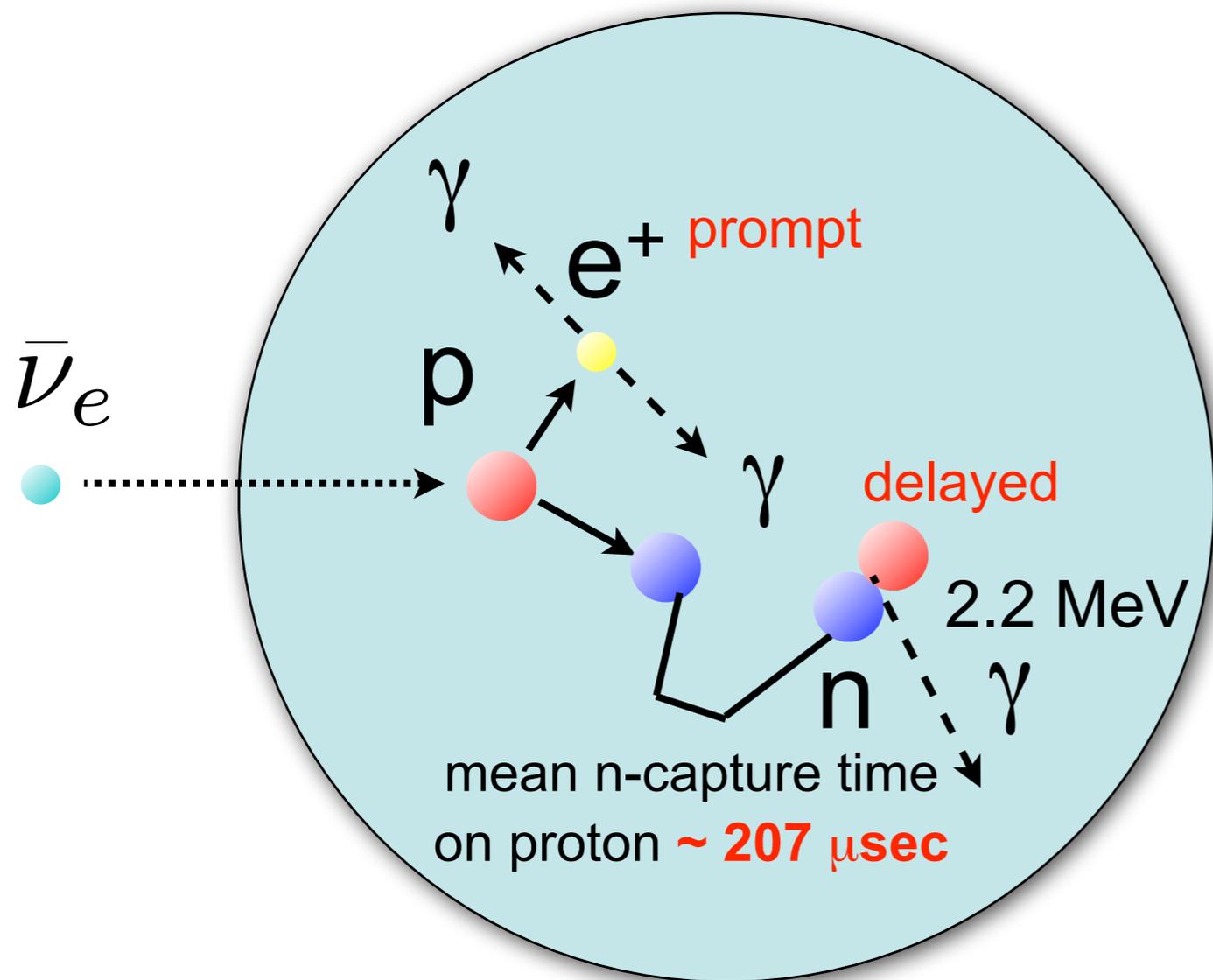
- (1) Radiogenic heat from **U, Th, K decay**
- (2) Release of gravitational energy through accretion or metallic core separation
- (3) Latent heat from the growth of inner core



Radiogenic heat : α -decay or β -decay emitting "anti-neutrinos"

Anti-Neutrino Detection

inverse beta-decay reaction



Tight background rejection
by delayed coincidence

(a) accidental B.G. discrimination

- $0.5 < \Delta T < 1000 \mu\text{s}$
- $\Delta R < 2 \text{ m}$
- $1.8 \text{ MeV} < E_{\text{delayed}} < 2.6 \text{ MeV}$ or
 $4.0 \text{ MeV} < E_{\text{delayed}} < 5.8 \text{ MeV}$
- $0.9 \text{ MeV} < E_{\text{prompt}} < 8.5 \text{ MeV}$
- $R_{\text{prompt}}, R_{\text{delayed}} < 6.0 \text{ m}$
- L-selection from 6 parameters

(b) μ spallation cut

- $\Delta T_{\mu} > 2 \text{ s}$ after showing μ
- $\Delta T_{\mu} > 2 \text{ s}$ or $\Delta L > 3 \text{ m}$ after
non-showering μ ($\Delta Q < 10^6 \text{ p.e.}$)

Reactor and Geo Neutrino Analysis

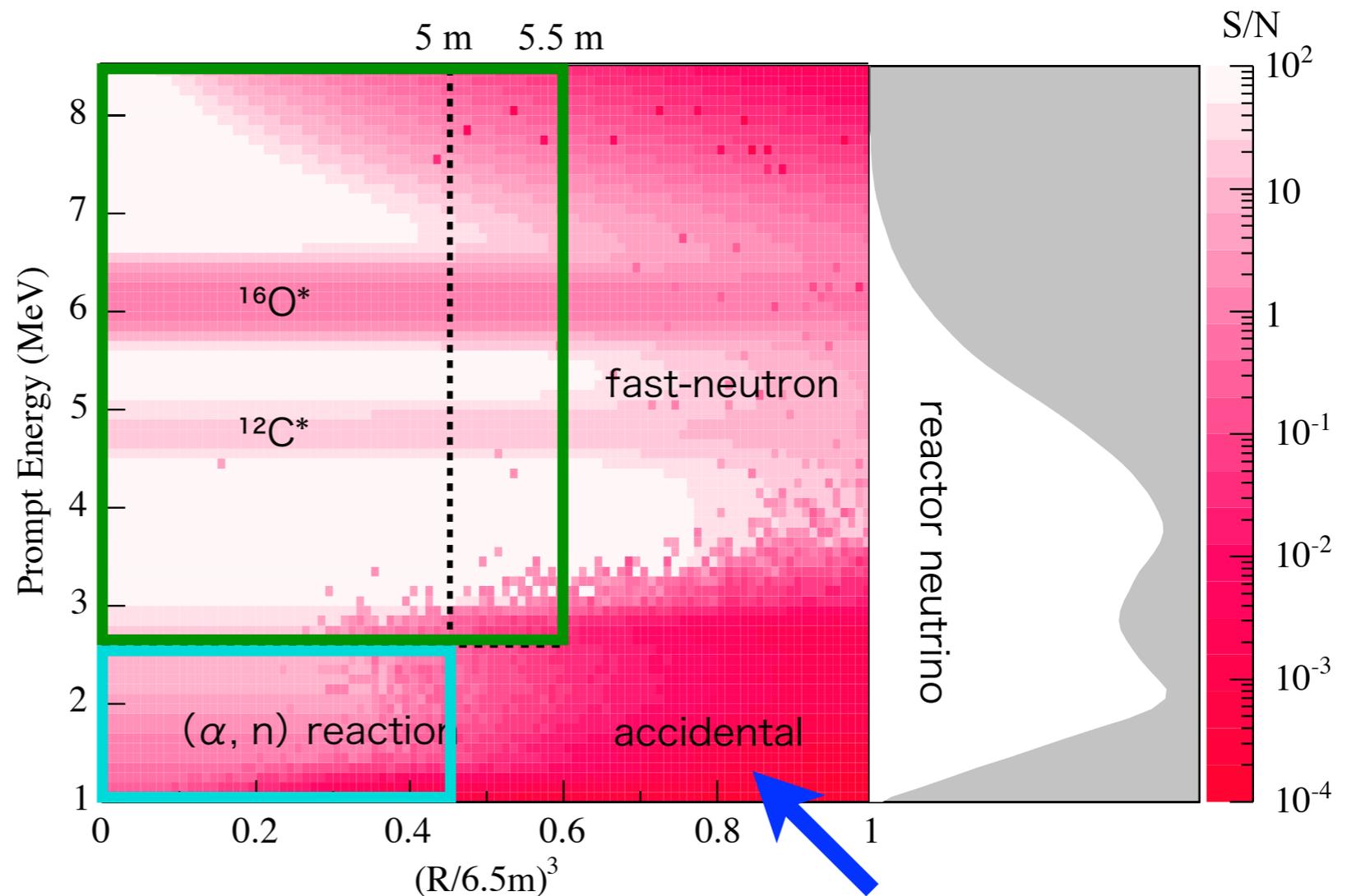
previous result

S / B ratio map (energy v.s. radius)

separated analysis window for reactor and geo neutrinos

reactor neutrino
(2.6 - 8.5 MeV, R 5.5 m)

geo neutrino
(0.9 - 2.6 MeV, R 5.0 m)



large accidental B.G.
caused by external γ -rays

Analysis improvement

- (1) efficient **accidental** background rejection
- (2) combined analysis of **reactor** and **geo** neutrinos

Anti-Neutrino Event Selection

(a) Accidental B.G. discrimination

discriminator based on 5 parameters (E_d , ΔR , ΔT , R_p , R_d)

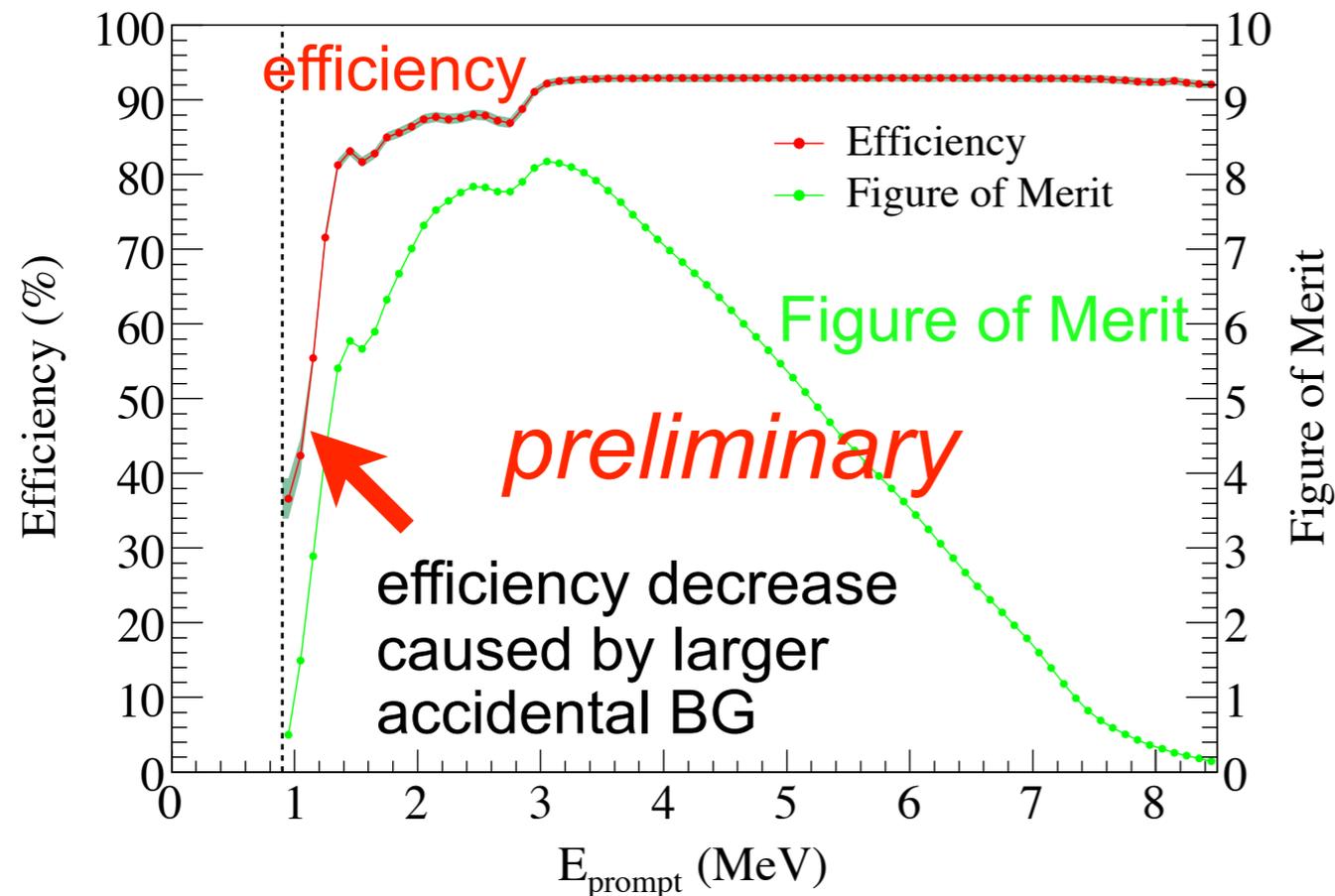
$$L_{\text{ratio}} = \frac{f_{\bar{\nu}}}{f_{\bar{\nu}} + f_{\text{accidental}}} \quad f : \text{PDF}$$

Selection : Maximize "Figure of Merit" $\frac{S}{\sqrt{S + B_{\text{accidental}}}}$

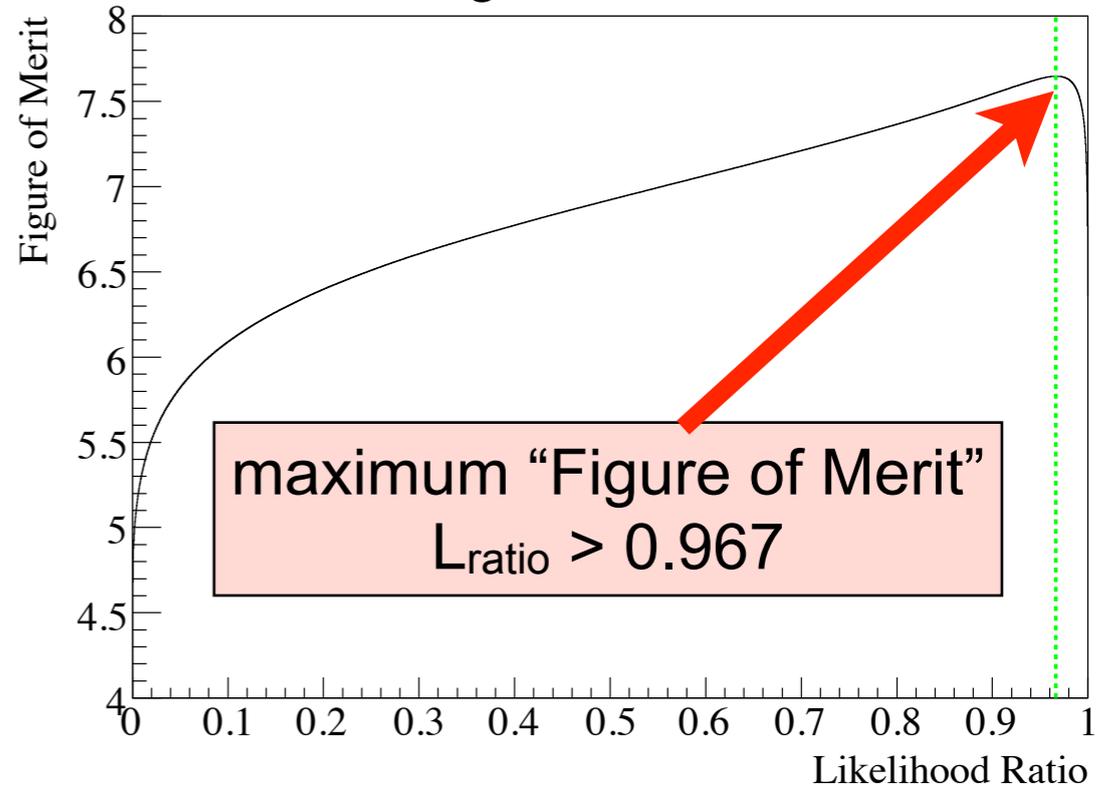
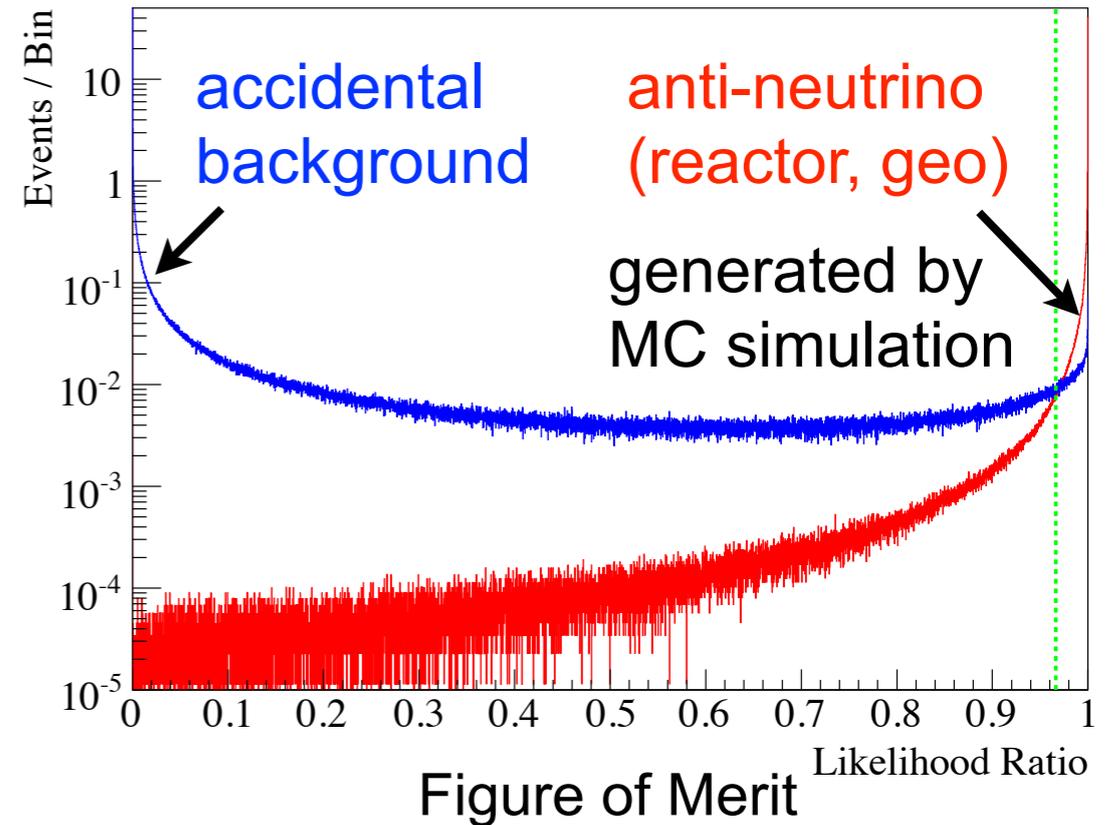
(b) μ spallation cut

- $\Delta T_{\mu} > 2$ s after showing μ ($\Delta Q > 10^6$ p.e.)
- $\Delta T_{\mu} > 2$ s or $\Delta L > 3$ m after non-showering μ

Detection efficiency



$2.2 < E_{\text{prompt}} < 2.3$ MeV

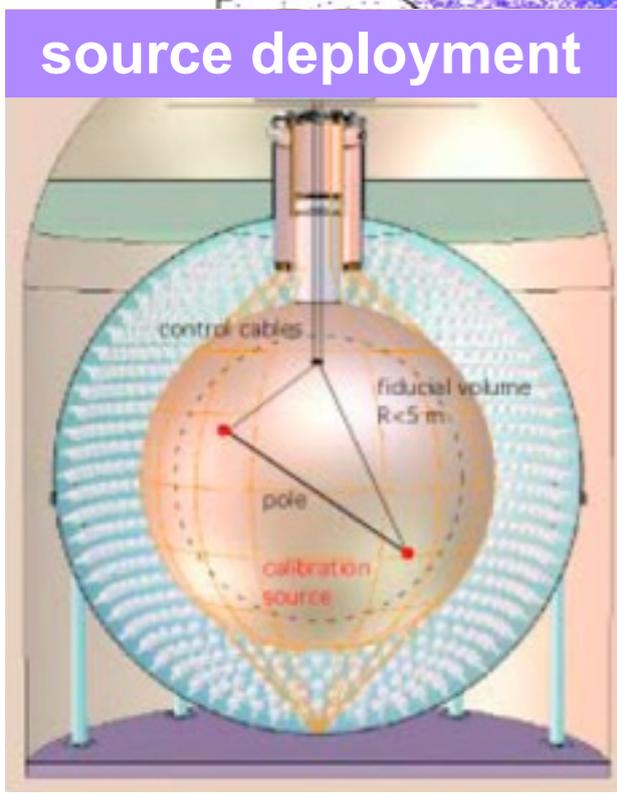
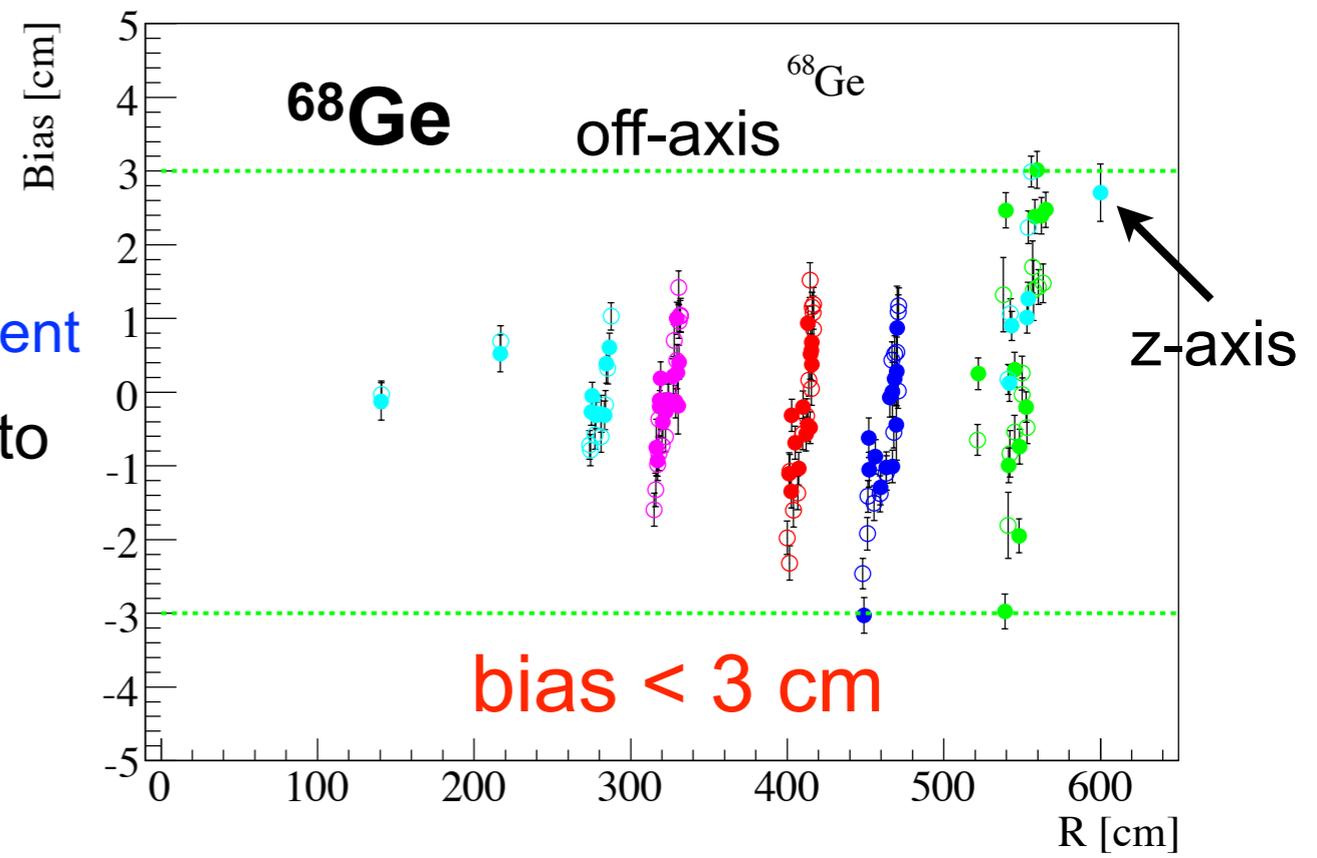
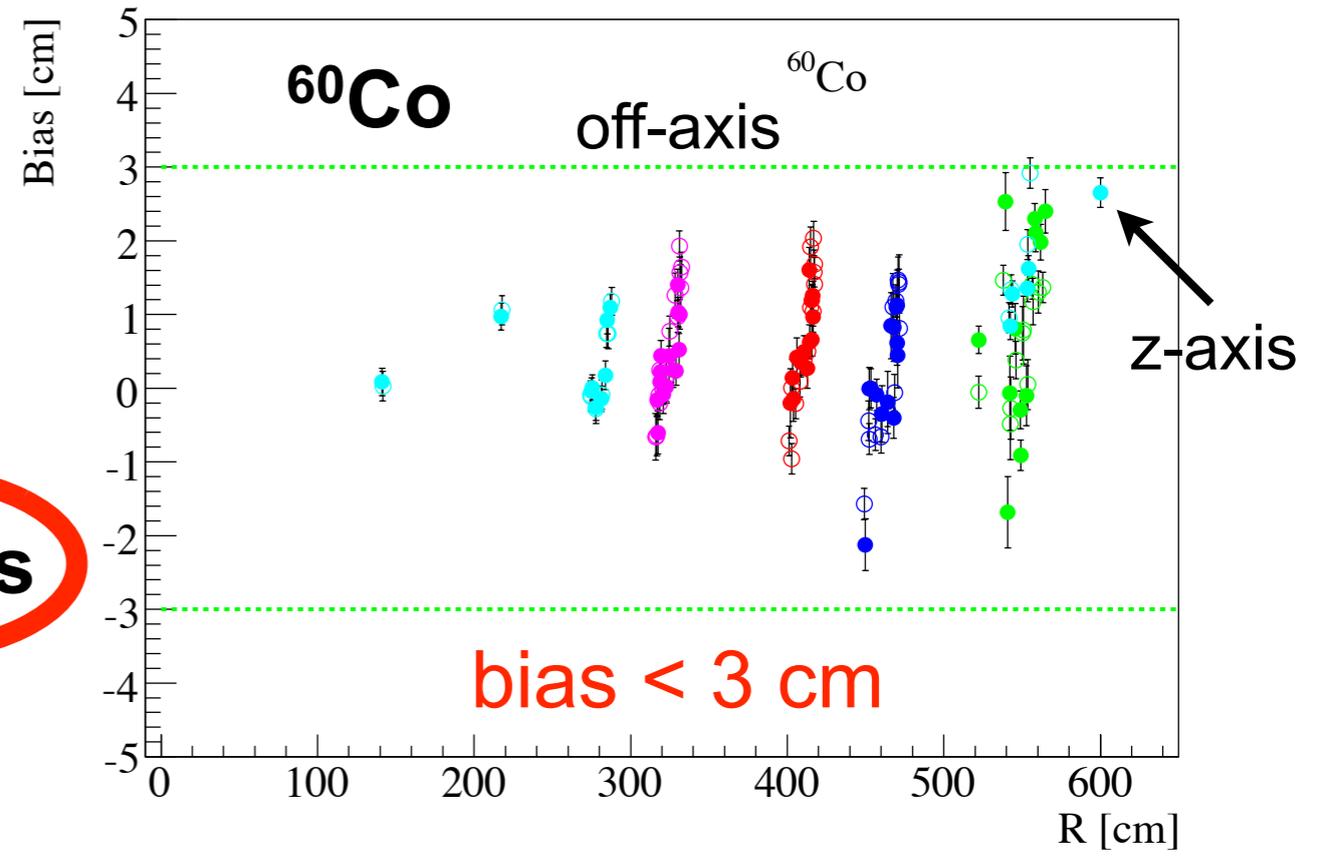
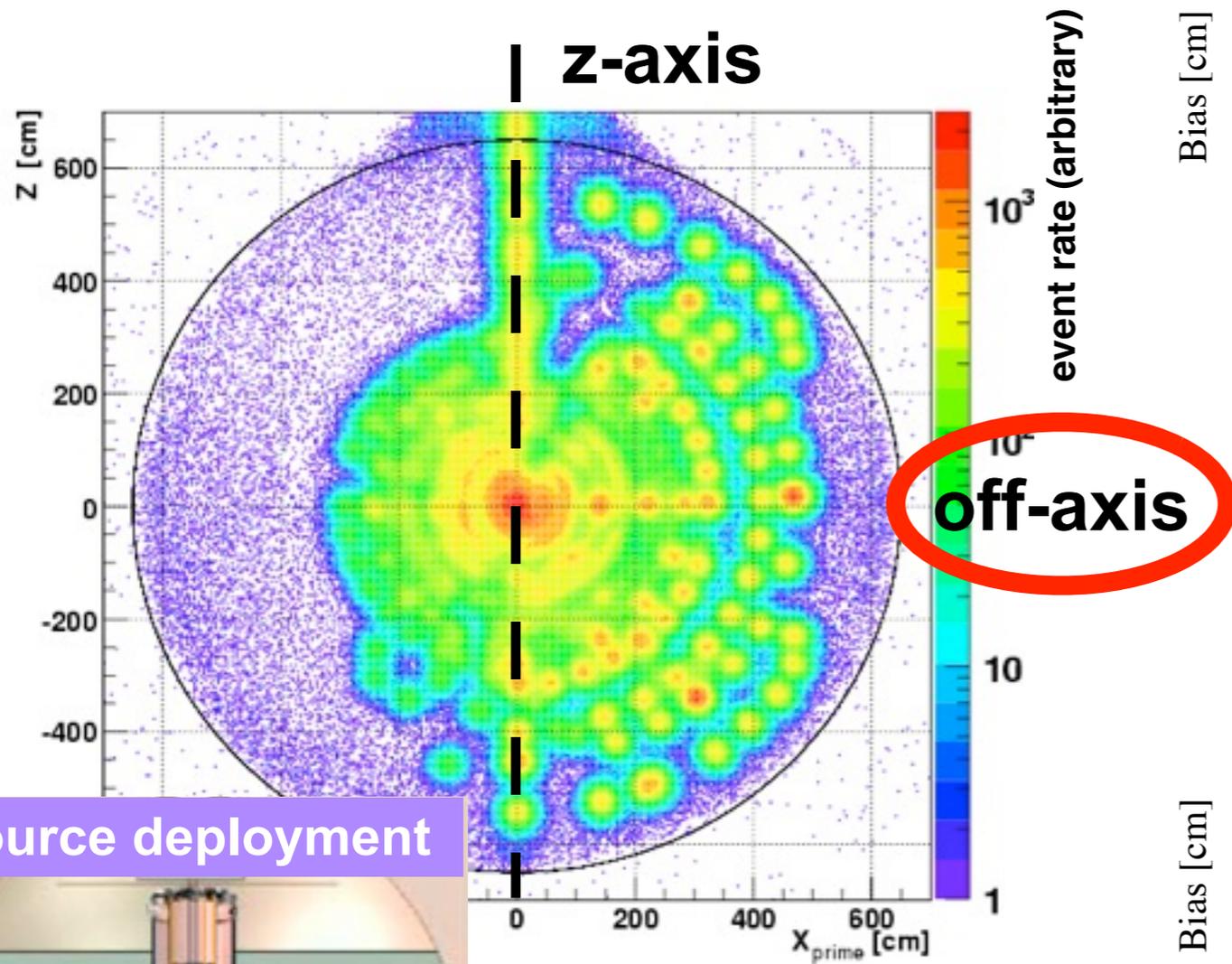


Systematic Uncertainty

before / after purification

	Detector-related (%)		Reactor-related (%)	
Δm_{21}^2	Energy scale	1.8 / 1.8	$\bar{\nu}_e$ -spectra [21]	0.6 / 0.6
Rate	Fiducial volume	1.8 / 2.5	$\bar{\nu}_e$ -spectra	1.4 / 1.4
	Energy scale	1.1 / 1.3	Reactor power	2.1 / 2.1
	$L_{cut}(E_p)$ eff.	0.7 / 0.8	Fuel composition	1.0 / 1.0
	Cross section	0.2 / 0.2	Long-lived nuclei	0.3 / 0.4
	Total	2.3 / 3.0	Total	2.7 / 2.8

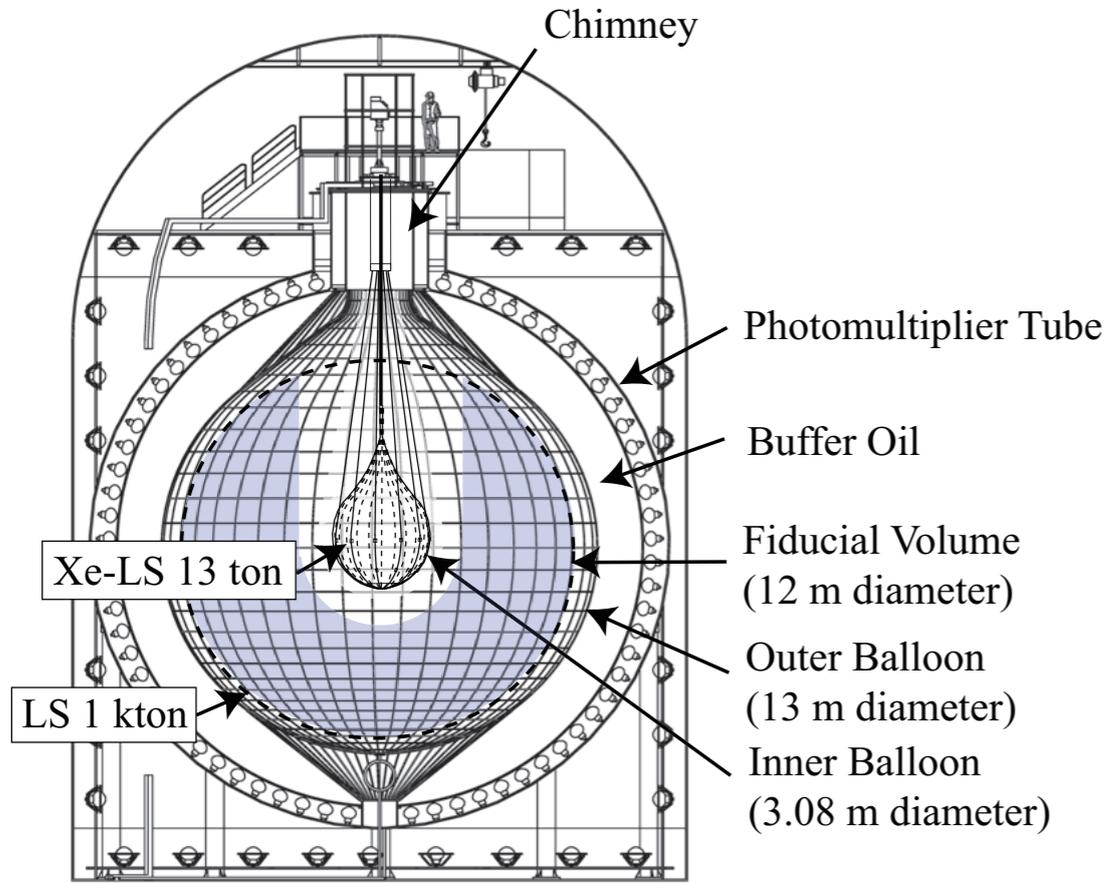
Full Volume Calibration



“4pi calibration” system for the off-axis source deployment

bias < 3 cm corresponds to 1.8% volume uncertainty

cross-checked by ¹²B/¹²N uniformity



- Vertex cut conditions

To minimize accidental coincidences, we apply **Xe LS cut** for KamLAND-Zen Phase in $R < 6.0\text{m}$ fiducial volume.

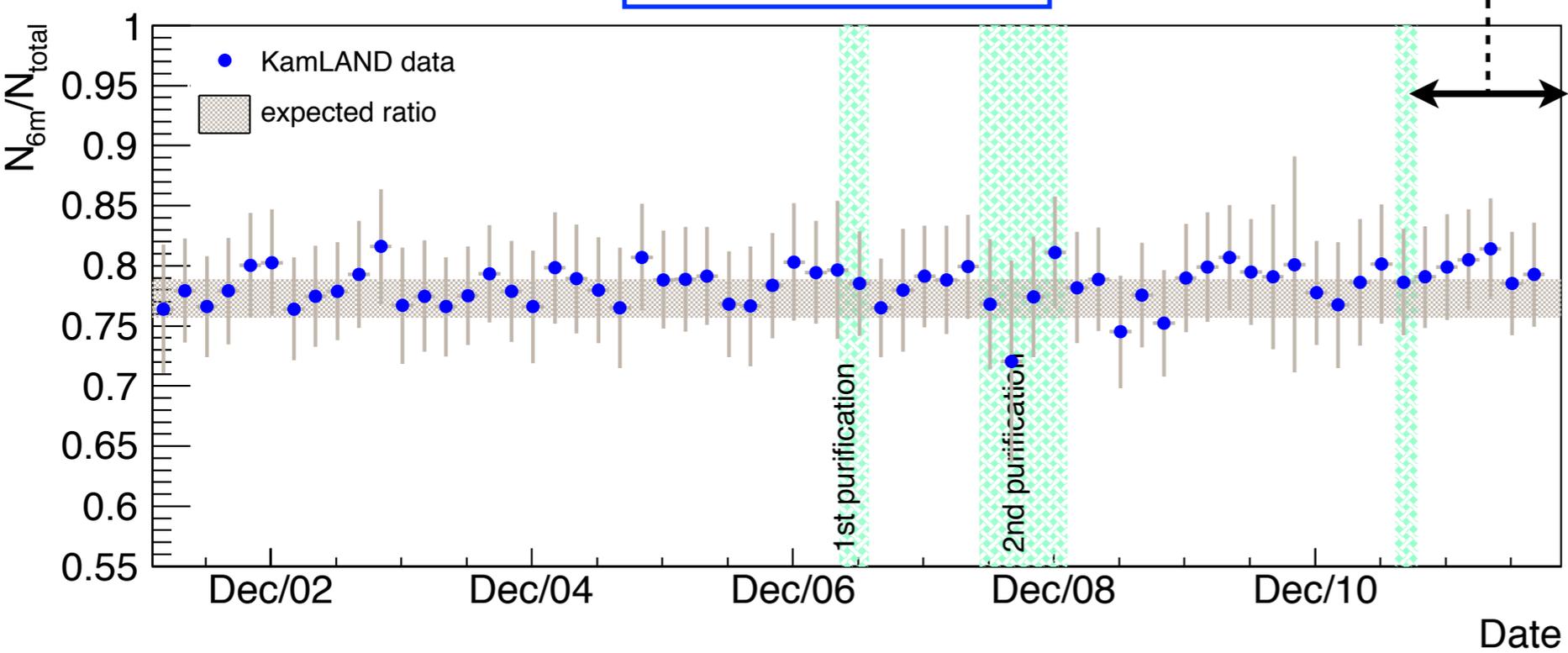
$R > 2.5\text{m}$, cylinder cut ($\rho > 2.5\text{m}$, $Z > 0$)
(cut out volume 16.6% of $R < 6\text{m}$)

- Data stability of KamLAND region

^{12}B $N_{6\text{m}}/N_{\text{total}}$ vs time

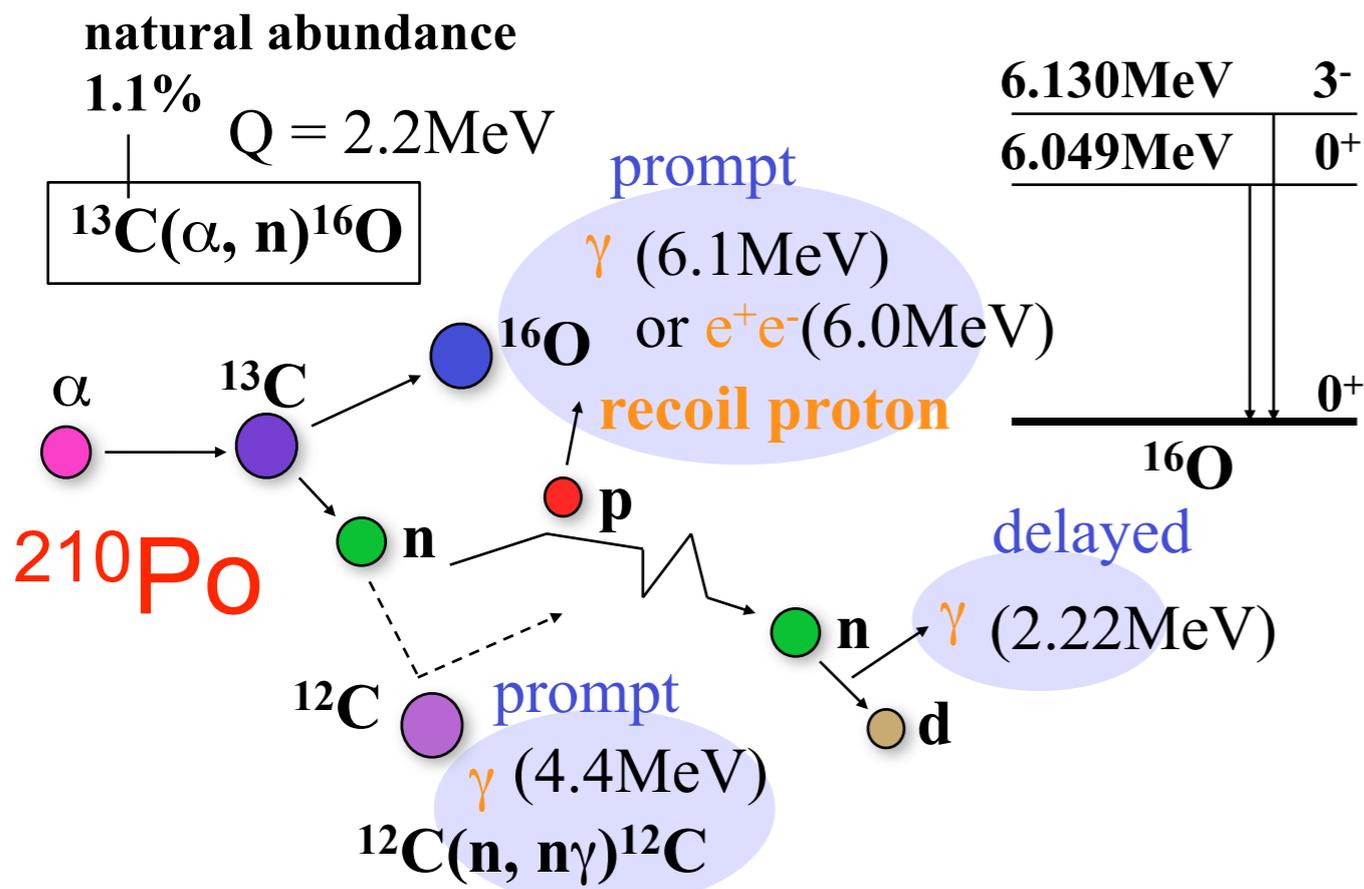
with Xe LS Cut

KamLAND-Zen Phase

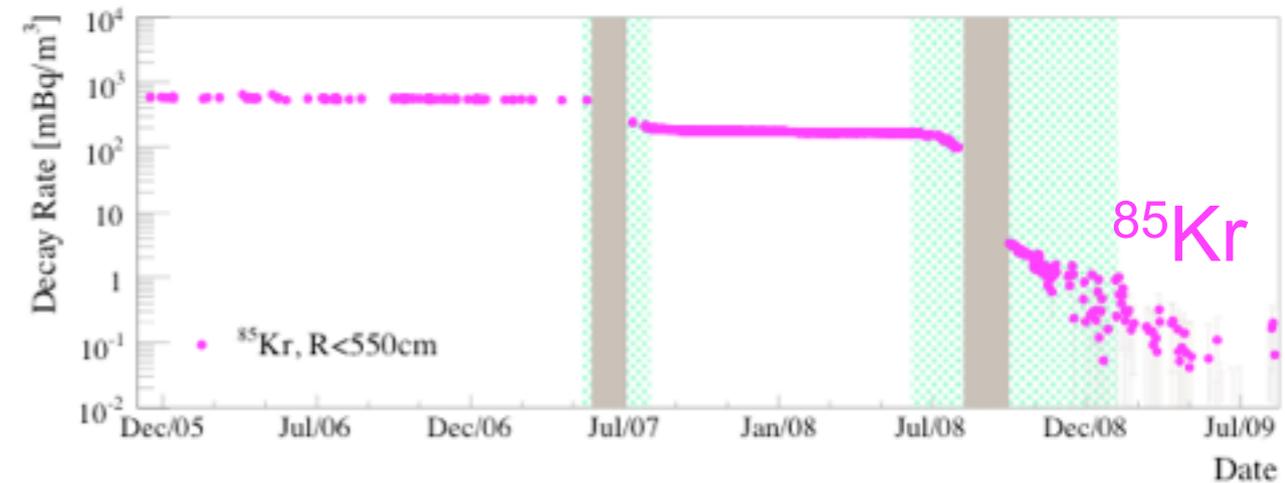
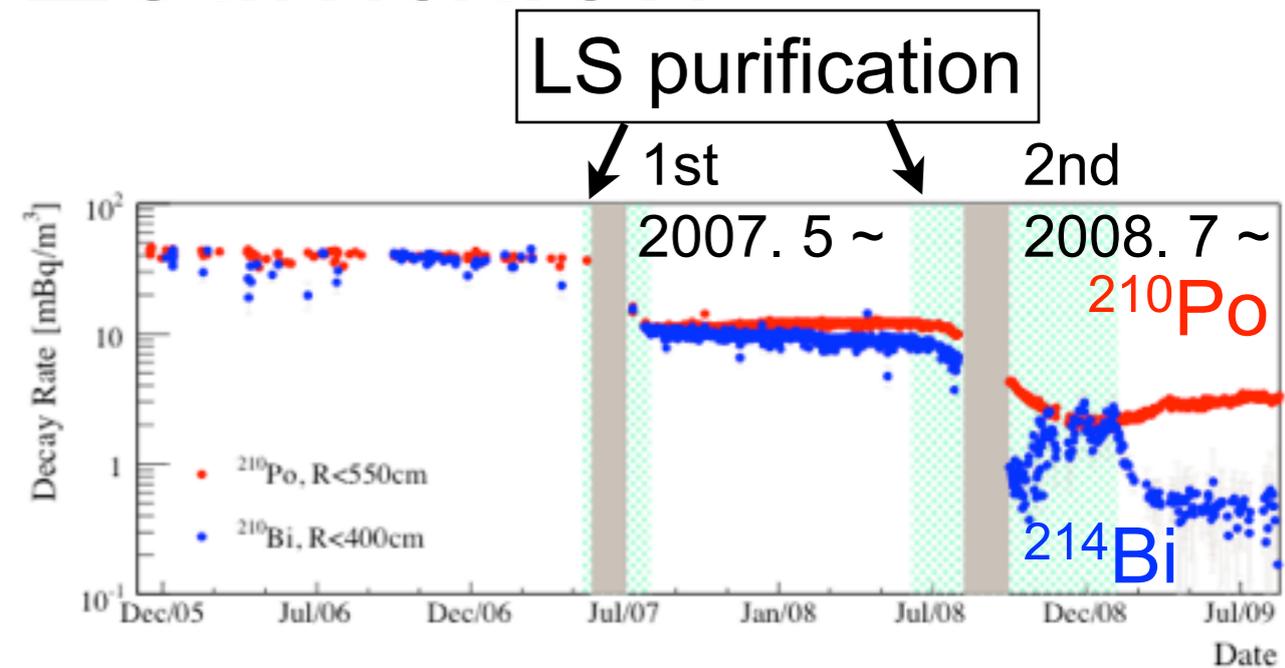
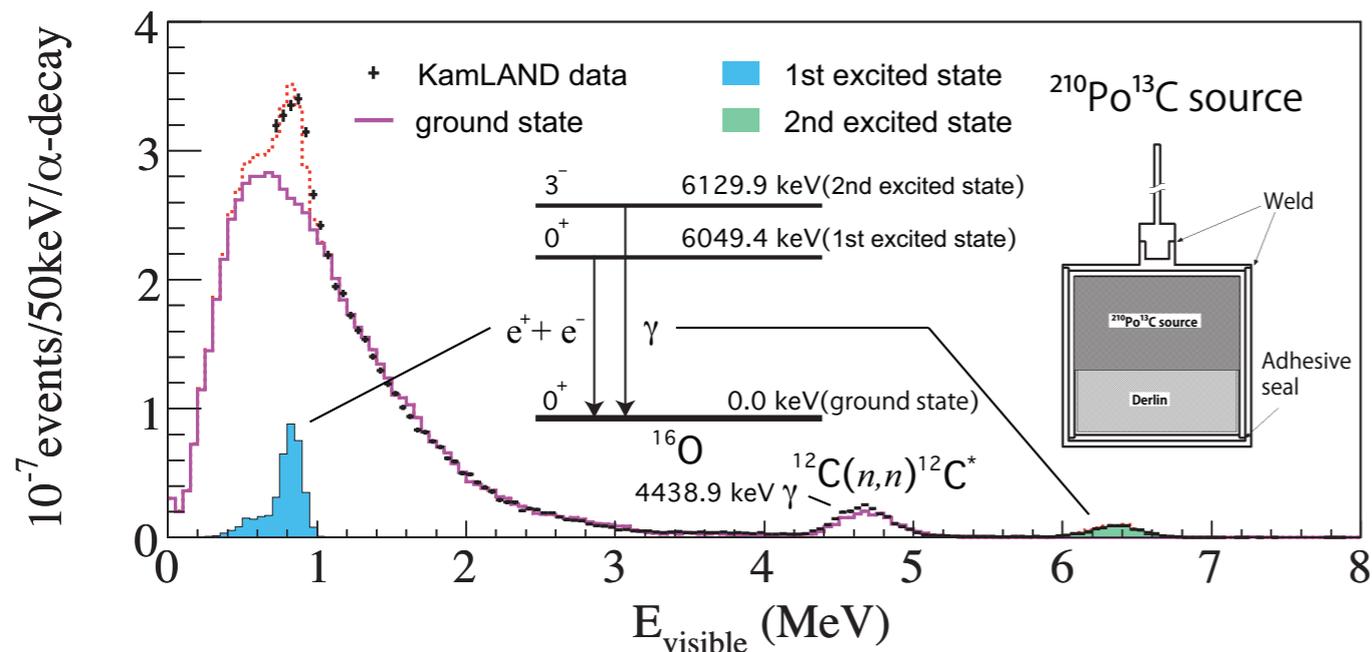


- Event rate has been stable.
- difference before and after purification : 2.5%

Background Estimation



in-situ calibration with $^{210}\text{Po}^{13}\text{C}$



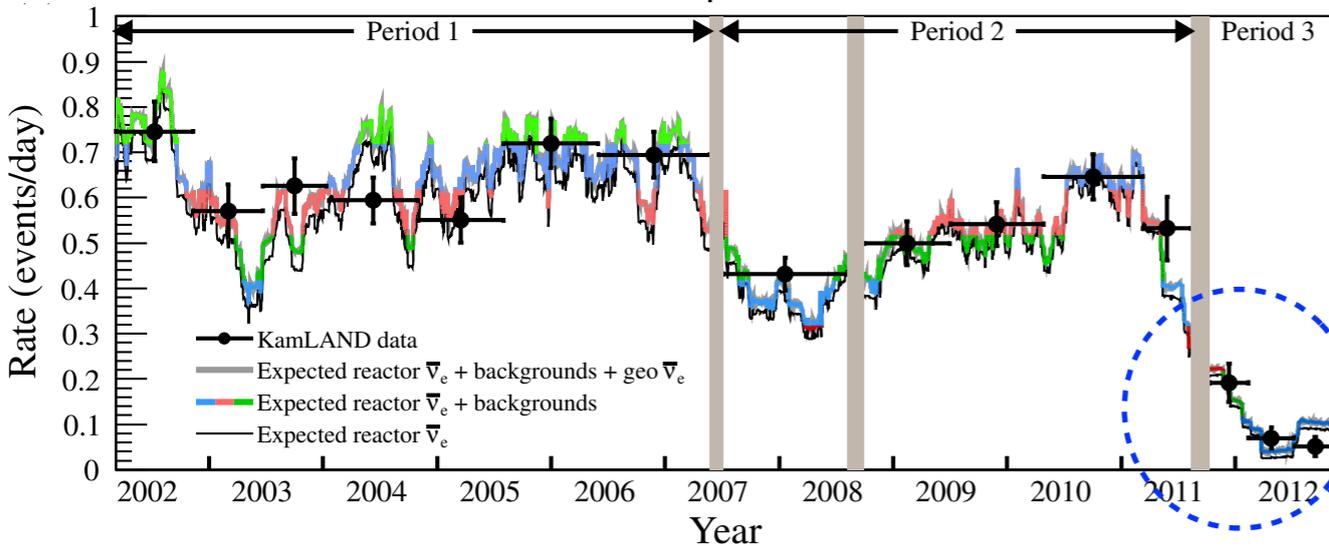
(1) dominant BG source (α, n) has been reduced by down to $\sim 1/20$

(2) determination of the cross section is improved by in-situ calibration

uncertainty: 11% for ground state

Correlation Plot

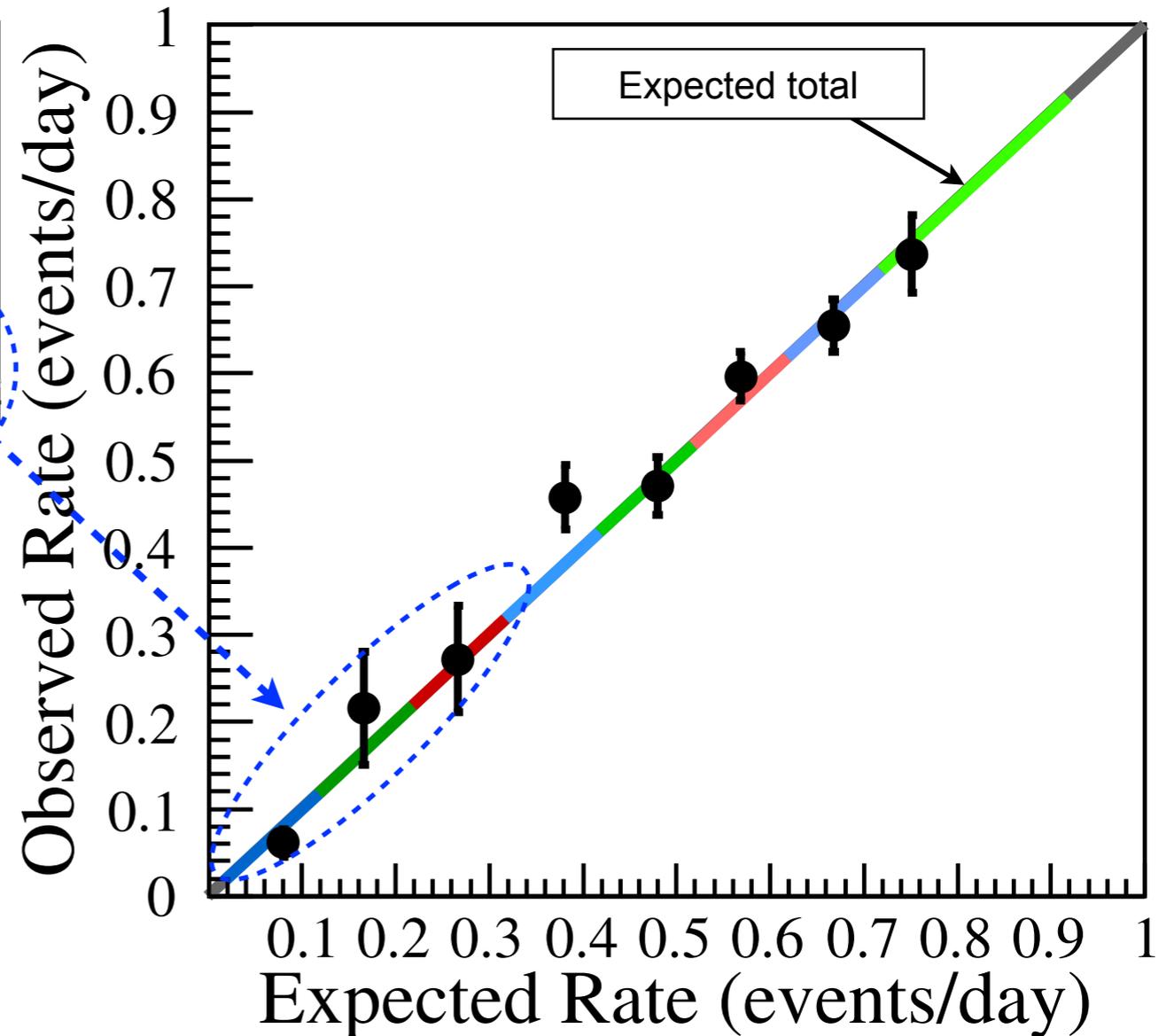
$2.6 < E_p < 8.5 \text{ MeV}$



provide good data to
confirm our background



“Rate + Shape + Time” analysis



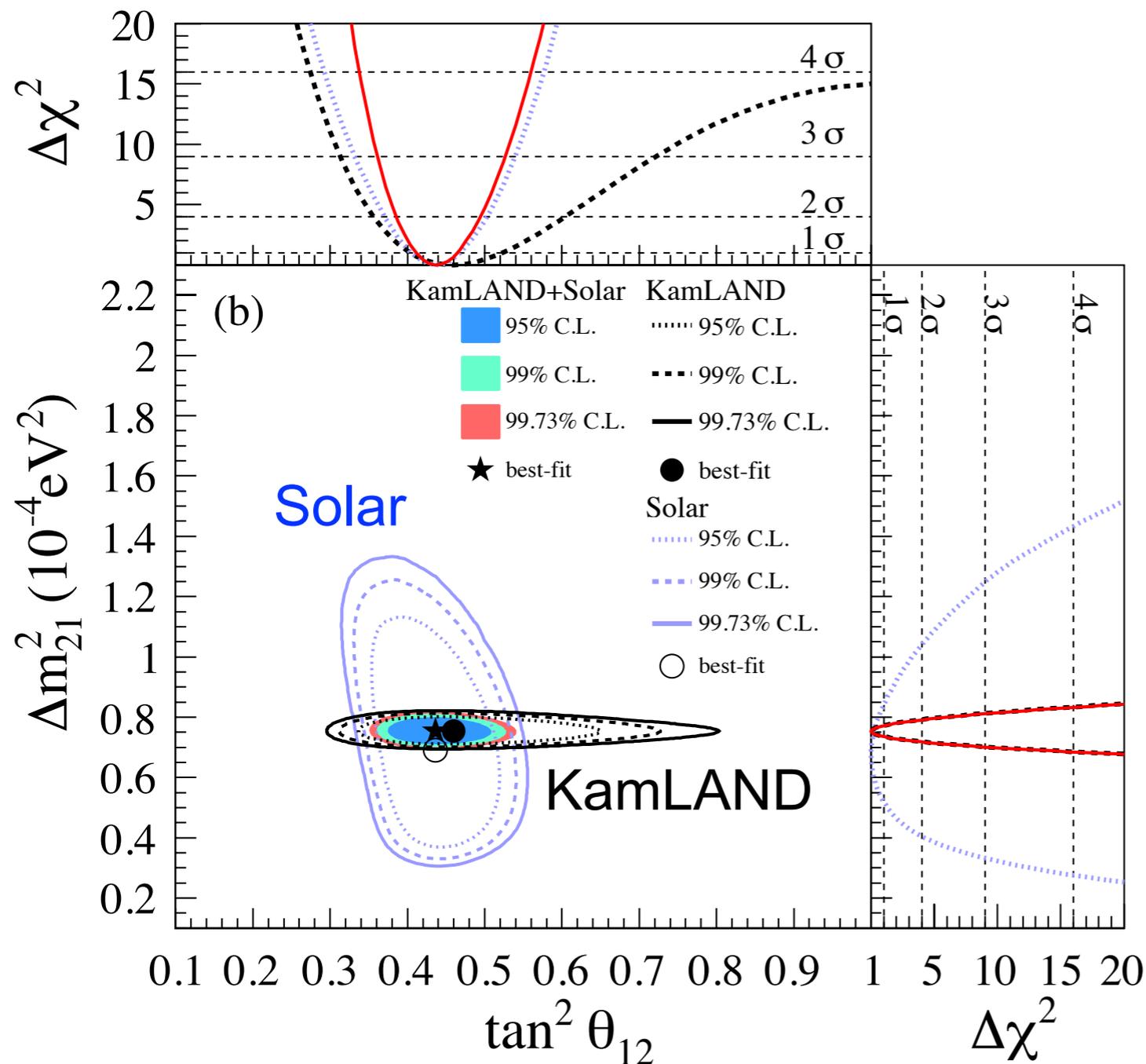
$$\begin{aligned} \chi^2 = & \chi_{\text{rate}}^2(\theta_{12}, \theta_{13}, \Delta m_{21}^2, N_{\text{BG}1 \rightarrow 5}, N_{\text{U,Th}}^{\text{geo}}, \alpha_{1 \rightarrow 4}) \\ & - 2 \ln L_{\text{shape}}(\theta_{12}, \theta_{13}, \Delta m_{21}^2, N_{\text{BG}1 \rightarrow 5}, N_{\text{U,Th}}^{\text{geo}}, \alpha_{1 \rightarrow 4}) \\ & + \chi_{\text{BG}}^2(N_{\text{BG}1 \rightarrow 5}) + \chi_{\text{sys}}^2(\alpha_{1 \rightarrow 4}) \\ & + \chi_{\text{osci}}^2(\theta_{12}, \theta_{13}, \Delta m_{21}^2) . \end{aligned}$$

time dependent

3-Flavor Oscillation Parameters

θ_{12} : **Solar** constraint is dominant

Δm^2 : **KamLAND** constraint is dominant



survival probability

$$P_{ee}^{3\nu} = \cos^4 \theta_{13} P_{e'e'}^{2\nu} + \sin^4 \theta_{13}$$

electron density

matter effect

$$N_{e'} \rightarrow N_e \cos^2 \theta_{13}$$

atmospheric oscillation length is completely averaged out

~~Δm_{31}^2~~

solar + KamLAND + θ_{13} experiments

$$\Delta m_{21}^2 = 7.53_{-0.18}^{+0.18} \times 10^{-5} \text{ eV}^2$$

$$\tan^2 \theta_{12} = 0.436_{-0.025}^{+0.029}$$

$$\sin^2 \theta_{13} = 0.023_{-0.002}^{+0.002}$$

Reference Earth Model

UCC U : 2.8 ppm / Th : 10.7 ppm

MCC U : 1.6 ppm / Th : 6.1 ppm

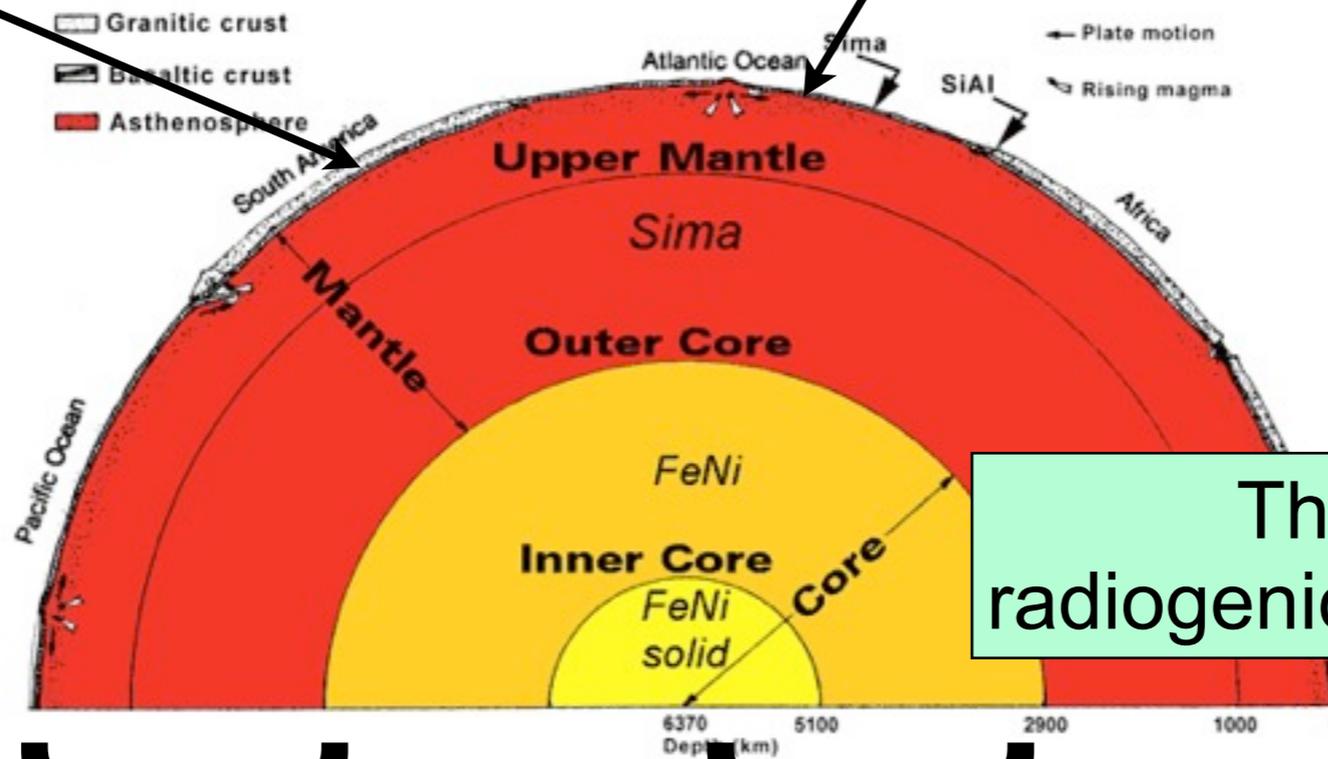
LCC U : 0.2 ppm / Th : 1.2 ppm

Rudnick et al. (1995)

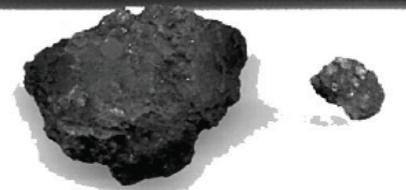
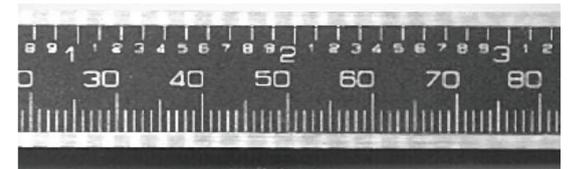
continental crust

oceanic crust

U : 0.10 ppm / Th : 0.22 ppm



chondrite meteorite



Th/U ~ 3.9
radiogenic heat ~ 16 TW

mantle

core

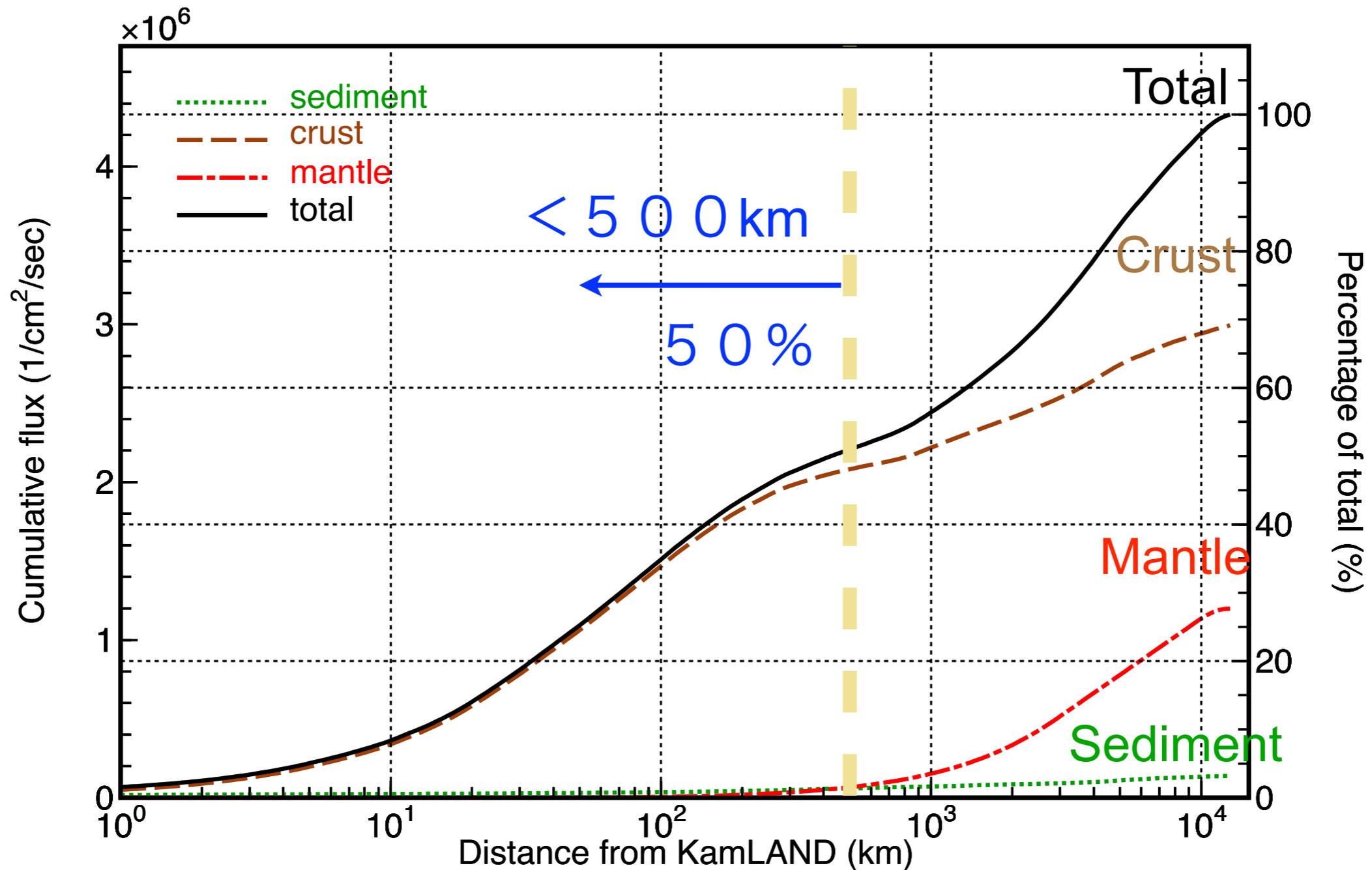
U : 0.012 ppm / Th : 0.048 ppm

U : 0 ppm / Th : 0 ppm

no U/Th in core

Mantle = BSE (Primitive Mantle) - Crust

Distance and Cumulative Flux

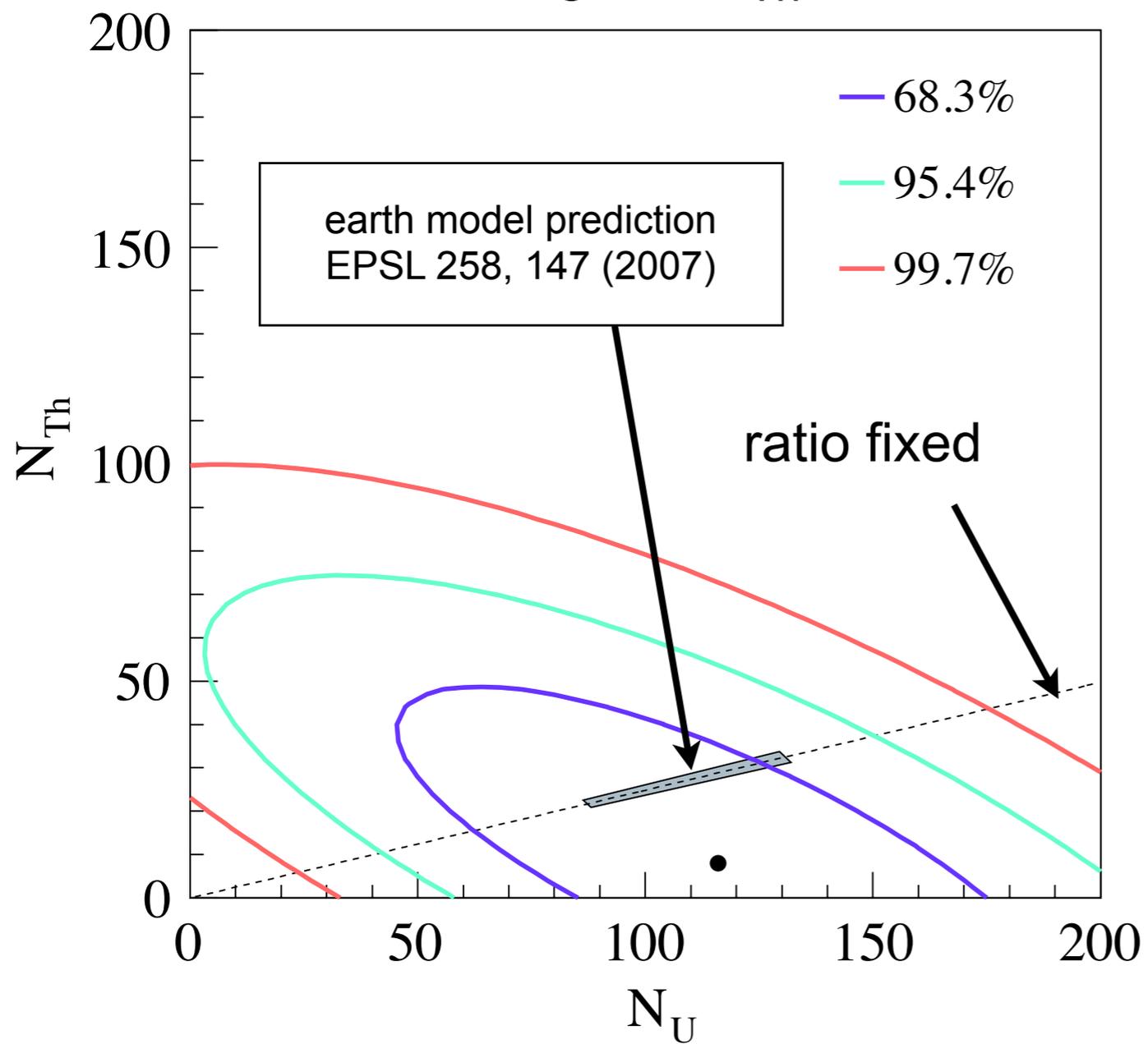


neutrino oscillation

$$P(E, L) \sim 1 - \frac{1}{2} \sin^2 2\theta_{12} \quad (\text{constant suppression})$$

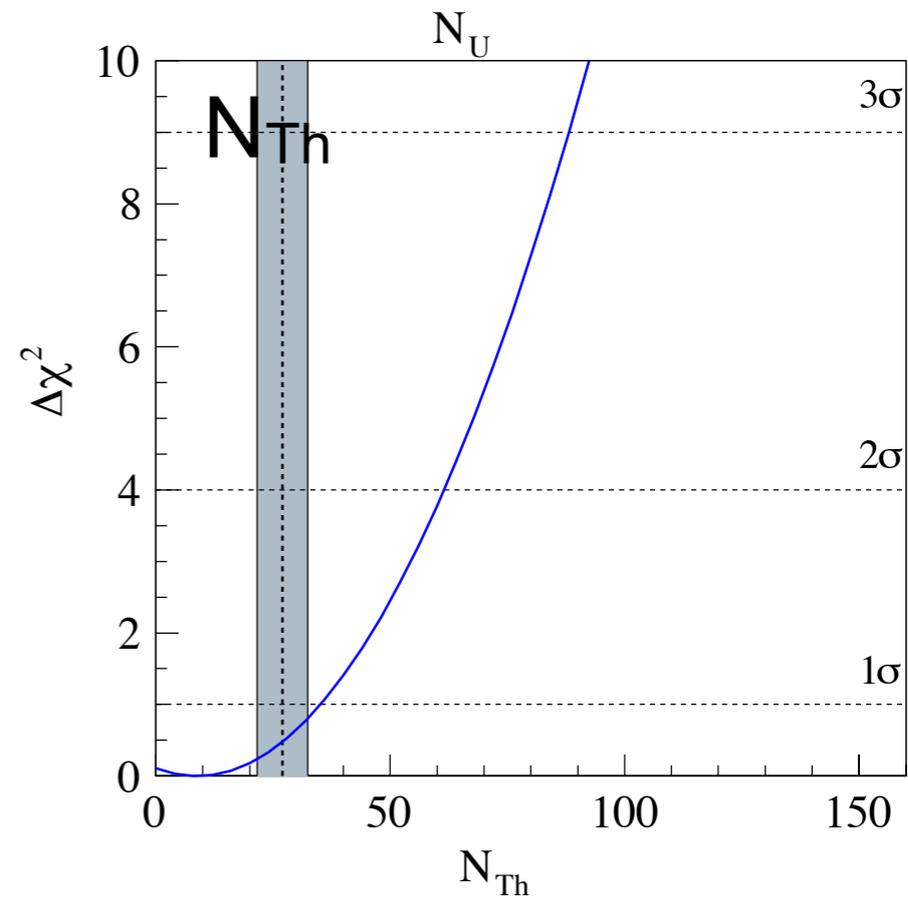
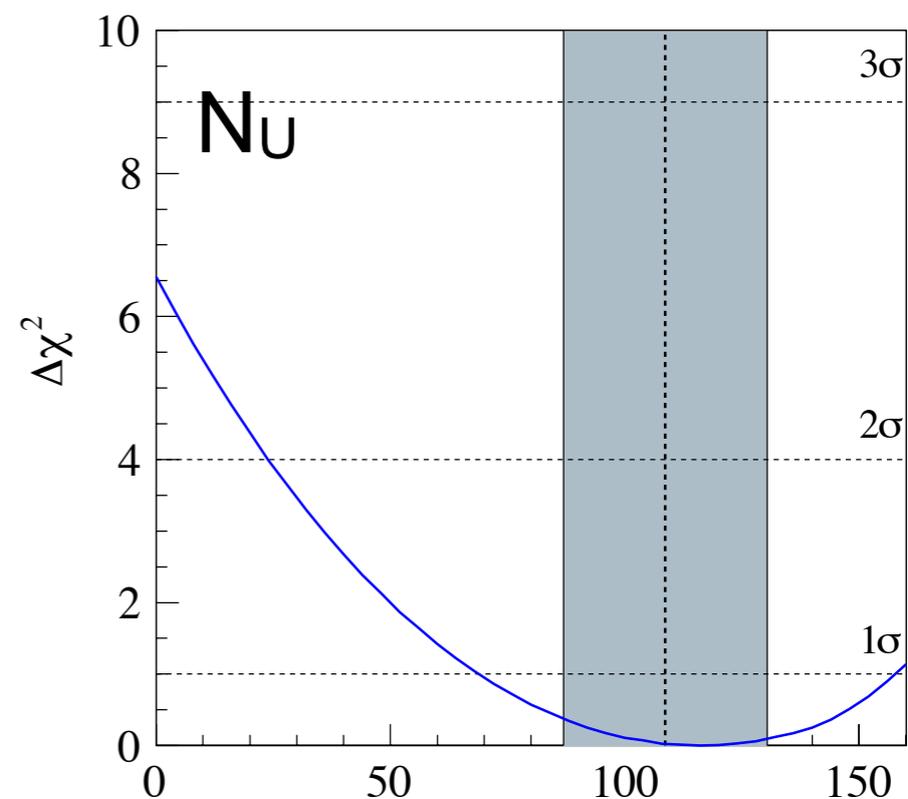
50% of the total flux originates within 500 km

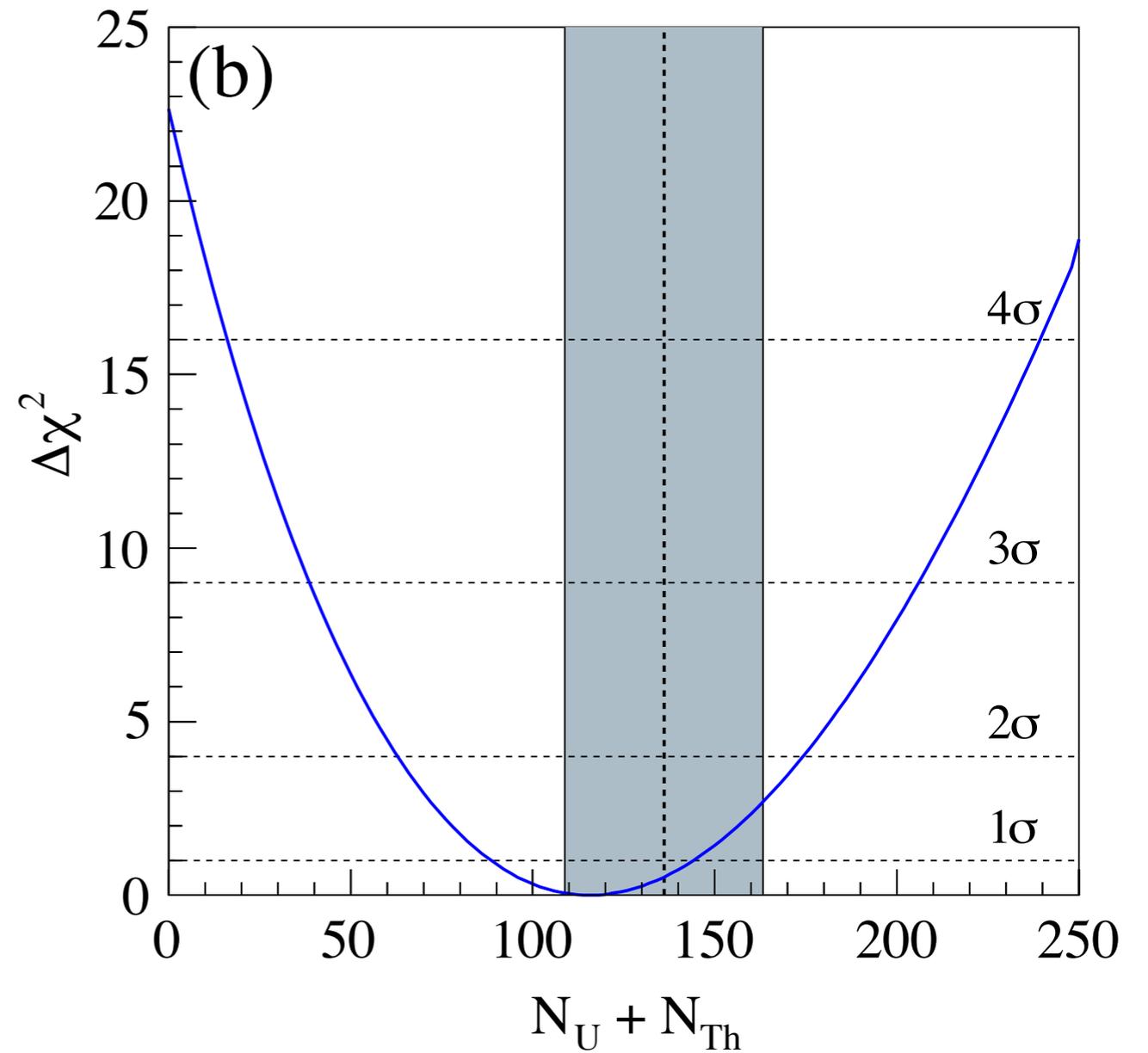
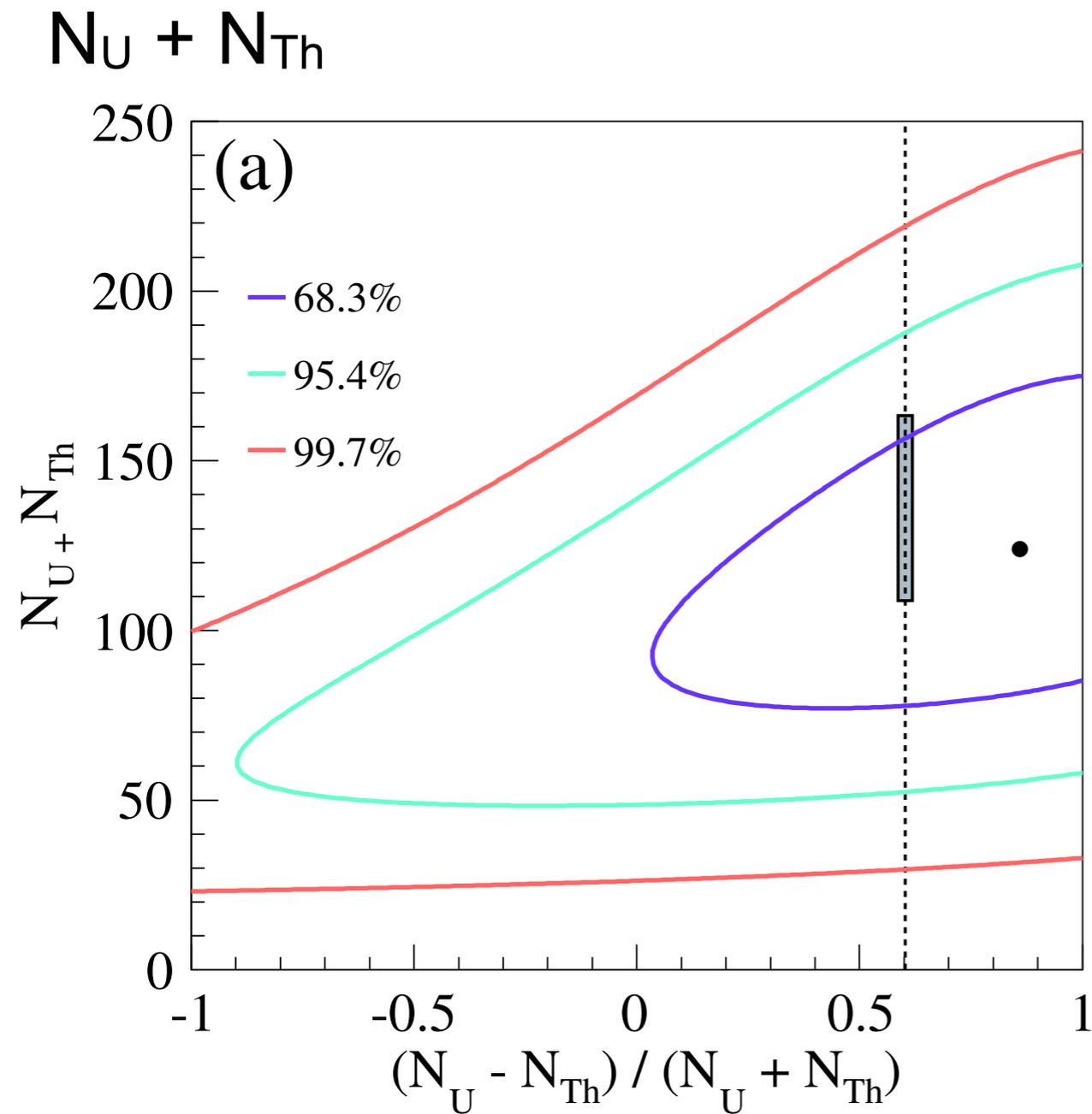
N_U vs N_{Th}



best-fit $(N_U, N_{Th}) = (116, 8)$

$N_U = 0$ signal : **rejected at 2.6σ (99.0%)**



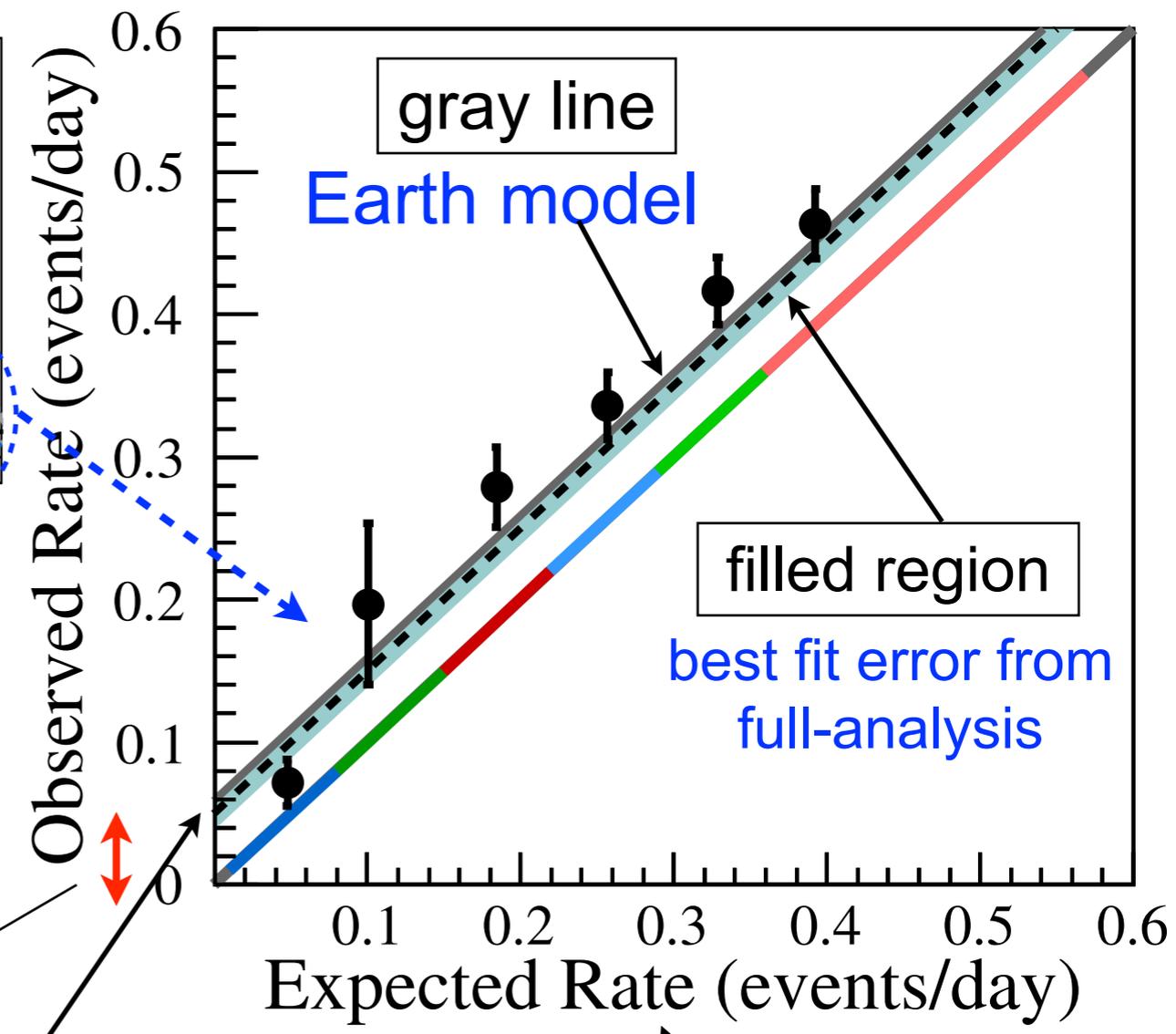
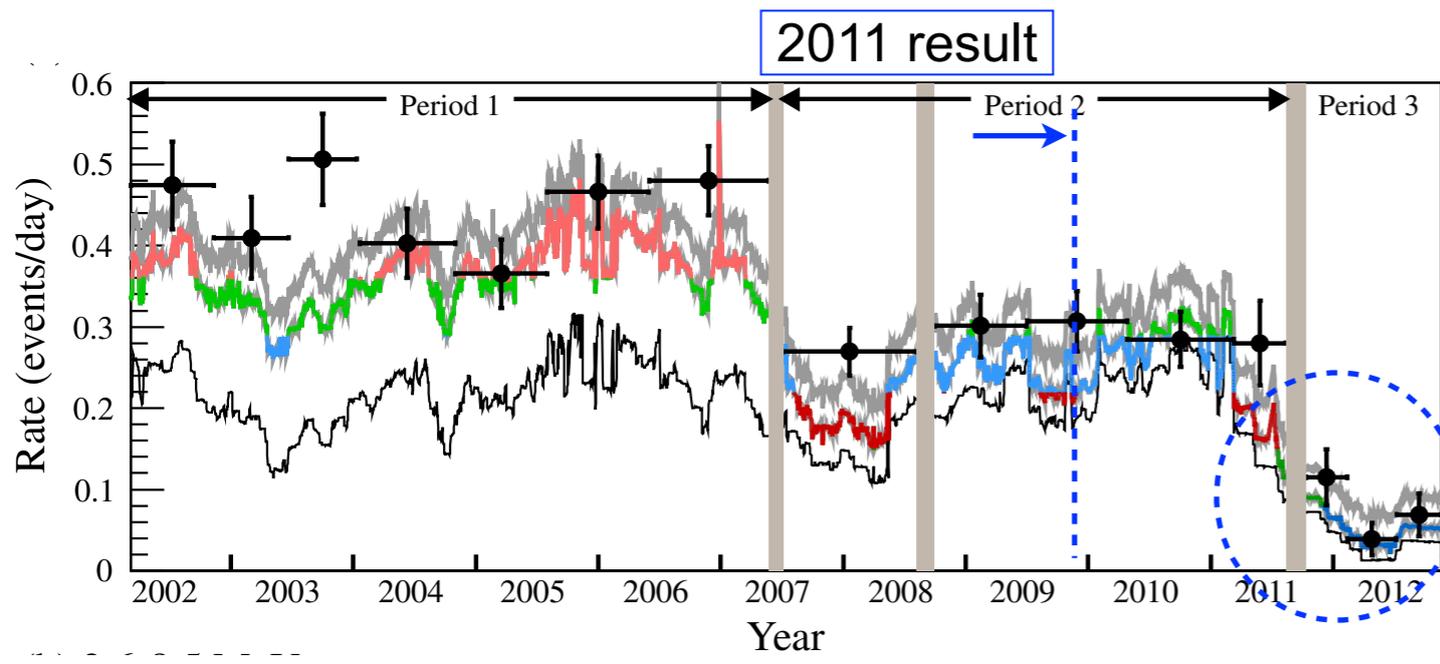


best-fit $N_U + N_{Th} = 116^{+28}_{-27}$

Flux : $3.4^{+0.8}_{-0.8} \times 10^6 \text{ cm}^{-2}\text{s}^{-1}$

0 signal rejected at 99.9998% C.L. (2×10^{-6})

- Expected Rate vs Observed Rate (0.9-2.6 MeV)



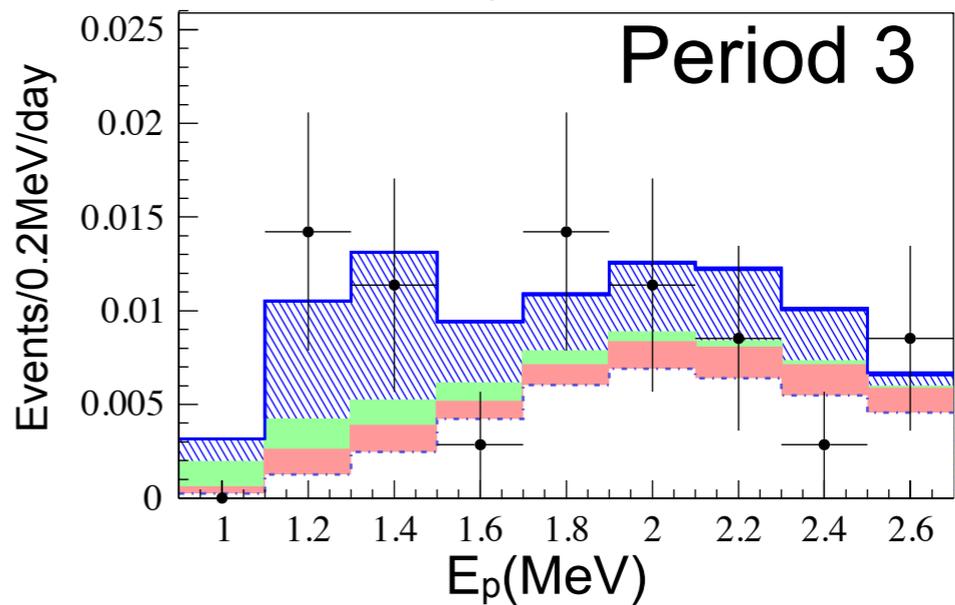
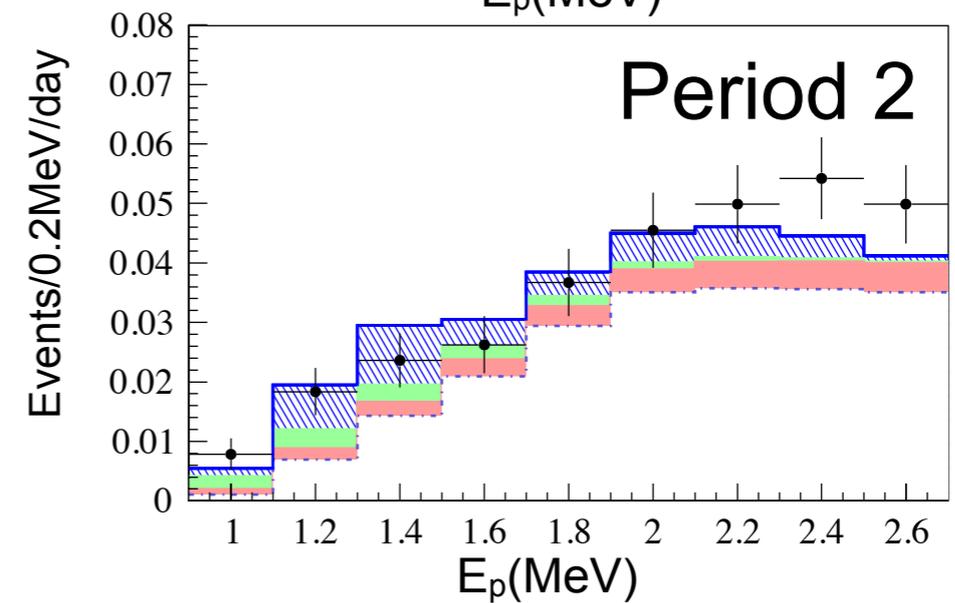
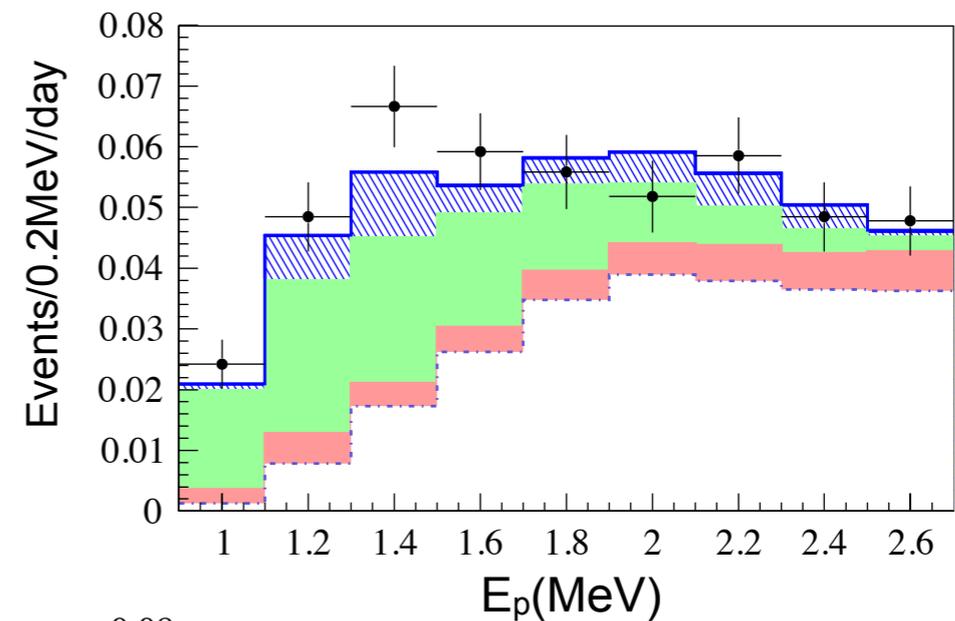
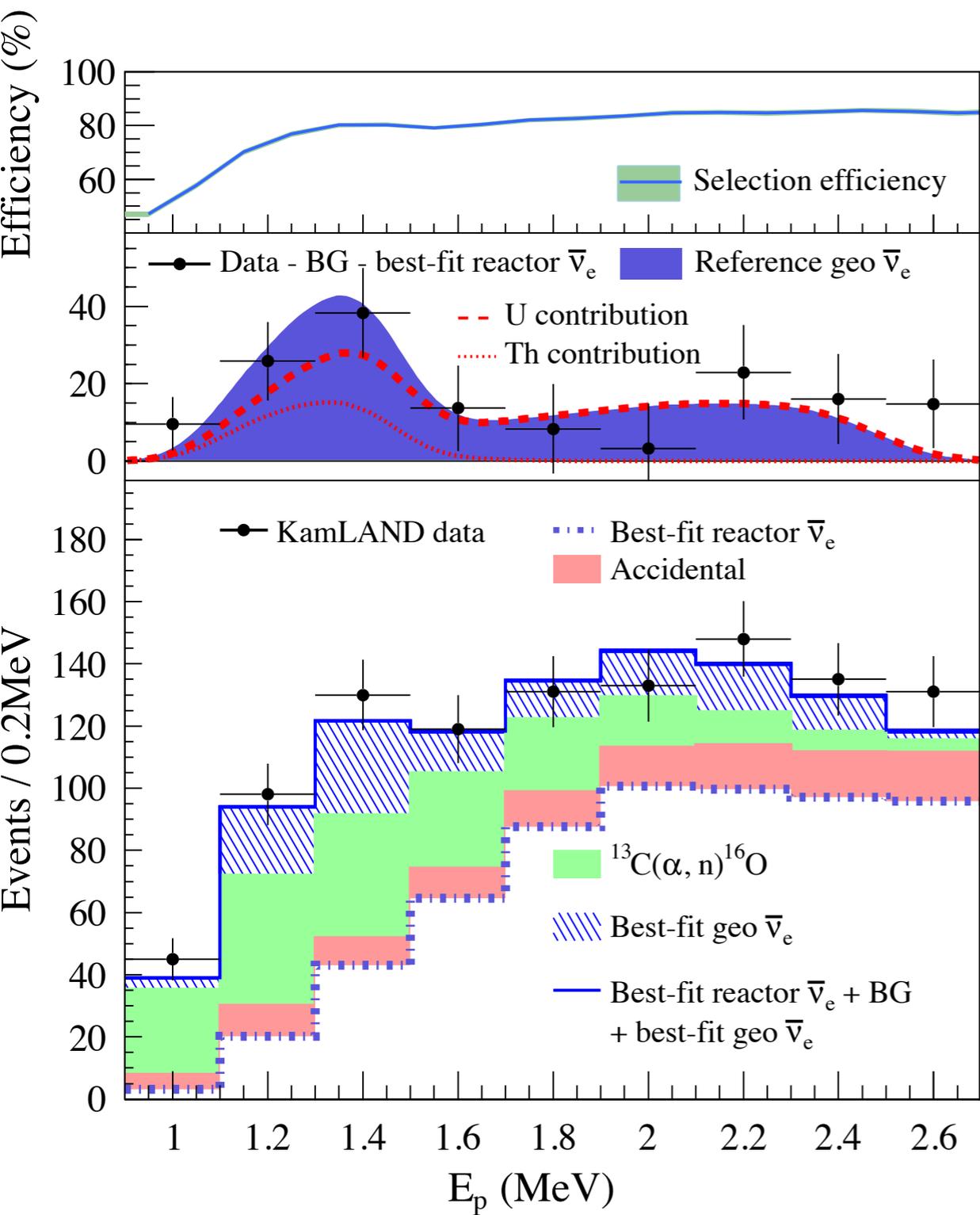
- Lower three data points can be added by using low-reactor operation period.

- Strong correlation between expected and observed event rate.

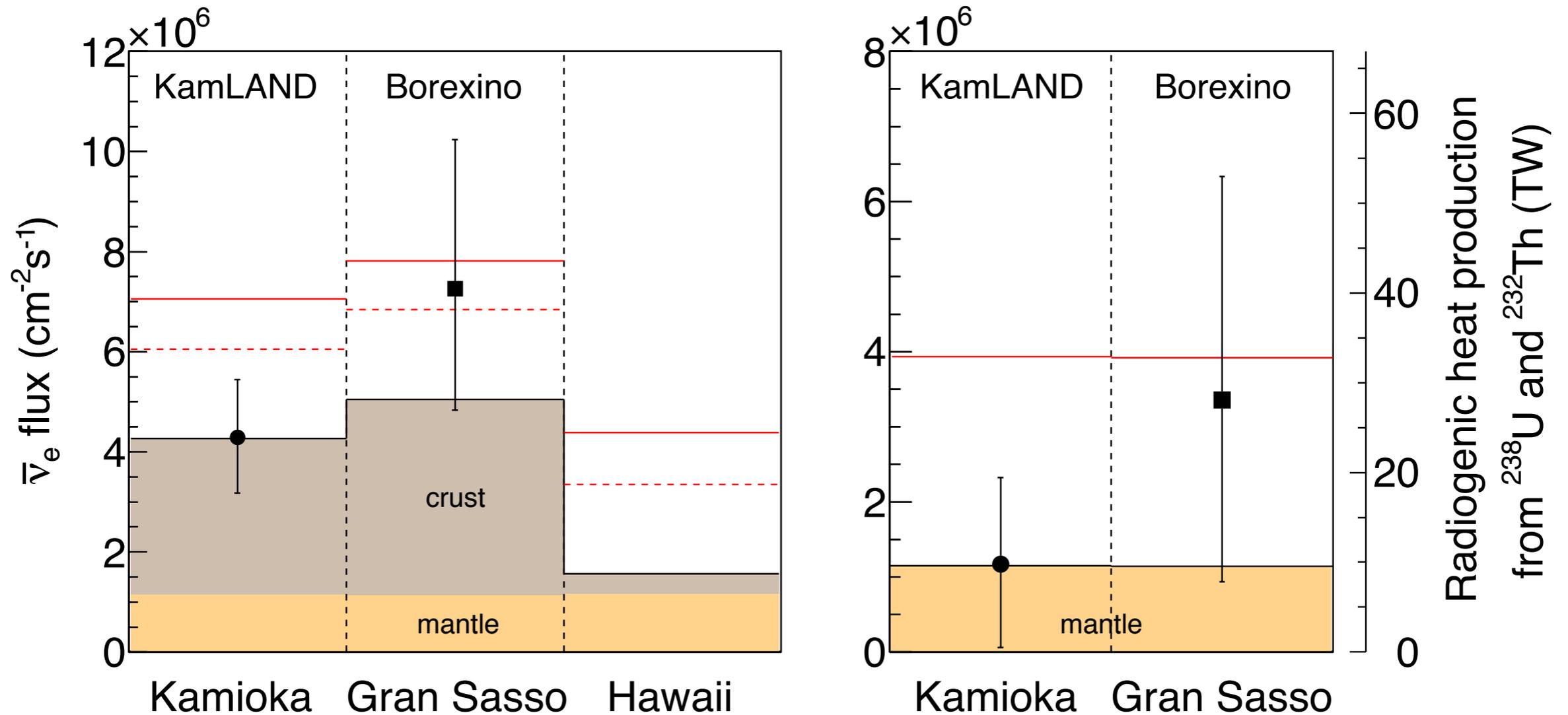
constant contribution of **geo-neutrino**

dotted line's intercept
 best fit from full-analysis (Rate+Shape+Time)
 116 +28/-27 events

reactor anti-neutrino + other backgrounds



Test of Fully-Radiogenic Model

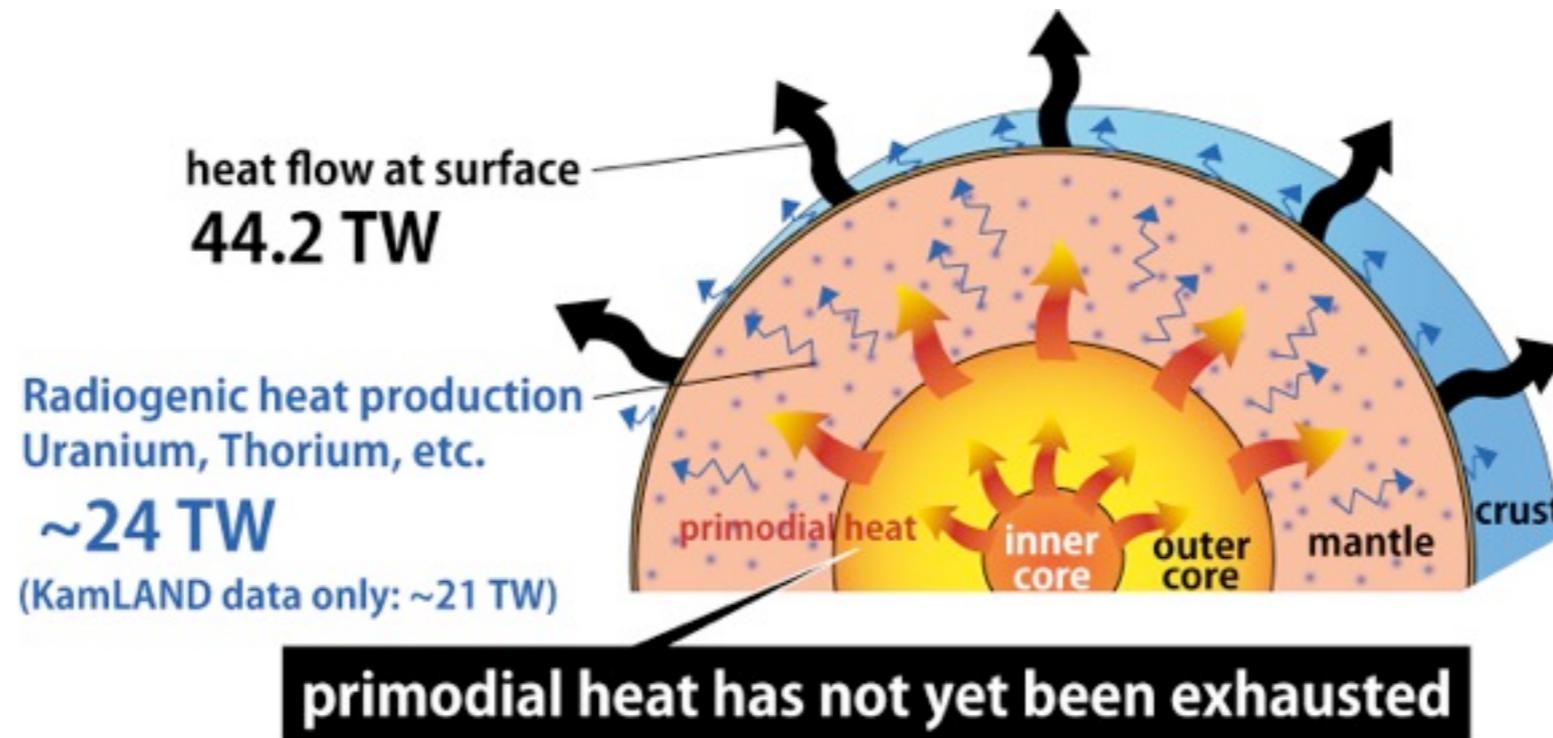


- Radiogenic heat production from ^{238}U and ^{232}Th is $20.0^{+8.8}_{-8.6}$ TW
- **Fully-radiogenic models** are disfavored at

KamLAND only 98.1% C.L. KamLAND + Borexino 97.2% C.L.

Measured Heat Balance

Result from geo neutrino experiment



“heat flow at surface: 44.2 ± 1.0 TW” – “radiogenic heat production: $24.3^{+8.8}_{-8.6}$ TW”

→ **primordial heat when the Earth formed**

- Geo neutrino data showed radiogenic heat ~24 TW contributes only half of Earth's total outgoing heat flux
- Geo reactor at the center of the Earth is constrained

KamLAND < 5.2 TW (90% C.L.)

Borexino < 3 TW (95% C.L.)