

Snoplus: Recent results and status

NSD Staff Meeting

Morgan Askins

21 September 2021

Snoplus Group at Berkeley

Group Lead

Gabriel Orebi Gann



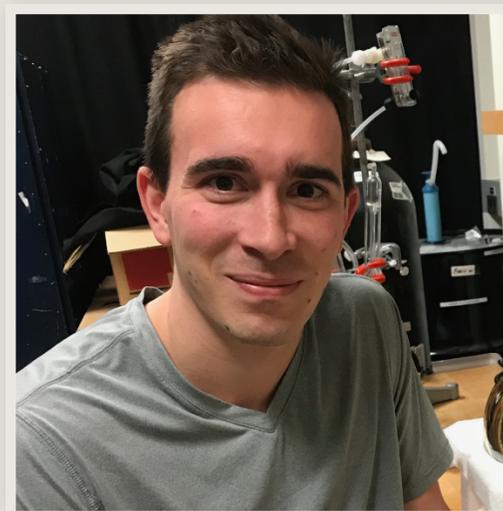
<https://underground.physics.berkeley.edu/>

Postdocs

Morgan Askins



Tanner Kaptanoglu



Graduate Students

Ed Callaghan



Max Smiley



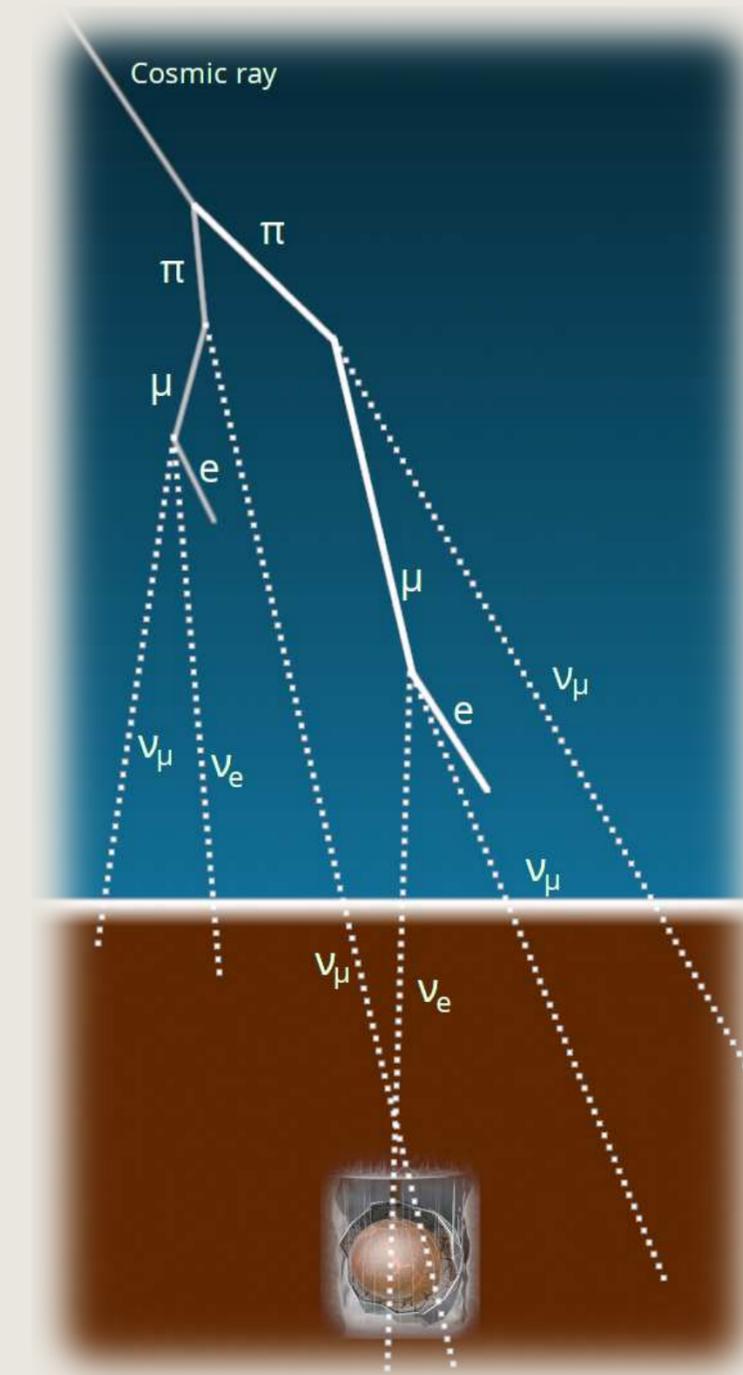
Snoplus Detector

Located deep underground in Sudbury, Ontario at SNOLab, the SNO+ detector is an upgrade to the SNO detector, designed to explore many neutrino-related physics topics, and ultimately search for neutrino-less double beta decay.

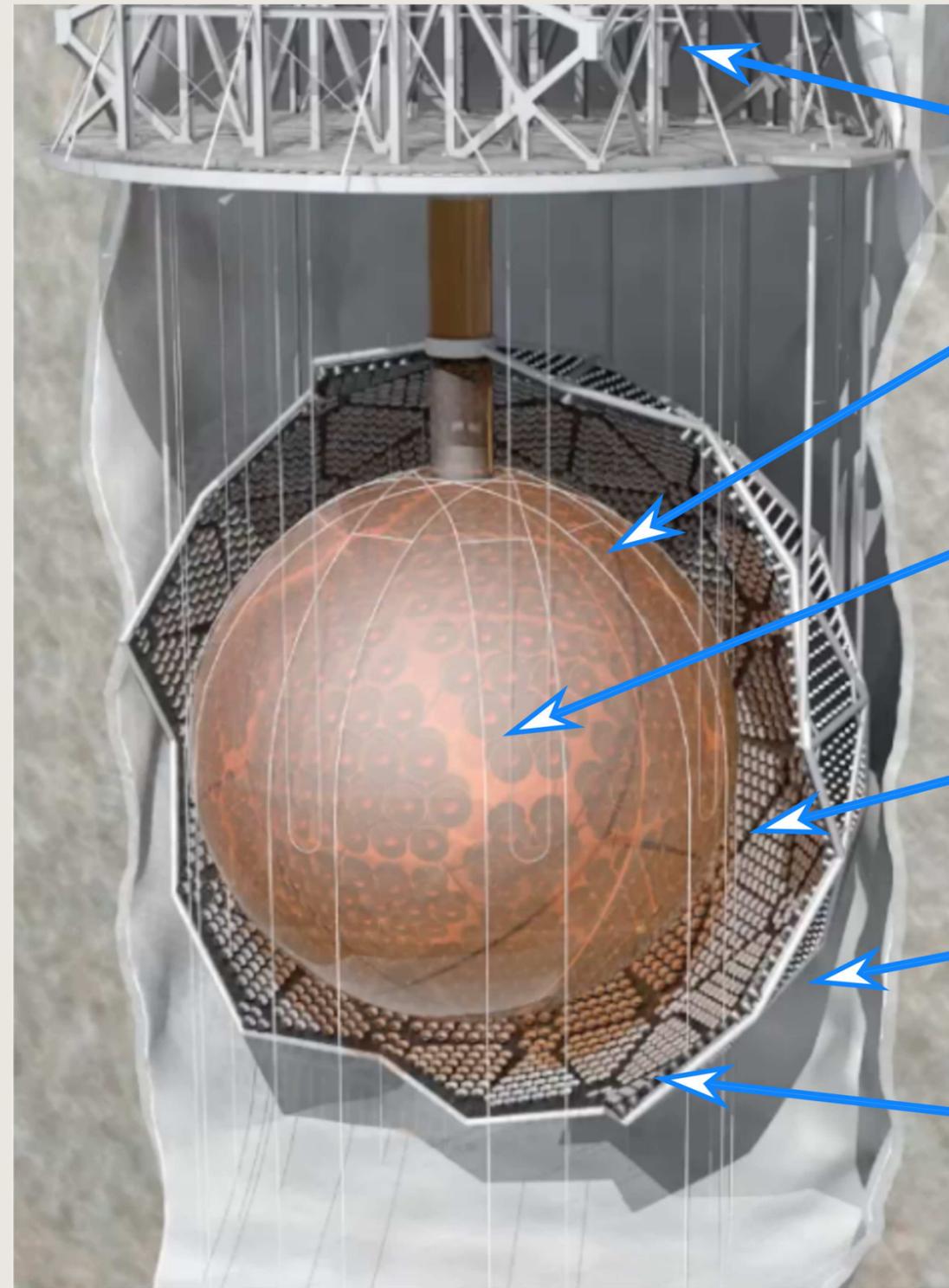
- Modern underground lab provides high levels of cleanliness
- 6800 ft. overburden significantly reduces muon flux—ideal for rare event searches.

Full detector details:

JINST 16 (2021) 08, P08059



Snoplus Detector



Overburden: 6800 ft. (5890 m.w.e.)

Acrylic Vessel (12 m Diameter)

905 Tonnes H₂O (phase I)

780 Tonnes Te-doped Scintillator (phase II)

1700 Tonnes H₂O Inner Buffer

5700 Tonnes H₂O Outer Buffer

~9300 PMTs

SNO+ Physics Goals

→ Water Phase

- Search for invisible nucleon decay
- Measure ^8B solar ν flux
- Characterize backgrounds from detector components
- Measure neutron capture cross-section in water

→ Pure Scintillator Phase

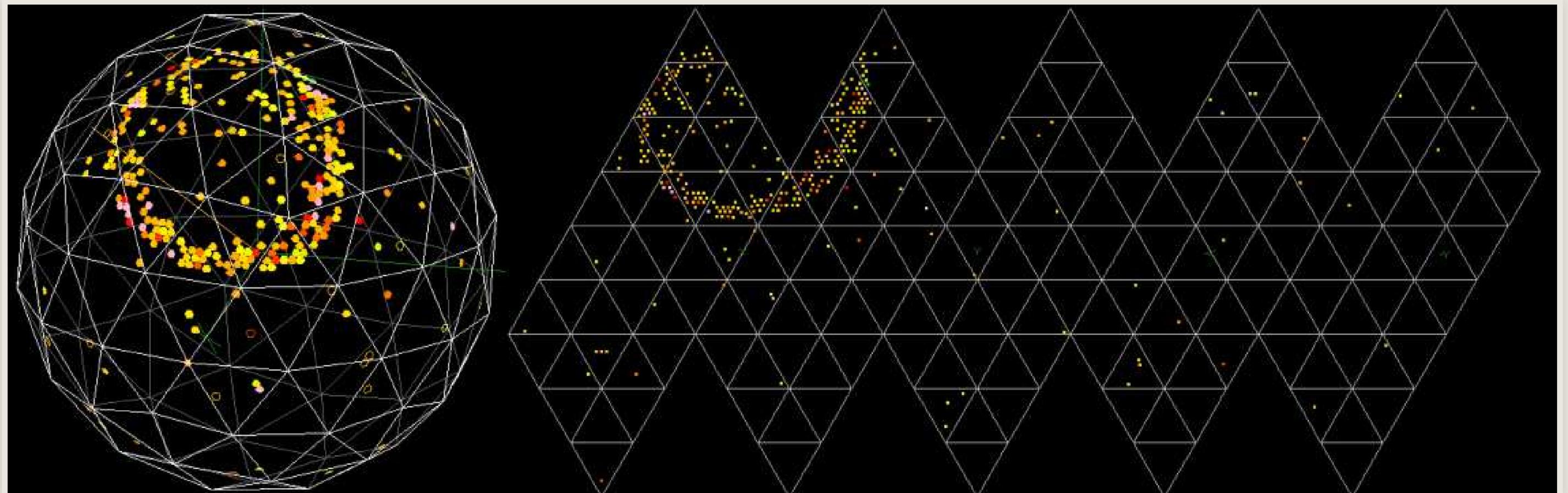
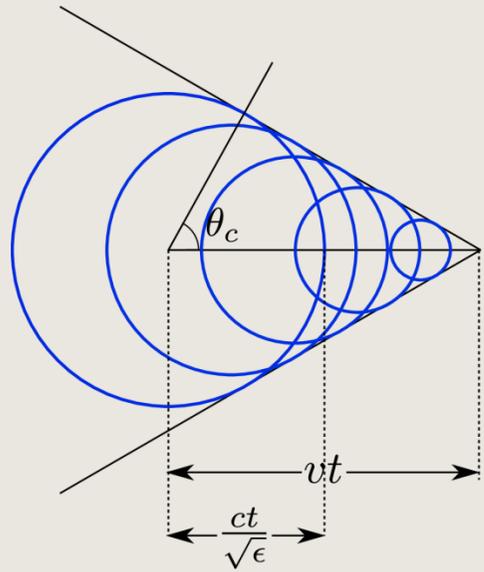
- Measure reactor and geo $\bar{\nu}_e$
- Ready to handle supernova burst measurements
- Characterize scintillator backgrounds
- Precision solar ν measurements
- Demonstration of directionality in liquid scintillator

→ Te-loaded Scintillator Phase

- Search for neutrinoless double beta decay
- Continuation of pure scintillator physics

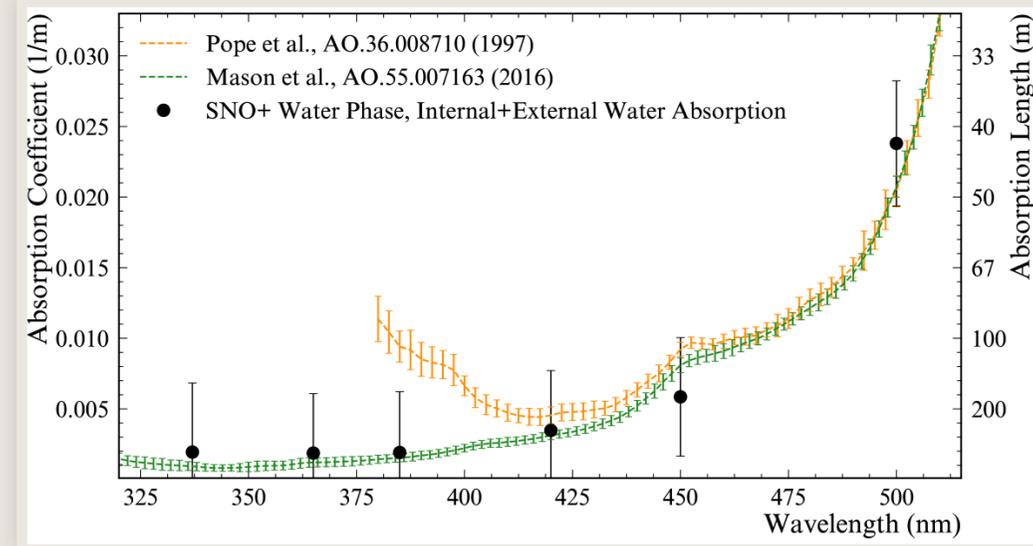
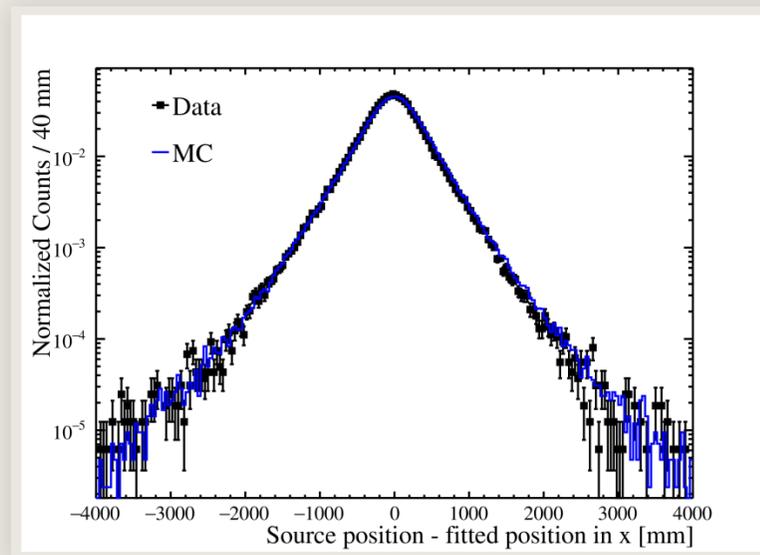
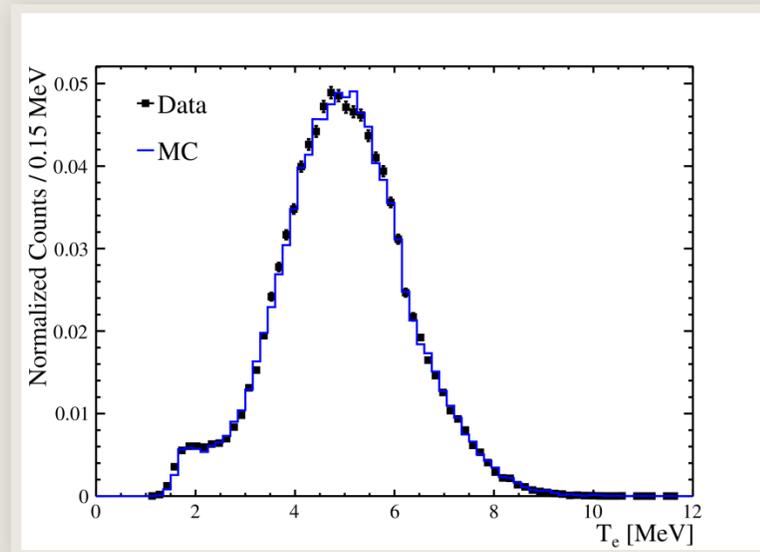
Water Phase

- Water phase used for detector commissioning, as well as physics data taking.
- Light emission only through Cherenkov radiation.
- Cherenkov provides directional information useful for solar ν .
- Energy resolution is statistically limited by low light output.



Water phase calibration

Tagged ^{16}N source used for reconstruction calibration and evaluation of systematics.

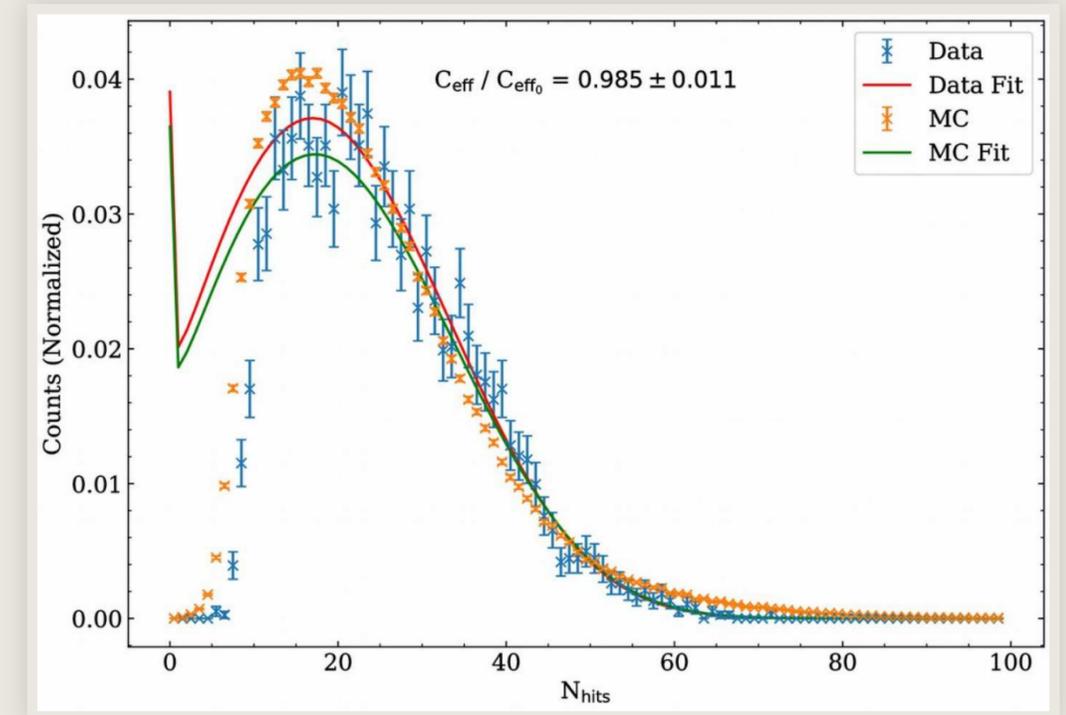
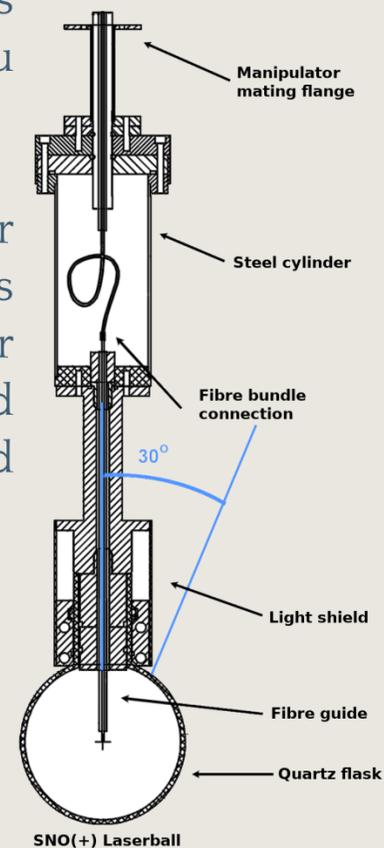


Pure Cherenkov light source developed by Berkeley, using ^8Li β -decays to produce light directly in an acrylic sphere.

Optical Properties measured via an in-situ laser diffusion sphere.

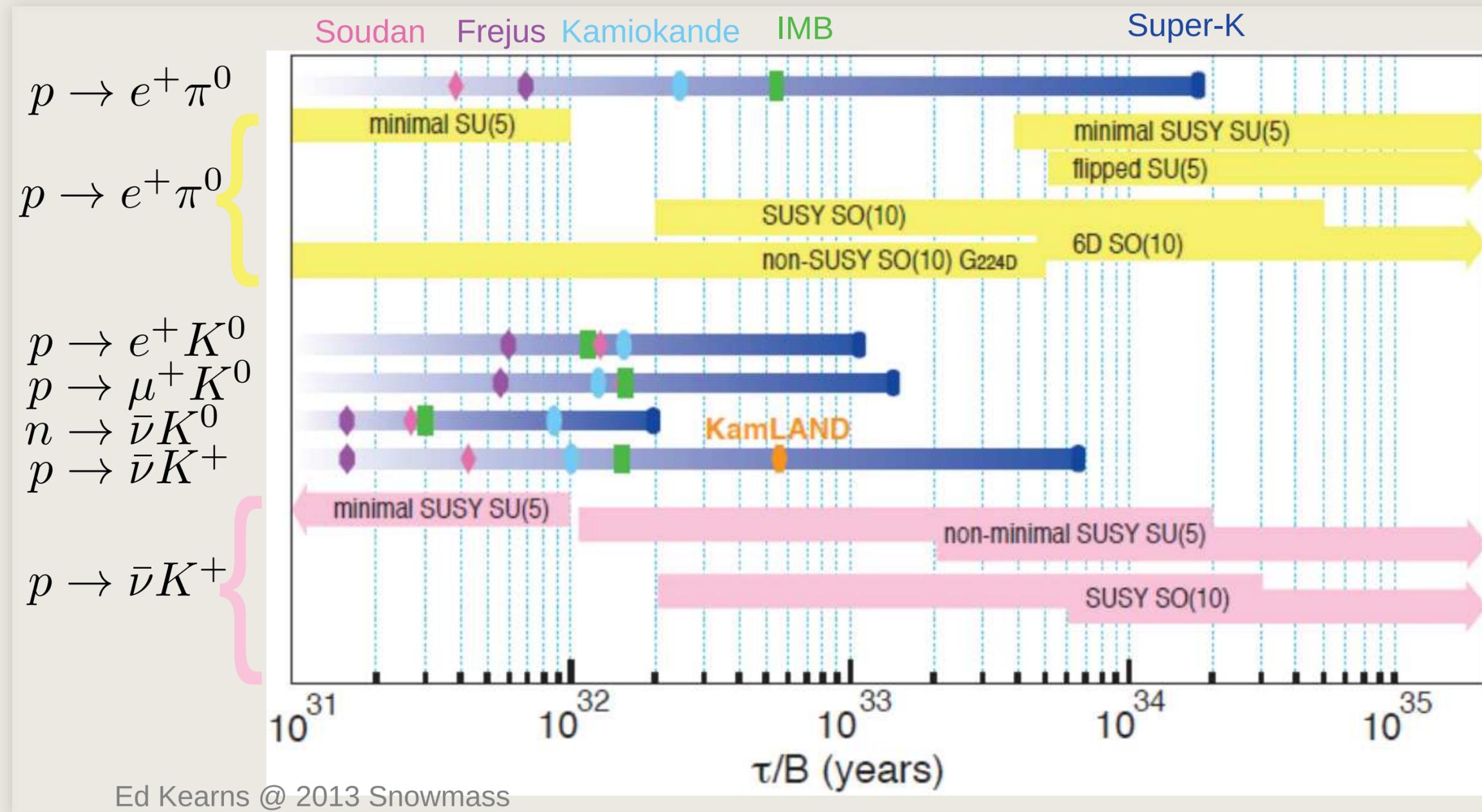
Attenuation of the water and acrylic vessel, as well as the PMT angular response measured and verified with a deployed ^{16}N source.

[arxiv:2106.03951](https://arxiv.org/abs/2106.03951)



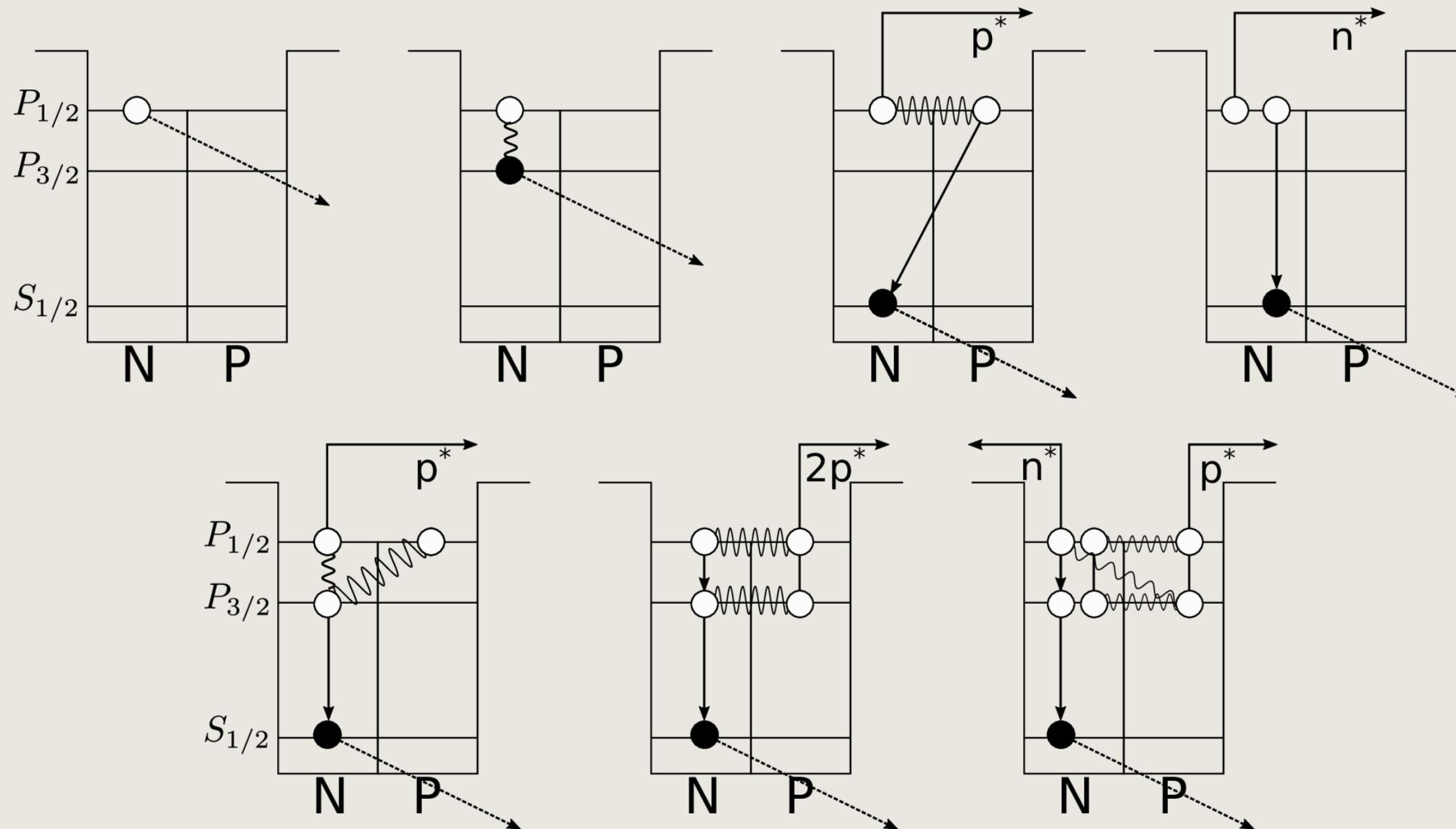
Nucleon Decay

The search for nucleon decay (particularly on free protons) has been the motivation for many large experiments.



Invisible Modes in SNO+

Deexcitation Mode of Neutron Hole in ^{15}N



Proton Decay

$$^{16}\text{O} \rightarrow ^{15}\text{N}^*$$

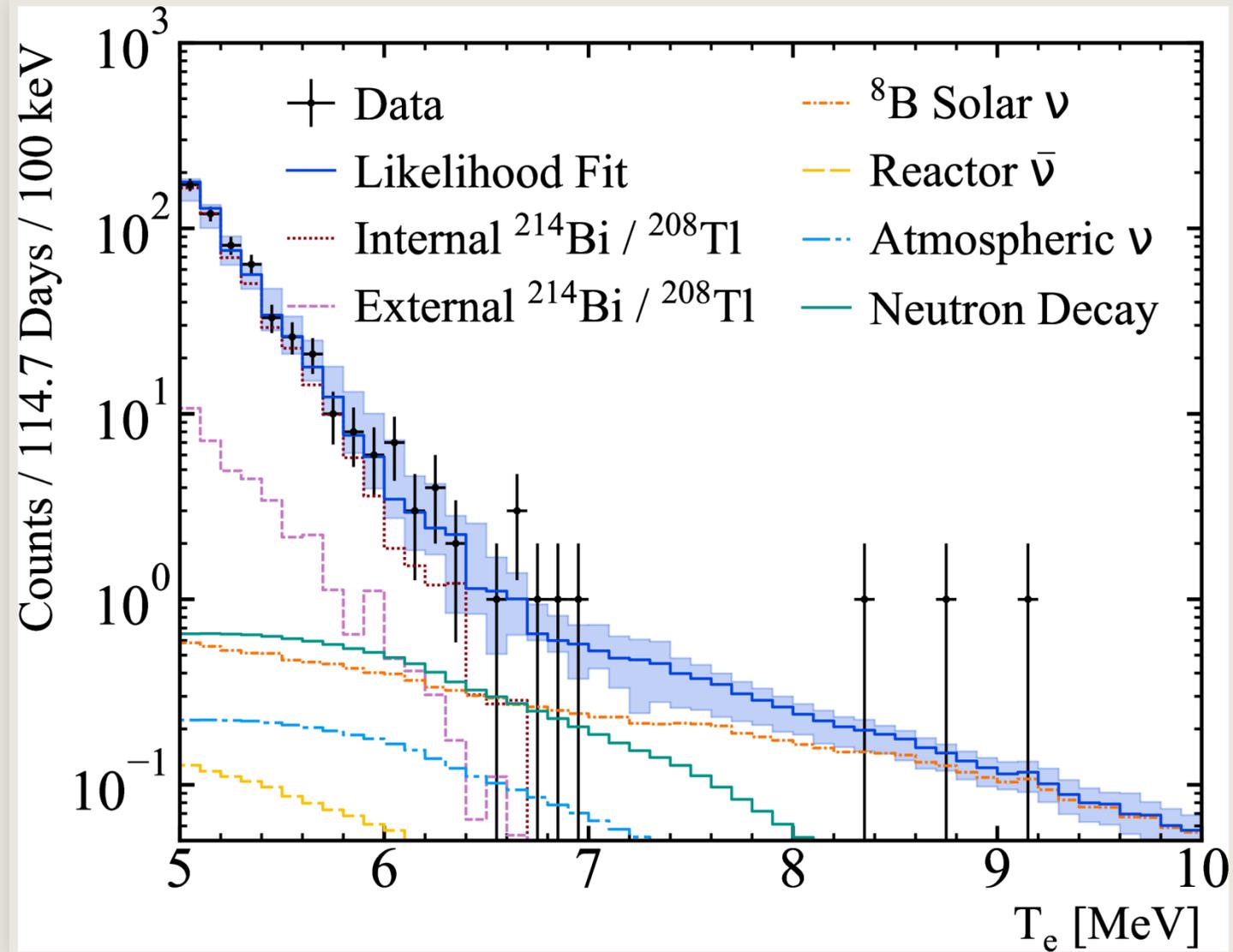
3%: 9.93 MeV γ
 41%: 6.32 MeV γ

Neutron Decay

$$^{16}\text{O} \rightarrow ^{15}\text{O}^*$$

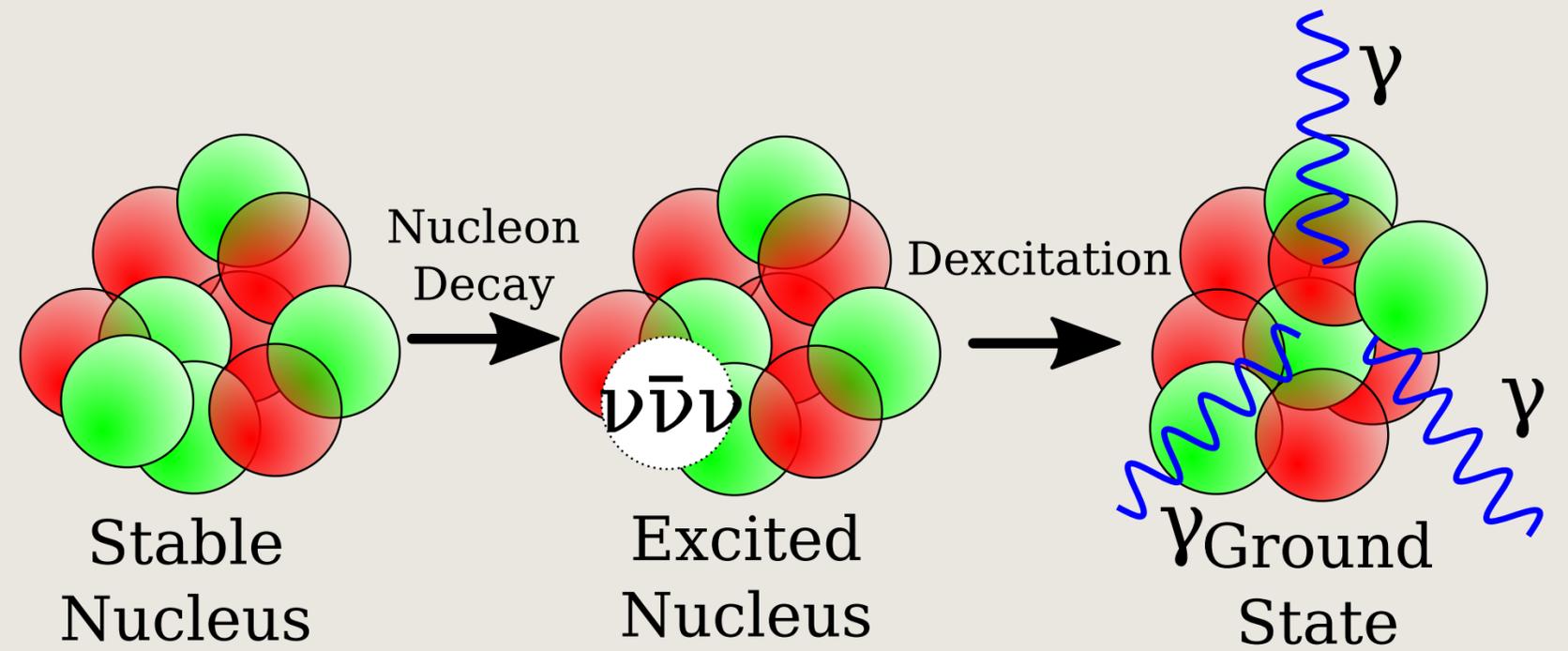
44%: 6.18 MeV γ

First Results: Invisible Nucleon Decay



Phys. Rev. D 99 (2019) 3, 032008

Decay Mode	Partial Lifetime
n	2.5×10^{29}
p	3.6×10^{29}
pp	4.7×10^{28}
pn	2.6×10^{28}
nn	1.3×10^{28}

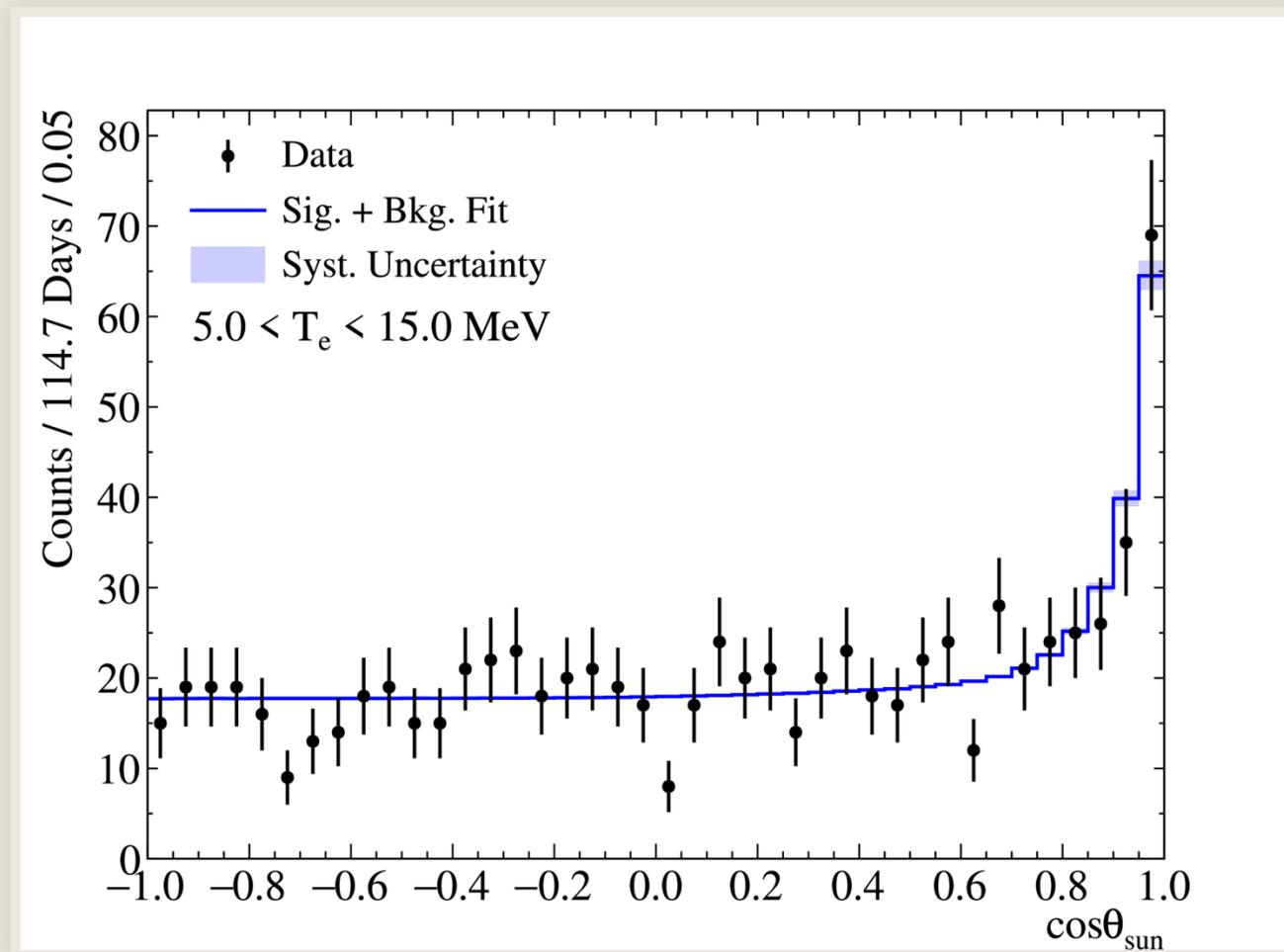


First Results: Solar Flux

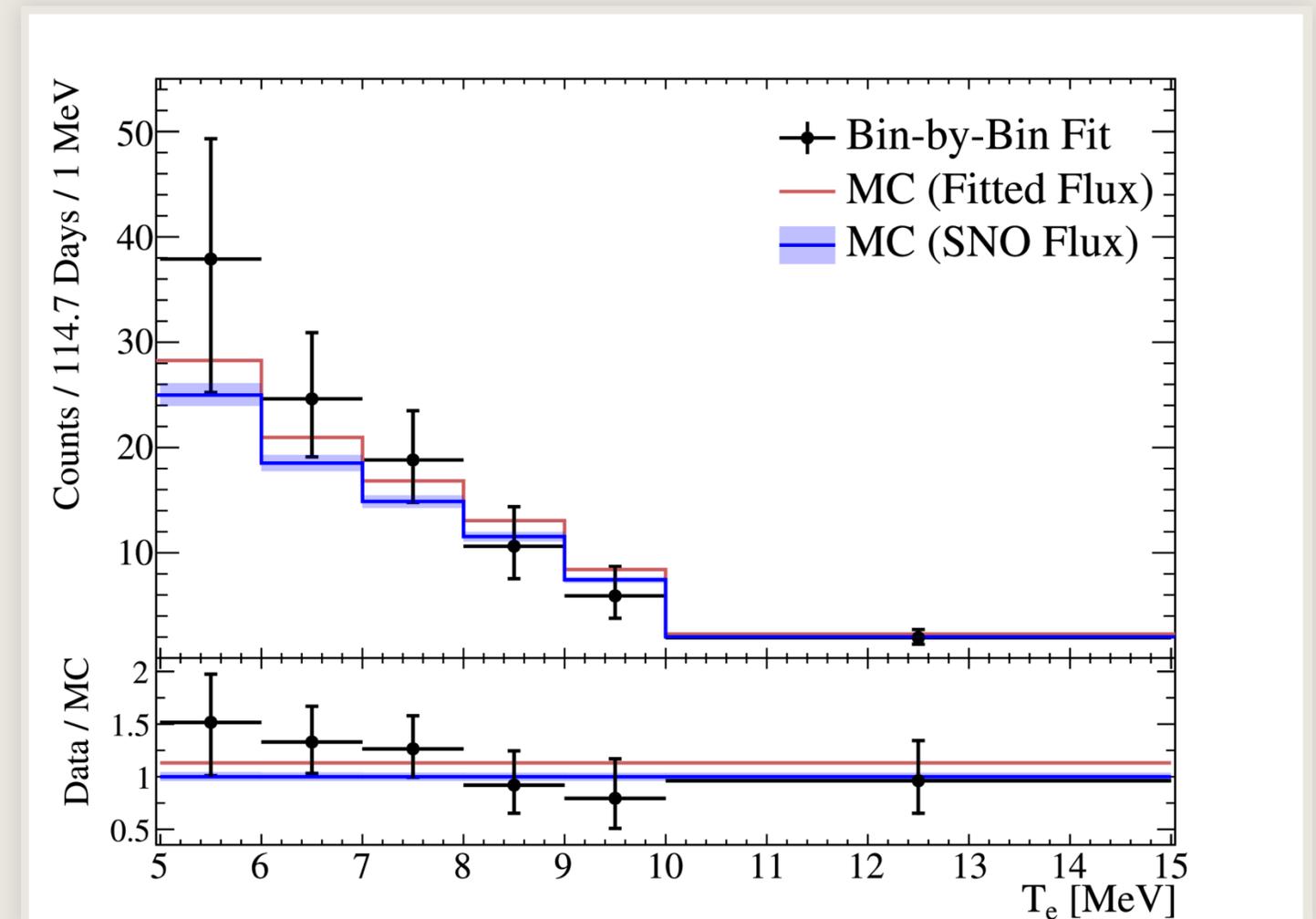
Event direction in water shows clear evidence for solar neutrinos on top of a flat background.

Measured flux consistent with SNO and Super-K

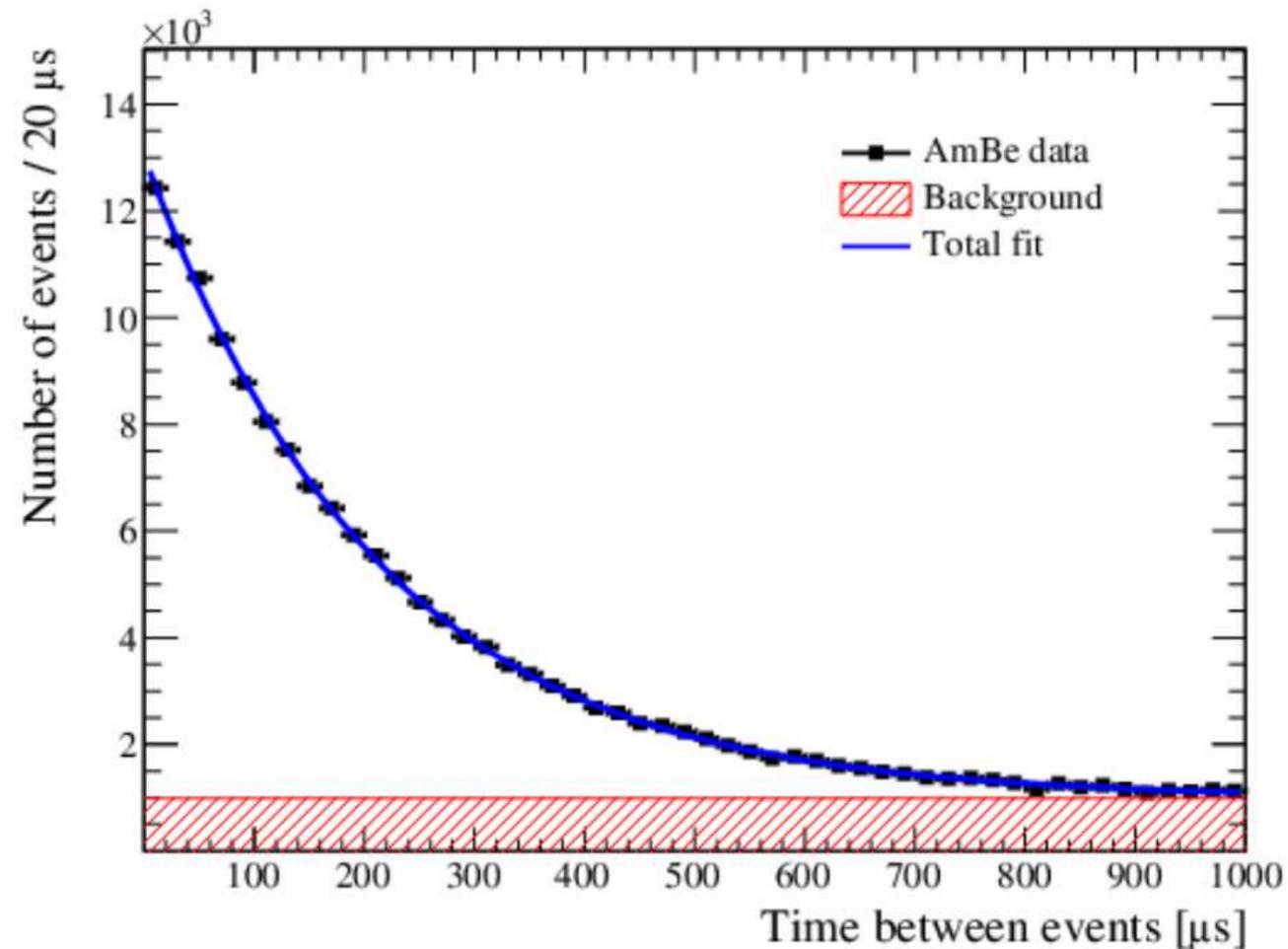
$$\Phi_{8B} = 5.95_{-0.71}^{0.75} (\text{stat})_{-0.10}^{0.13} (\text{syst}) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$$



Phys. Rev. D. 99 (2019) 1, 012012



First Results: Neutron Capture



Measurement of thermal neutron-proton capture cross-section with high detection efficiency in pure water.

Efficiency [%]	τ [μs]
49.09 ± 0.39	$202.35^{+0.87}_{-0.76}$

With a thermal cross-section consistent with dedicated experiments

$$\sigma_H = 336.3^{+1.2}_{-1.5} \text{ mb}$$

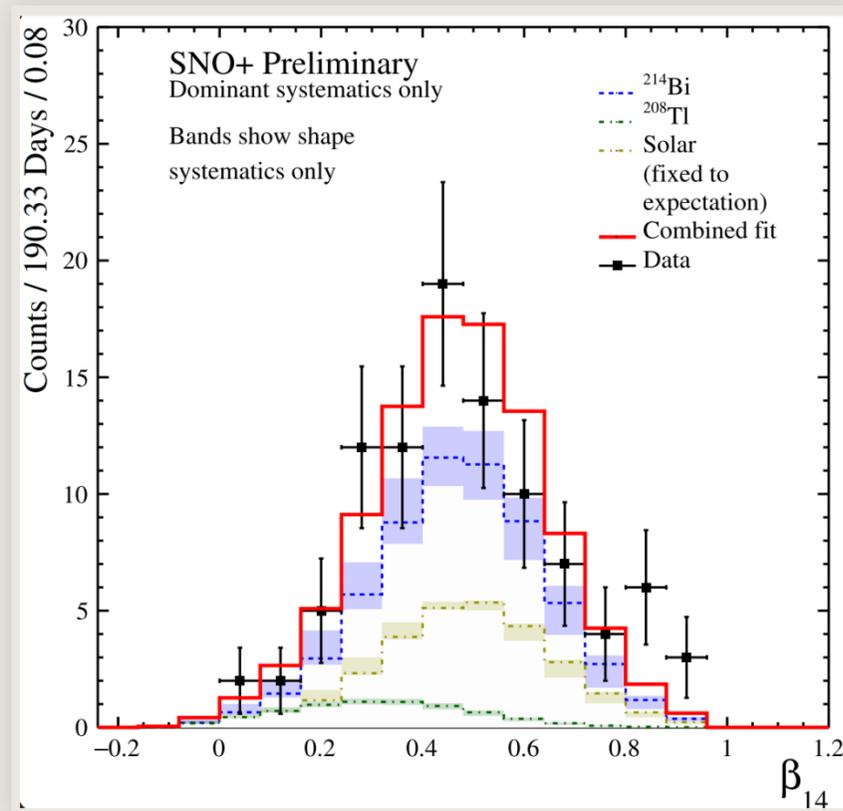
Phys. Rev. C. 102 (2020) 1, 014002

Imminent water results

New low-background dataset triples the total livetime and greatly improves sensitivity due to decrease internal backgrounds.

	gU/gH_2O	gTh/gH_2O
Dataset I	$(3.6 \pm 0.9^{+1.0}_{-0.7}) \times 10^{-14}$	$< 1.3 \times 10^{-14}$ (95%CL)
Dataset II	$(3.2 \pm 0.7^{+1.1}_{-0.9}) \times 10^{-15}$	$< 1.1 \times 10^{-15}$ (95%CL)

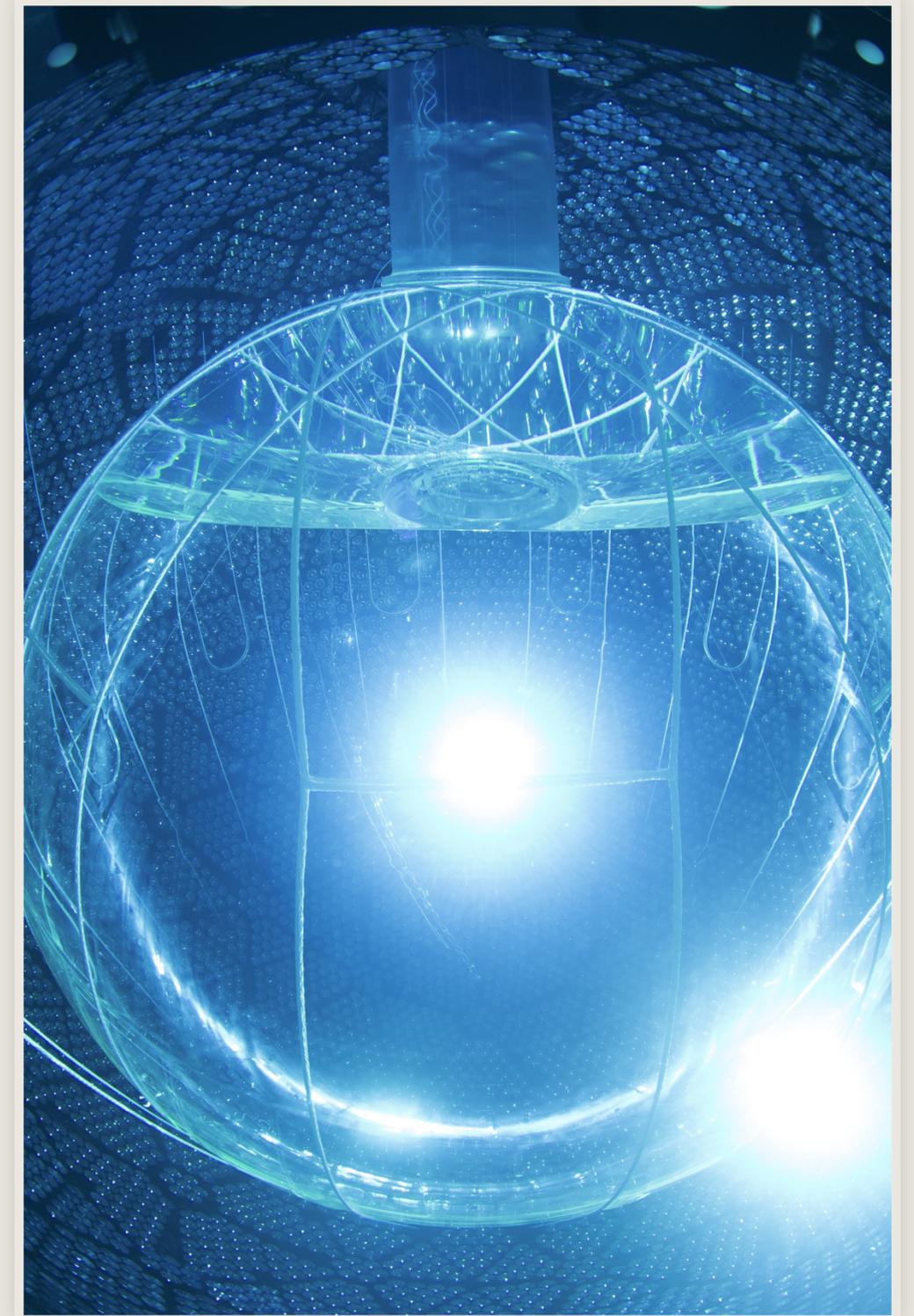
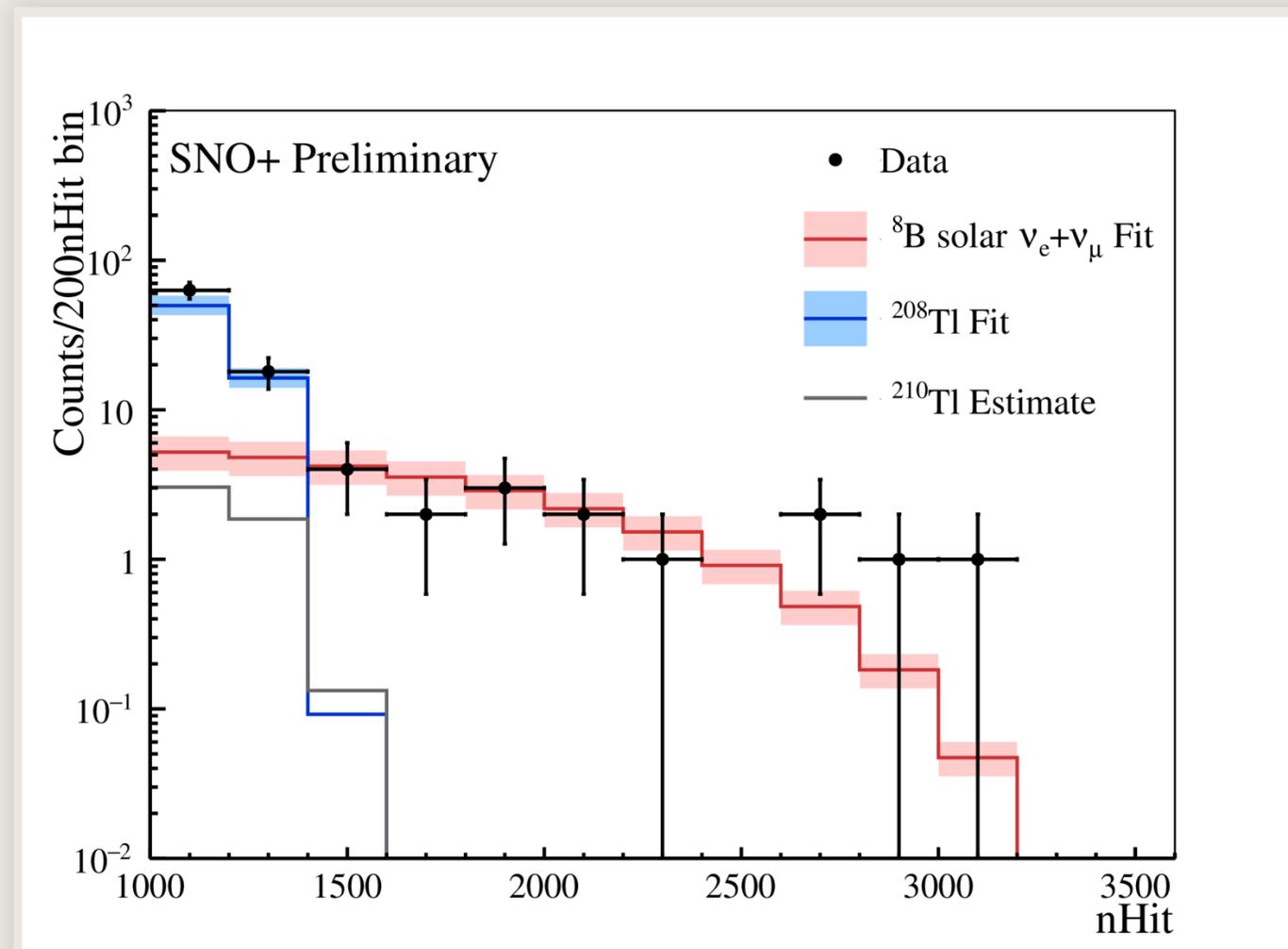
- † Improved sensitivity to invisible nucleon decay.
- † Measurement of reactor antineutrinos.
- Improved measurement of the solar 8B flux.



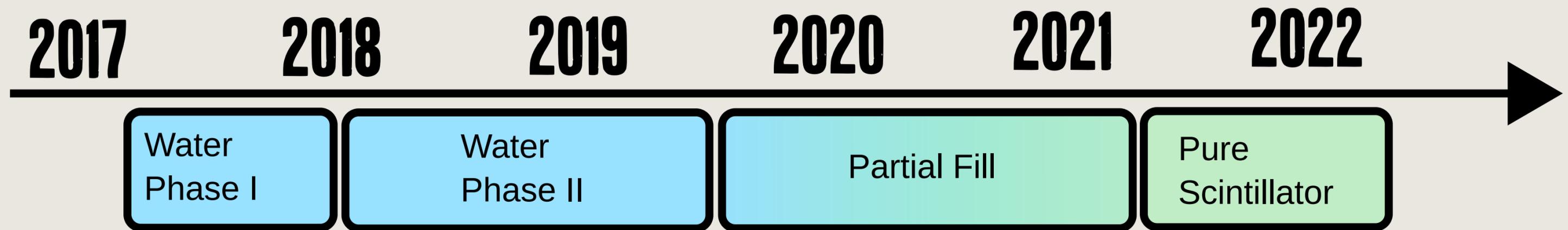
† *Berkeley lead efforts.*

Partial-fill

Extended filling phase provided opportunity to test analysis methods prior to completing the detector. This included measurements of various internal backgrounds as well as ^8B solar neutrinos.



Status and Timeline

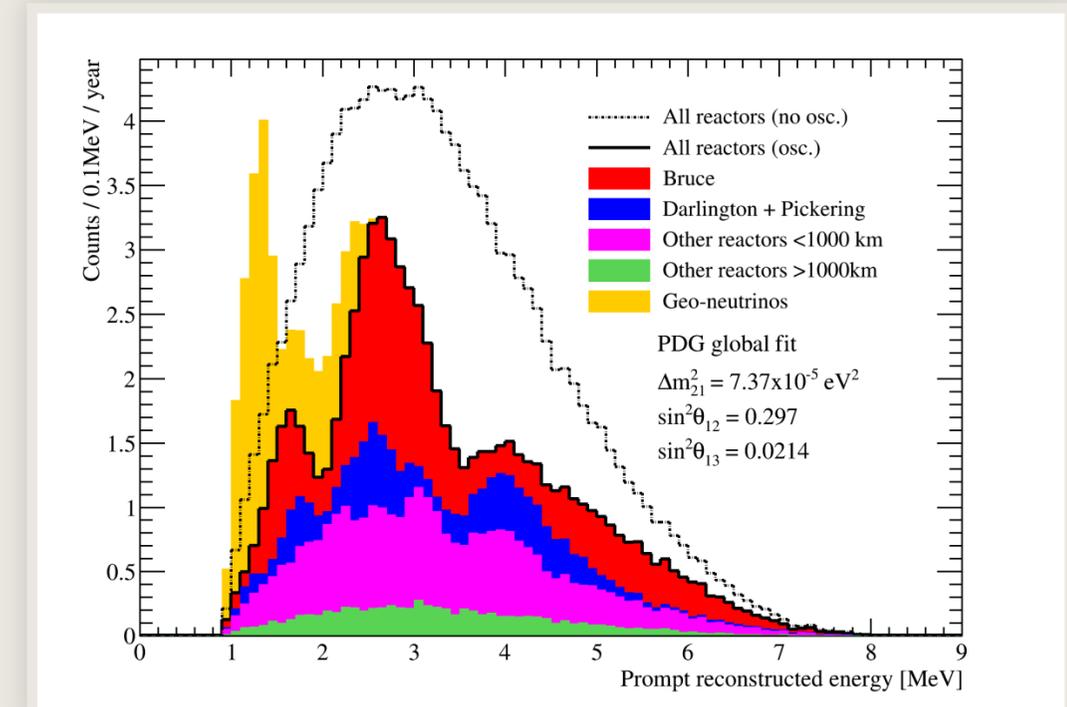
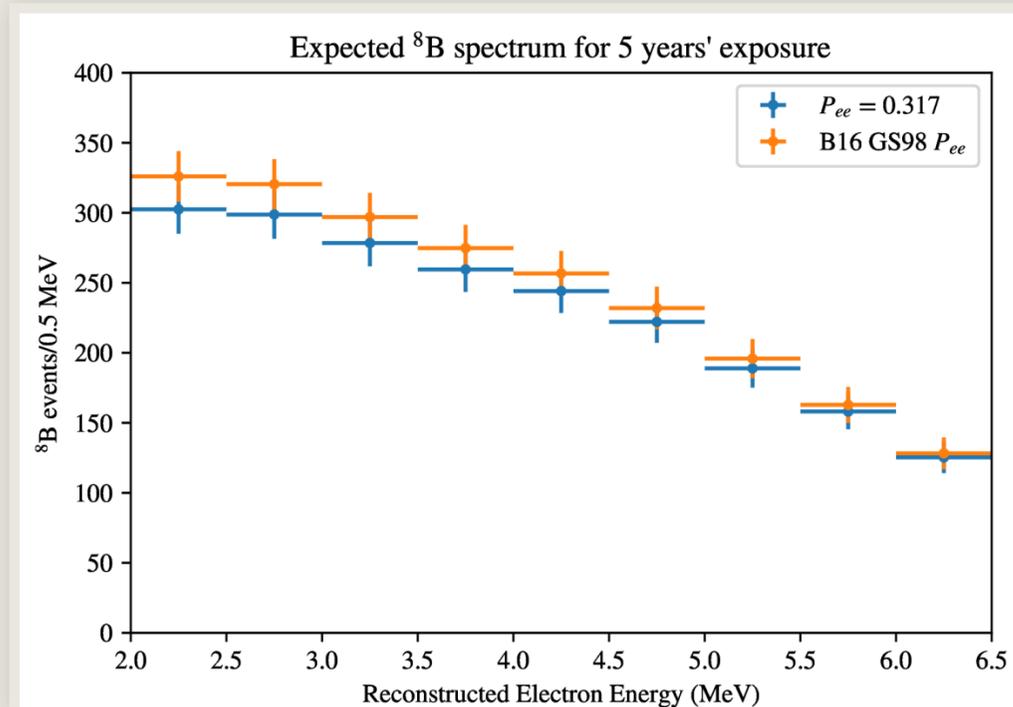
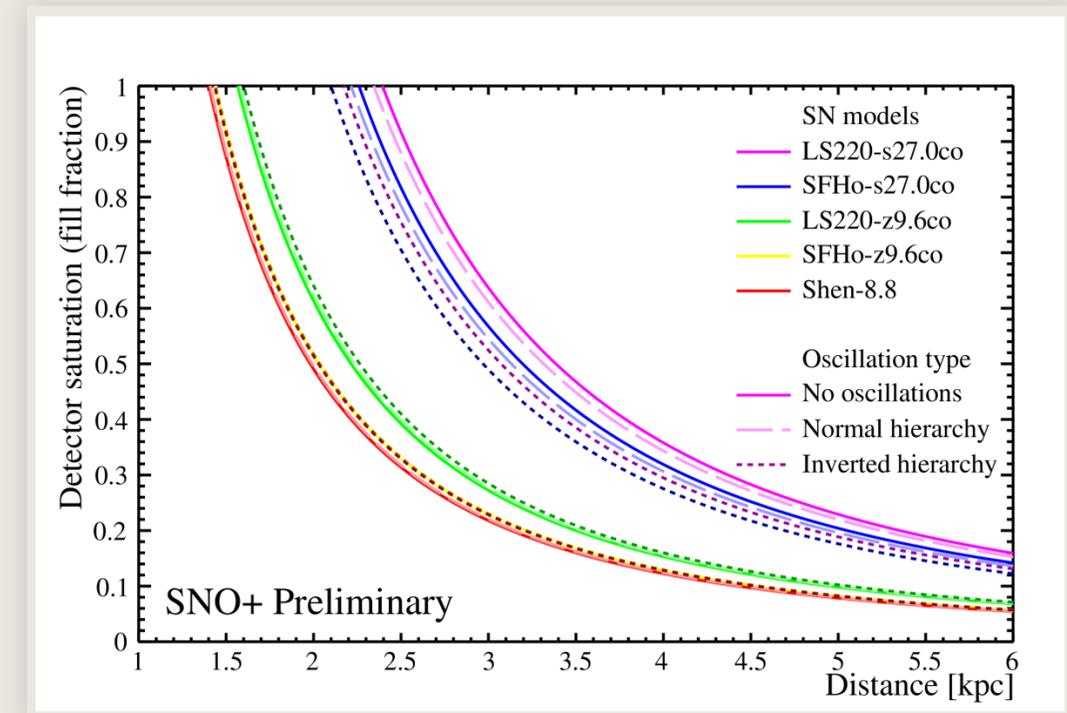


- Data from phase I water analyzed and published
- Phase II water data ready with papers under review
 - NEW: Nucleon Decay, Solar, and Reactor anti-neutrino results coming soon.
- Partial fill backgrounds analyzed to inform future analysis, understand the trigger system, and "practice" the final analysis
- Detector is currently completely filled with liquid scintillator. PPO is still being incrementally added to hit the target concentration.
- Tellurium is expected to be mixed into the detector in 2022, following a stable scintillator phase.

Full scintillator

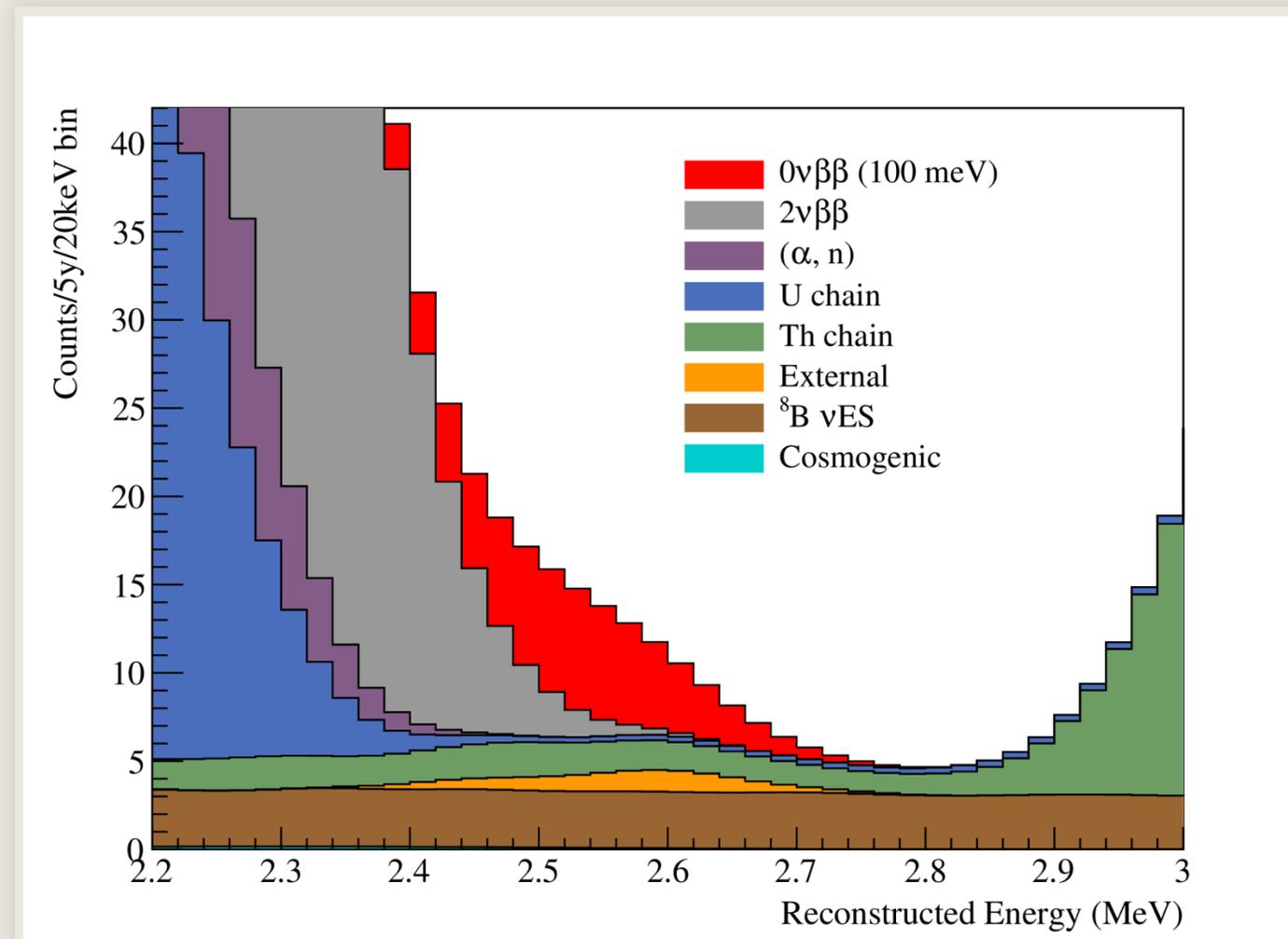
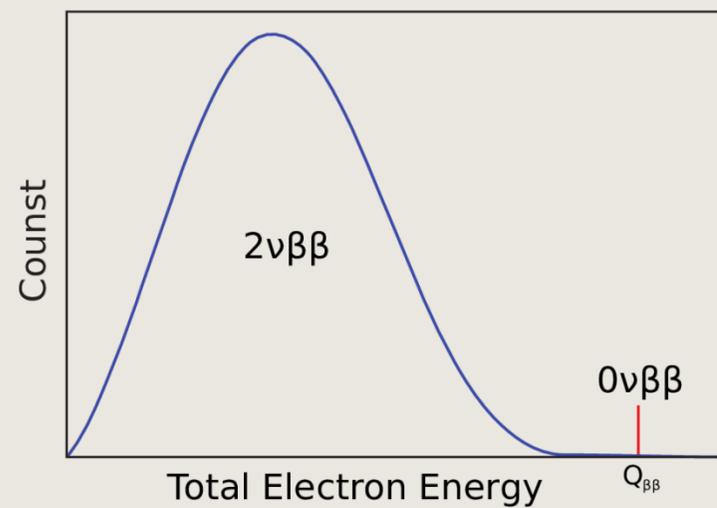
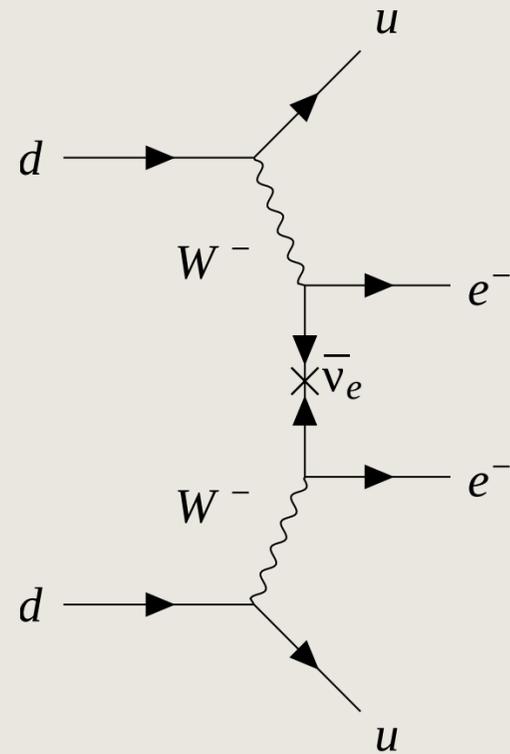
Physics Program:

- Supernova monitoring (join SNEWS)
- Sensitivity to various solar neutrinos (background dependent)
- Reactor and geoneutrino measurements

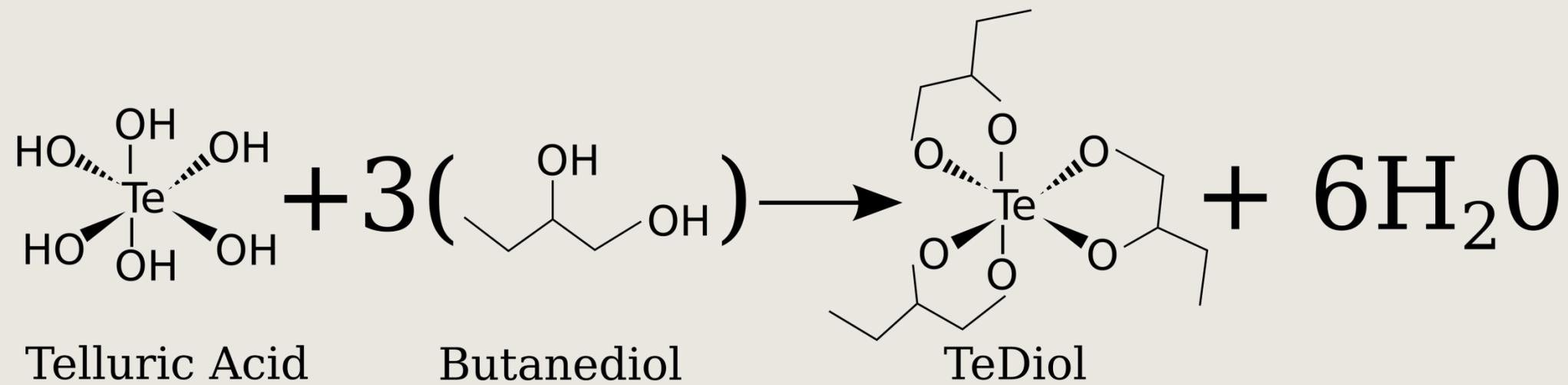


Primary Physics: $0\nu\beta\beta$

With a 2.6 MeV endpoint, the $0\nu\beta\beta$ signal sits in a region of well known backgrounds that can be suppressed through fiducialization and constrained via sideband analyses.



Te-Loading



- Scintillator consists of Linear alkyl benzene (LAB) + 2 g/l PPO.
- Telluric acid dissolved into the scintillator using Butanediol.
- High optical transparency and 50 times more light compared with water.
- Initial loading of 0.5% natural Te by mass.
- Plant is constructed and preparing to fill in 2022.



$0\nu\beta\beta$ Prospects

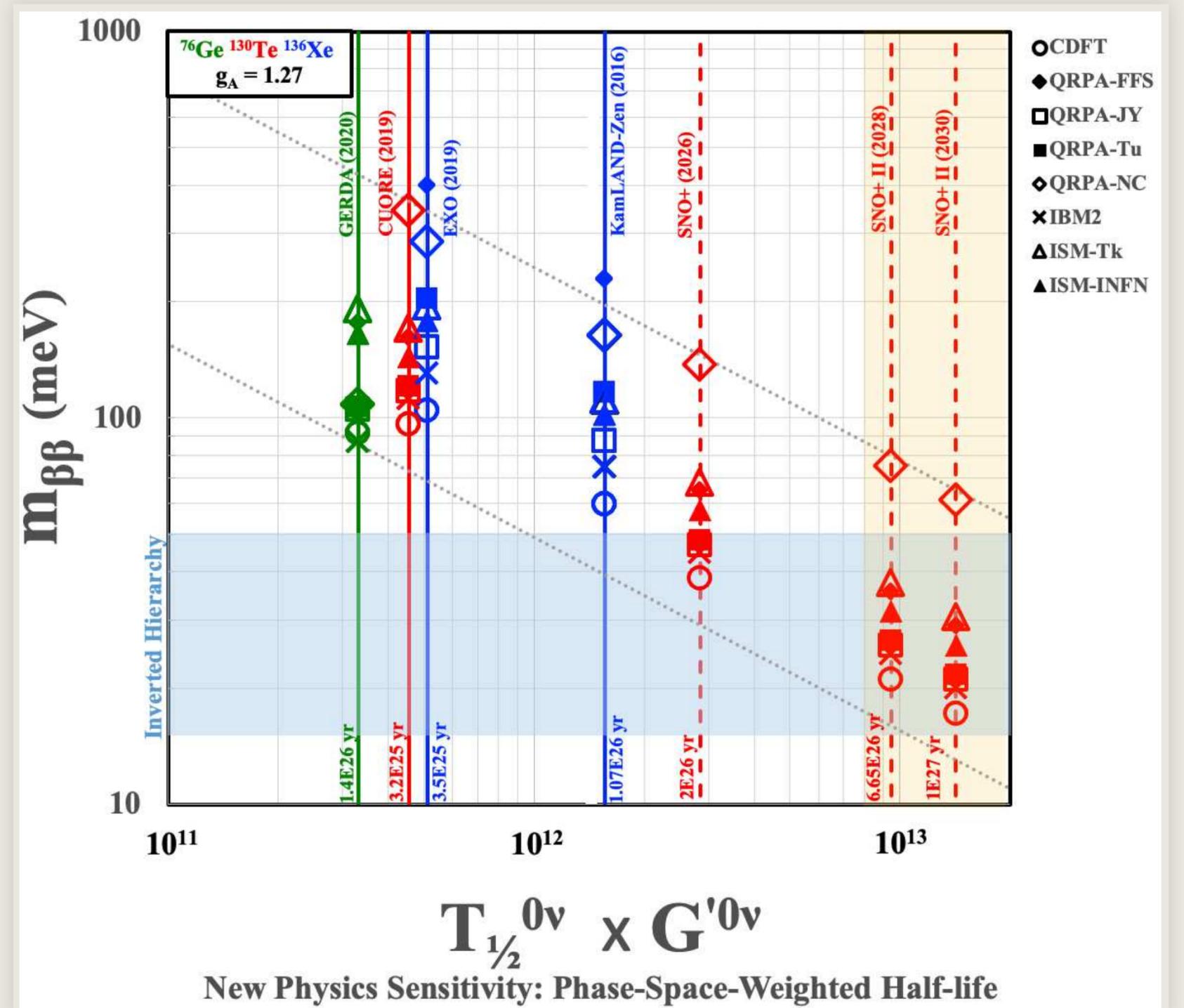
→ Counting analysis for 5 years in an optimized ROI:

- Expected half-life sensitivity $\tau > 2 \times 10^{26}$ years
- $m_{\beta\beta}$ range 37-89 meV (model dependent)

→ More detailed spectral analysis should further constrain backgrounds and improve the signal sensitivity

→ SNO+ technique is scalable

- Increased loading up to 3% with good light yield and stability
- Component upgrades



Conclusions

- The SNO+ detector has completed water phase and completely filled the detector with liquid scintillator.
- PPO additions are ongoing and Tellurium additions are scheduled 2020 (much of the delay and uncertainty related to limited access and scheduling due to COVID).
- Final water analyses completing and beginning to take data for upcoming scintillator results.

